Physisorption and ortho-para conversion of molecular hydrogen on solid surfaces

K. Fukutani, T. Sugimoto

^aInstitute of Industrial Science, The University of Tokyo, Komaba, Meguro-ku, Tokyo 153-8505, Japan ^bDepartment of Chemistry, School of Science, Kyoto University, Oiwake-cho, Sakyo-ku, Kyoto 606-8502, Japan

Abstract

Molecular hydrogen exists in nuclear-spin isomers of ortho and para species according to the total nuclear spin. These species are correlated to the rotational states with even and odd rotational quantum numbers because of the symmetry of the total wavefunction with respect to the permutation of the two nuclei. Although interconversion between the ortho and para states is slow in an isolated state, the conversion is promoted in a physisorption state via interaction with surfaces of not only magnetic but also diamagnetic materials. In a physisorption state, the rotational motion of hydrogen molecules is modified due to the potential anisotropy. The physisorption properties and interconversion rate of the ortho and para hydrogen have recently been investigated on well-defined surfaces, which allow detailed comparison with theory. Furthermore, relative abundance of the ortho and para hydrogen in astronomical circumstances has been reported in recent years, which often shows a value out of equilibrium with the environment temperature. Physisorption and ortho-para conversion on the surfaces of interstellar media are expected to enable deeper understanding of astronomical phenomena. In this article, we review recent progress of experimental and theoretical studies on the physisorption and ortho-para conversion of molecular hydrogen and its relevance to the recent astronomical observation.

Keywords: Molecular hydrogen, ortho-para conversion, physisorption

Preprint submitted to Elsevier

July 19, 2013

Email addresses: fukutani@iis.u-tokyo.ac.jp (K. Fukutani), toshiki@kuchem.kyoto-u.ac.jp (T. Sugimoto)

1 1. Introduction

Hydrogen is the most abundant element in the universe, and ubiquitous 2 in our society. While atomic hydrogen is chemically reactive because of the 3 unpaired electron, molecular hydrogen is rather inert due to its closed shell 4 nature. A remarkable feature of molecular hydrogen is that it is classified 5 into nuclear-spin isomers designated as ortho and para species according to 6 the total nuclear spin [1, 2, 3]. The nuclear spins of proton and deuteron 7 are 1/2 and 1, respectively. Since the indistinguishability of identical nuclei 8 dictates that the total wave function be antisymmetric or symmetric with g respect to the exchange of the two nuclei depending on their spins, the ortho 10 and para species are correlated with the rotational states with odd and even 11 quantum numbers. This was first pointed out theoretically [4, 5, 6] and con-12 firmed experimentally in late 1920's [7, 8]. Presence of nuclear-spin isomers 13 is manifestation of the quantum mechanics, and therefore has gained much 14 attention in both physics and chemistry. 15

In an isolated state, the transition between the ortho and para species is 16 strictly forbidden, the theoretical time scale being 10^{20} s, which is longer 17 than the age of the universe. Interaction with other substances, however, 18 brings about a perturbation, which promotes the conversion between the 19 ortho and para species. The ortho-para (o-p) conversion is, on one hand, 20 of fundamental interest because it includes nuclear-spin flip and rotational-21 state transition, on the other hand technologically important in hydrogen 22 liquefaction. Hydrogen is an efficient energy source as used in fuel cells and 23 rocket fuel, and storing hydrogen in a liquid form is a high-density storage 24 method. When normal hydrogen with the o-p ratio of 3 is liquefied, about 25 40% of the originally stored hydrogen evaporates within 100 h because of 26 rotational-energy release originating from the o-p conversion slowly occurring 27 in liquid hydrogen [9]. This boil-off problem is circumvented with a use of 28 proper catalysts such as iron hydroxides and chromium oxides that enhance 29 o-p conversion during liquefaction and allow for storing hydrogen in the para 30 form. 31

When molecular hydrogen interacts with solid surfaces, the molecule is adsorbed on surfaces either molecularly or dissociatively. While the molecular adsorption occurs through the van der Waals interaction and electric multipole interaction induced by the surface electric field, which is referred to as physisorption, dissociative adsorption is realized via orbital hybridization, which is called chemisorption. The o-p conversion occurs via both

schemes of adsorption: Once H_2 is dissociatively chemisorbed on a solid 38 surface and associatively desorbs from the surface, the relative nuclear-spin 39 direction can be changed implying interchange between the ortho and para 40 states. In a physisorption state, on the other hand, the o-p transition is pro-41 moted through magnetic interaction with surfaces. The former mechanism is 42 sometimes referred to as a chemical process, while the latter is called a physi-43 cal process. Since the adsorption energy for physisorption and chemisorption 44 is typically around 30 meV and 1 eV, respectively, the physical and chemical 45 mechanisms are important at low and high temperatures, respectively. In the 46 present review, we focus on the physical o-p conversion process, in particular 47 the progress made after the previous review article by Ilisca [10]. 48

In the physical conversion, the first step is physisorption in the poten-49 tial well. Important factors include the sticking probability and residence 50 time on the surface. It is noted that the energy scale of the physisorption is 51 similar to that of the rotational motion of hydrogen molecules. Since the ro-52 tational energy of hydrogen molecules is relatively large compared with those 53 of heavier molecules, the rotational motion is retained in the physisorption 54 as recognized as the quantum rotor. Upon physisorption, furthermore, the 55 rotational levels are modified by the anisotropic potential. Since the rota-56 tional states are different between the ortho and para species, the transition 57 between the two species is accompanied by rotational-energy dissipation or 58 excitation. In this regard, the energy level in both the center of mass motion 59 and rotational motion is of importance for the physical o-p conversion. The 60 physisorption properties of hydrogen molecules are described in Sec. 2.2. 61

The second point in the physical conversion is the nuclear-spin flip that changes the nuclear-spin multiplicity of hydrogen molecules. This requires a perturbation that includes magnetic interaction. The magnetic interaction originates from either the spin or orbital motion of electrons and nuclei. The fundamental concept of the nuclear-spin flip is described in Sec. 3.2.

From the experimental point of view, how to probe the nuclear spin of 67 hydrogen and distinguish the ortho and para species was not a trivial prob-68 lem. Since the ortho and para hydrogen has different rotational energies, two 69 experimental approaches are to measure the nuclear-spin state directly and 70 probe the rotational state. The former is nuclear magnetic resonance, and 71 the latter includes the thermal-conductivity measurement originating from 72 the specific-heat difference of the ortho and para species and the rotational 73 spectroscopy using light, neutrons and electrons, which is reviewed in Sec. 4. 74 Since the discovery of the ortho and para species of molecular hydro-75

gen, the adsorption properties and interconversion of ortho and para hy-76 drogen were thoroughly investigated on a variety of surfaces by employing 77 the thermal-conductivity measurement. Whereas the physisorption energy 78 of the ortho and para hydrogen was found to be different [11, 12, 13], the 79 o-p conversion rate was measured on surfaces of magnetic materials including 80 magnetically dilute and dense systems in 1950's – 1960's, which are reviewed 81 in [10]. The o-p conversion process was interpreted on the basis of the Wigner 82 model where the inhomogeneous magnetic field on magnetic surfaces induces 83 the nuclear-spin flip as shown in Sec. 3.2. The o-p conversion rate was fur-84 thermore found to be appreciably affected by external magnetic fields and 85 magnetic phase transition, which also acquired much attention [14, 15, 16]. 86 Although basic understanding of the o-p conversion was developed and many 87 interesting findings were reported in this period, molecular-level understand-88 ing was not satisfactory mainly because most of the experiments were per-89 formed for powder and porous samples, of which surfaces were neither well-90 defined nor well-characterized. It should be pointed out that most metal 91 surfaces are readily oxidized with O_2 and/or H_2O even though the experi-92 ments were intended on clean metal surfaces, and that oxide surfaces might 93 be partially reduced as a result of interaction with hydrogen even if oxide 94 surfaces were initially prepared. 95

In early 1980's, new experimental results suggesting o-p conversion on 96 diamagnetic metals were reported using sophisticated electron energy loss 97 spectroscopy (EELS) experiments [17, 18, 19, 20]. This was actually the 98 beginning of a new era for hydrogen o-p conversion studies, in that the ex-99 periments were performed on well-defined single-crystal surfaces in an ultra-100 high vacuum (UHV) condition. Stimulated by these papers, theoretical ef-101 forts were made [21], and a new conversion model was constructed [22]. The 102 progress in this period as well as historical survey is given in a comprehensive 103 review by Ilisca |10|. 104

Along with this theoretical advance, the effects of molecular orientation 105 on the o-p conversion were theoretically investigated in detail, and a new pro-106 posal to control the conversion was presented [23]. While the theory provides 107 a solid basis for the o-p conversion on metal surfaces, experimental data were 108 not satisfactory because of the difficulty to detect the spin state of hydro-109 gen molecules on well-defined solid surfaces. In 2000's, a new experimental 110 technique probing the ortho and para hydrogen on surfaces was developed 111 by combining the resonance-enhanced multiphoton ionization (REMPI) with 112 desorption techniques [24, 25, 26]. With this technique, the adsorption prop-113

erties of ortho and para hydrogen and the o-p conversion time have been measured on solid surfaces.

In recent years, the o-p ratio of molecular hydrogen in space and plan-116 etary environments has been experimentally observed, which has acquired 117 much attention and is believed to provide us with information on astro-118 chemistry, astrophysics and planetary science [27, 28, 29, 30, 31, 32]. In 119 astrochemistry, possible o-p conversion on surfaces of interstellar media is 120 of considerable interest and importance. In this regard, the o-p conversion 121 has been investigated on ice surfaces, which are dominant interstellar media, 122 in laboratory experiments, and a new conversion model has been proposed 123 [33]. Related to this o-p conversion work, the o-p ratio upon H_2 formation on 124 surfaces has been investigated both experimentally and theoretically, which 125 are described in Sec. 6 along with recent astronomical observation results. 126 In view of the current status of the hydrogen o-p conversion on well-defined 127 solid surfaces, recent progress in this field is reviewed in this article. 128

¹²⁹ 2. Fundamental features of molecular hydrogen

2.1. Electronic, rotational, and nuclear-spin states: classification into nuclearspin isomers

Molecular hydrogen is composed of two protons and two electrons. The 132 total wavefunction of a molecule consists of nuclear wavefunction and elec-133 tronic wavefunction, each wavefunction being a product of orbital and spin 134 functions. Let $\vec{r_1}$ and $\vec{r_2}$ are the position vectors of two electrons in the hy-135 drogen molecule with $\vec{R_a}$ and $\vec{R_b}$ the position vectors of two nuclei as shown 136 in Fig. 1. The spin vectors of the electrons and nuclei are defined as $\vec{s_i}$ 137 (i=1,2) and $\vec{i_j}$ (j=a,b), respectively, and the quantum number of the total 138 electron spin $(\vec{S} = \vec{s}_1 + \vec{s}_2)$ and nuclear spin $(\vec{I} = \vec{i}_a + \vec{i}_b)$ of the molecule are 139 expressed as S and I, respectively. 140

Within the adiabatic approximation, the total wavefunction can be described by a product of the nuclear part $(\chi_n \phi_n)$ and electronic part (ϕ_e) by separating the center of mass motion with $\vec{R} = \vec{R_a} - \vec{R_b}$:

$$\psi(\vec{r_1}, \vec{r_2}, S, \vec{R}, I) = \phi_e(\vec{r_1}, \vec{r_2}, S; \vec{R}) \chi_n(I) \phi_n(\vec{R}), \tag{1}$$

where $\chi_n(I)$ and $\phi_n(\vec{R})$ represent the nuclear-spin part and spatial part of the wavefunction, respectively. Figure 2 shows the adiabatic potentials for



Figure 1: Definition of the position and spin vectors of electrons and nuclei in molecular hydrogen. $\vec{r_i}(i=1,2)$ and $\vec{R_j}(j=a,b)$ are the position vectors of the electrons and nuclei, respectively, and $\vec{s_i}$ (i=1,2) and $\vec{i_j}$ (j=a,b) are the spin vectors of the electrons and nuclei, nuclei, respectively.



Figure 2: Adiabatic potentials for molecular hydrogen. The electron configurations are X ((1s σ_g)²), B and b ((1s σ_g)(2p σ_u)), E,F ((1s σ_g)(2s σ_g)) and C ((1s σ_g)(2p π_u)) [34, 35, 36, 37].

¹⁴⁶ molecular hydrogen, where the ground state of X and excited states of b, B,¹⁴⁷ C and E,F are shown.

The nuclear part is physically understood as the nuclear motion on the adiabatic potentials. The nuclear wavefunction is further decomposed into the radial $(R = |\vec{R}|)$ and angular (θ, ϕ) parts, which represent the vibrational and rotational motions:

$$\phi_n(\vec{R}) = L_v(R) Y_{J,M}(\theta, \phi). \tag{2}$$

In the first approximation, the vibrational motion corresponds to the vibra-152 tion in harmonic potentials with the vibrational frequency described by the 153 quadratic curvature of the potential around the potential-energy minimum. 154 The orbital motion, on the other hand, is characterized by its angular mo-155 mentum J with the wavefunction described by the spherical harmonics, and 156 the energy is described as BJ(J+1) with a rotational constant B in the rigid 157 rotor approximation. The rotational constant is expressed by the moment 158 of inertia that is determined by the atomic mass and interatomic distance. 159 The rotational state has a degeneracy of (2J + 1). If the effects of anhar-160 monicity on the vibration and the change of the interatomic distance due to 161 vibrational excitation on the rotation are taken into consideration, the total 162 energy of the molecule E is described as [38] 163

$$E = T_e + \omega_e (v + \frac{1}{2}) - \omega_e \chi_e (v + \frac{1}{2})^2 + [B - \alpha_e (v + \frac{1}{2})]J(J+1) - \frac{4B^3}{\omega_e^2}J^2(J+1)^2,$$
(3)

where T_e and ω_e are the constants representing the electronic and vibrational energies with a vibrational quantum number of v, and χ_e and α_e express the correction due to the anharmonicity and change of the interatomic distance upon vibrational excitation, which are listed in Tables 1 and 2 for H₂ and D₂.

The H₂ and D₂ molecules consist of two protons with nuclear spin 1/2 and deuterons with nuclear spin 1, respectively, i.e. the proton and deuteron are the Fermi and Bose particles, respectively. Because of the quantum statistics, the total wavefunction of H₂ and D₂ must be antisymmetric and symmetric with respect to permutation of the two nuclei, respectively. In the following, the transformation of the wavefunction due to the two-nuclei permutation is described.

Table 1: Electronic energy (T_e) , vibrational constants (ω_e and $\omega_e \chi_e$) and rotational constants (B and α_e) for H₂ in the unit of cm⁻¹ [39].

[00]					
	T_e	ω_e	$\omega_e \chi_e$	B	α_e
Х	0	4401.21	121.33	60.853	3.062
В	91700	1358.09	20.888	20.0154	1.1845
Е	100082.3	2588.9	130.5	32.68	1.818

Table 2: Electronic energy (T_e) , vibrational constants (ω_e and $\omega_e \chi_e$) and rotational constants (B and α_e) for D₂ in the unit of cm⁻¹ [39].

	T_e	ω_e	$\omega_e \chi_e$	B	α_e
Х	0	3115.5	61.82	30.4436	1.0786
В	91697.2	963.08	11.038	10.068	0.4198
Е	100128.1	1784.42	48.1	16.37	0.6764

The nuclear-spin function is expressed according to the total nuclear spin, 176 I, and I is either 1 or 0 for H_2 and 2, 1, or 0 for D_2 . The spin eigenfunctions 177 are expressed by sums of spin functions as shown in Figs. 3 and 4. Here, each 178 ket corresponds to a product of two nuclear-spin functions with the numbers 179 denoting the z components of the two spin states. As clearly seen in the 180 formulas of Figs. 3 and 4, $\chi_n(I=1)$ and $\chi_n(I=0)$ of H₂ are symmetric and 18 antisymmetric with respect to the permutation of the two spins, respectively. 182 For D₂, on the other hand, whereas $\chi_n(I=2)$ and $\chi_n(I=0)$ are symmetric, 183 $\chi_n(I = 1)$ is antisymmetric with two-spin exchange. Of these states, the 184 state with a larger spin multiplicity is called ortho and the other is called para 185 species. The ortho and para H_2 are 3-fold and 1-fold degenerate, respectively, 186 while the degeneracy of ortho and para D_2 is 6 and 3, respectively. 18

The exchange of the position vectors of the two nuclei correspond to the transformation of the polar and azimuthal angles as $\theta \to \pi - \theta$ and $\phi \to \pi + \phi$ as illustrated in Fig. 5. The rotational wavefunction is then transformed as Eq. 4. Since the two-nuclei exchange does not affect the inter-nuclear distance R, the vibrational wavefunction of diatomic molecules does not change the sign irrespective of the vibrational quantum number as

$$L_{v}(R) \rightarrow L_{v}(R)$$

$$Y_{J,M}(\theta,\phi) \rightarrow Y_{J,M}(\pi-\theta,\pi+\phi) = (-1)^{J} Y_{J,M}(\theta,\phi).$$
(4)

$$\begin{split} \chi_n(I=1,I_z=1\sim-1) & \chi_n(I=0,I_z=0) \\ & \left\{ \begin{aligned} & \left|\frac{1}{2},\frac{1}{2}\right\rangle & \frac{1}{\sqrt{2}} \left(\left|\frac{1}{2},-\frac{1}{2}\right\rangle - \left|-\frac{1}{2},\frac{1}{2}\right\rangle\right) \\ & \frac{1}{\sqrt{2}} \left(\left|\frac{1}{2},-\frac{1}{2}\right\rangle + \left|-\frac{1}{2},\frac{1}{2}\right\rangle\right) \\ & \left|-\frac{1}{2},-\frac{1}{2}\right\rangle \end{aligned} \right.$$

Figure 3: Spin eigenfunctions χ_n of H₂ for I=1 and 0 with the z component (I_z) of the total nuclear spin \vec{I} . The first and second numbers in the kets represent the z components of the two spins. I=1 and 0 correspond to o-H₂ (odd J) and p-H₂ (even J), respectively.

Table 3: Symmetry with respect to two-nuclei permutation. S and A denote symmetric and antisymmetric states, respectively.

		\mathbf{S}	А
Electron		Σ_g^+, Σ_u^-	Σ_u^+, Σ_g^-
Rotation (J)		even	odd
Nuclear spin (I)	H_2	1	0
	D_2	2, 0	1

As easily understood from Eq. 4, the rotational wavefunction changes its sign when J= odd.

Finally, the effect of the two-nuclei exchange on the electronic wavefunction is considered. With the two-nuclei permutation, the electron coordinates are changed as inversion with respect to the molecule center and mirror reflection with respect to a mirror plane through the molecular axis. Hence, the electron wavefunctions denoted as Σ_g^+ and Σ_u^- are symmetric, and those as Σ_g^- and Σ_u^+ are antisymmetric [p. 131 of Ref. [38]]. The symmetry relation is summarized in Table 3.

Since the total wavefunction of H_2 and D_2 is antisymmetric and symmetric with respect to the two-nuclei permutation as described above, the rotational state is correlated with the nuclear-spin state as depicted in Fig. 6 for the electronic states of Σ_g^+ and Σ_u^- . Shown on the left-hand side are para H_2 (p-H₂) and ortho D_2 (o-D₂) with even J, and ortho H_2 (o-H₂) and para D_2

$$\begin{split} \chi_n(I=2,I_z=2\sim-2) & \chi_n(I=1,I_z=1\sim-1) \\ \begin{cases} |1,1\rangle \\ \frac{1}{\sqrt{2}}(|1,0\rangle+|0,1\rangle) \\ \frac{1}{\sqrt{6}}(|1,-1\rangle+2|0,0\rangle+|-1,1\rangle) \\ \frac{1}{\sqrt{2}}(|0,-1\rangle+|-1,0\rangle) \\ \frac{1}{\sqrt{2}}(|0,-1\rangle+|-1,0\rangle) \\ |-1,-1\rangle \\ \chi_n(I=0) \\ \frac{1}{\sqrt{3}}(|1,-1\rangle-|0,0\rangle+|-1,1\rangle) \end{split}$$

Figure 4: Spin eigenfunctions χ_n of D₂ for I=2, 1 and 0 with the z component (I_z) of the total nuclear spin \vec{I} . The first and second numbers in the kets represent the z component of each spin. I=2 and 0 correspond to o-D₂ (even J), and I=1 corresponds to p-D₂ (odd J).



Figure 5: Transformation of the coordinates due to exchange of the two-nuclei position vectors that corresponds to reverse the sign of the relative position vector as $\vec{R} \to -\vec{R}$. The polar and azimuthal angles are transformed as $\theta \to \pi - \theta$ and $\phi \to \pi + \phi$.

 $_{208}$ (p-D₂) with odd J are shown on the right-hand side.

The partition function describing the population of each rotational level (E_J) at a temperature of T for o-H₂ ($Z_o(T)$) and p-H₂ ($Z_p(T)$) is described as

$$Z_{o}(T) = \sum_{J=\text{odd}} g_{(I=1)}(2J+1) \exp(-\frac{E_{J}}{k_{B}T})$$
$$Z_{p}(T) = \sum_{J=\text{even}} g_{(I=0)}(2J+1) \exp(-\frac{E_{J}}{k_{B}T}),$$
(5)

where $g_{(I=1)}=3$ and $g_{(I=0)}=1$ are the nuclear-spin degeneracies and (2J+1)212 represents the rotational-state degeneracy. The population ratio of $o-H_2$ and 213 $p-H_2$ (OPR: ortho-para ratio) in thermal equilibrium is then obtained by 214 $Z_{\rho}(T)/Z_{p}(T)$, which is shown in Fig. 7 as a function of temperature. At a 215 sufficiently low temperature, this value leads to 0, whereas it reaches 3 at a 216 high temperature. Similarly, the para-ortho ratio of D_2 becomes 0 and 1/2 at 217 low and high temperature limits, respectively, as obtained from the partition 218 function of D_2 . 219

220 2.2. Adsorption on solid surfaces

221 2.2.1. Adsorption potential

Molecular hydrogen is adsorbed on surfaces mainly in two schemes: dis-222 sociative chemisorption and molecular physisorption. In the former case, H_2 223 is dissociated and two hydrogen atoms are strongly bound via chemical bond 224 formation. In the latter case, on the other hand, H_2 is weakly trapped on 225 surfaces via the van der Waals interaction and electric multipole interaction. 226 The typical adsorption energy is 0.5 eV per H atom for chemisorption and 22 30 meV per H_2 molecule for physisorption [40, 41, 42]. Since the ortho-para 228 conversion via dissociative adsorption is out of the scope of this article, only 229 the physisorption properties of H_2 are described in this section. 230

Physisorption of H_2 is realized by competition between the attractive van 231 der Waals force and Pauli repulsion due to the overlap of the closed-shell 232 electron wavefunctions with substrate electrons. Physisorption of various 233 molecules was reviewed in a recent article [43]. As schematically shown in 234 Fig. 8, the substrate electron density is smeared out at the surface, which 235 determines the physisorption position. As discussed later, furthermore, the 236 electron density at the substrate plays a decisive role in the ortho-para con-237 version. As intuitively understood from Fig. 8, both the van der Waals force 238 and Pauli repulsion are dependent on the direction of the molecular axis. 239



Figure 6: Rotational-energy levels for low-J states of the ortho and para species of H₂ and D₂ in the electronic Σ_g^+ and Σ_u^- states. E = BJ(J+1) denotes the rotational energy with the rotational constant B and rotational quantum number J.



Figure 7: Ortho to para ratio (OPR) of H_2 calculated from the partition function for the rotational level as a function of temperature.



Figure 8: Schematic illustration of H₂ physisorbed on a surface at a certain distance. The surface electron density decays exponentially from the surface with increasing distance (z). H₂ with the $1s\sigma_g$ orbital exists in the tail region of the surface electron with the molecular-axis angle of Θ from the surface normal direction.



Figure 9: (a) Schematic figure of the physisorption potential due to the isotropic term V_0 as a function of the molecule-surface distance (z). The vibrational levels are denoted with the vibrational quantum number n. (b) Splitting of the rotational-energy level due to the $V_2(z)P_2(\cos\Theta)$ anisotropic term in the case of $V_2 < 0$.



Figure 10: Electron energy loss spectra taken for H₂, HD and D₂ physisorbed on Cu(100) at about 10 K. The two loss peaks in each spectrum (except the feature at 9.9 meV in HD) correspond to $n=0 \rightarrow 1$ and $n=0 \rightarrow 2$ excitations of the molecule–substrate vibration shown in Fig. 9(a). Reproduced by permission from [44]. The vibrational energy is smaller for heavier isotopes.

The physisorption potential $V(z, \Theta)$ of a molecule is mainly characterized by its well depth and anisotropy, which represent the electrostatic energy as functions of the distance (z) of the center of mass of the molecule from the surface and the polar angle (Θ) of the molecular axis with respect to the surface normal. The potential is then generally expanded in Legendre functions as

$$V(z,\Theta) = \sum_{\ell=0}^{\infty} V_{\ell}(z) P_{\ell}(\cos\Theta).$$
(6)

Since H_2 is a homonuclear molecule, ℓ is restricted to even integers.

The first term of this formula represents the isotropic potential, which is 247 schematically shown in Fig. 9(a). The well depth was theoretically evaluated 248 to be 10 - 40 meV depending on the substrate [45, 46], which were also 249 experimentally examined by selective adsorption experiments [47, 48, 49, 250 50, 51, 52, 53, 54, where molecular hydrogen is selectively adsorbed on 251 surfaces at particular conditions of the kinetic energy and incidence angle 252 corresponding to the bound states in the potential well. The adsorption 253 energy can also be experimentally evaluated by temperature-programmed 254 desorption (TPD) [55, 56, 57], which revealed the physisorption energy of 255 30 - 60 meV on ice and carbon surfaces. The vibrational states in the 256 physisorption well were directly probed with EELS by Andersson's group 25 [44]. Figure 10 shows the EELS data taken for H_2 , D_2 and HD. Several loss 258 features were observed at 7 – 15 meV, which were attributed to $n=0 \rightarrow 1$ and 259 $n=0 \rightarrow 2$ excitations in Fig. 9(a), which compared well with the results of 260 selective adsorption resonances. At low-symmetry sites like steps and kinks, 26 the physisorption potential is largely modified as compared to the flat terrace 262 site, and more strongly adsorbed H_2 has been identified [58, 59, 60, 61, 62, 63]. 263 On adatom (Au and Cu)-adsorbed Cu(100), on the other hand, D_2 was found 264 to be strongly adsorbed at the adatom site with a binding energy of about 265 40 meV [64]. Owing to recent advances of the density functional theory for 266 the van der Waals interaction, the physisorption energy and potential of H_2 267 are theoretically treated from first principles [65, 66]. 268

As mentioned in Sec. 2.1, the rotational energies of H₂ in the J=1 and 2 states are 14.7 and 43.9 meV, respectively, which are comparable to the physisorption energy. In contrast to heavier molecules, the rotational motion of molecular hydrogen is retained in the adsorption state. This is why molecular hydrogen is called a quantum rotor. Since the $V_{\ell}(z)P_{\ell}(\cos\Theta)$ ($\ell \geq 2$) terms are expected to be comparable or smaller than the physisorption well depth due to the isotropic term $(V_0(z))$, the effect of the anisotropic potential can be treated as a perturbation.

According to the perturbation theory, the first-order energy shift due to the rotational motion is described by

$$\Delta E^{(1)}(n, J, M) = \langle n | V_2 | n \rangle \langle JM | P_2(\cos \Theta) | JM \rangle$$

= $\frac{3 \langle n | V_2 | n \rangle}{2J + 3} (\frac{J^2 - M^2}{2J - 1} - \frac{J}{3}),$ (7)

where $\langle n|V_2|n\rangle$ is the matrix element of $V_2(z)$ with respect to the molecule-279 surface vibrational wavefunction in $V_0(z)$ and $|JM\rangle$ represents the rotational 280 wavefunction of $Y_{J,M}(\Theta, \phi)$ [67]. The integral of $\langle JM | P_2(\cos \Theta) | JM \rangle$ is easily 281 done following the formula for the angular momentum [68]. It is clear that 282 the rotational-state degeneracy is lifted as shown in Fig. 9(b). Since the 283 rotational wavefunction is unperturbed in the first-order perturbation, this 284 is called nearly free rotor here¹. The second-order energy term, on the other 285 hand, is expressed as 286

$$\Delta E^{(2)}(n,J,M) = \sum_{n',J',M'} \frac{|\langle n'|V_2|n\rangle\langle J'M'|P_2(\cos\Theta)|JM\rangle|^2}{E(nJM) - E(n'J'M')}.$$
 (8)

This second-order energy becomes significant when $\langle n'|V_2|n\rangle$ is in the same 287 order as E(nJM) - E(n'J'M'), i.e. 44 meV for J = 0 and 73 meV for J = 1288 in the case of n = n'. In this case, the rotational wavefunction is distorted 289 from that of the original J state because of mixing with other rotational 290 states, which is called hindered rotation here. The energy levels and wave-291 functions of such hindered rotation were theoretically analyzed in detail [69]. 292 Figure 11 shows the obtained energy levels as a function of the potential 293 anisotropy for J=0 and 1 states [69]. Historically, the anisotropic potential 294 was treated as either $C\cos^2\Theta$ or $C'\sin^2\Theta$ instead of $P_2(\cos\Theta)$ [69], where 295

¹The rotational state under an isotropic potential is sometimes called three-dimensional (3D) rotation. In molecular physics, on the other hand, the 3D rotation occasionally denotes the rotational motion of non-linear molecules, which have three rotational axes (Θ, ϕ, χ) as compared to linear molecules (Θ, ϕ) called 2D rotation. To avoid confusion, we use the expressions of free rotor, nearly-free rotor, and hindered rotor to represent the rotational motion under isotropic potential, small anisotropic potential, and large anisotropic potential, respectively, in this article.



Figure 11: Rotational energy for the J=1 and J=0 states as a function of the square root of the ratio of the anisotropic potential to the rotational constant. The anisotropic potential is either $C \cos^2 \Theta$ or $C' \sin^2 \Theta$. Without anisotropic potential, the molecule behaves as a free rotor (FR). While the molecule is in a nearly free rotor (NFR) state under a small anisotropic potential, the rotational motion is described as a hindered rotor (HR) under a large anisotropy. Adapted from [69].

C and C' are constants. Since the mean value of the anisotropic potential 296 with these formulas is positive, the energy level gets higher along with the 297 lifting of the degeneracy of the J=1 state as C or C' becomes larger. As the 298 anisotropic potential becomes infinitely large, the rotational motion tends 299 to be confined in either the surface-normal or parallel direction depending 300 on the sign of $V_2(z)$, which is sometimes designated one-dimensional (1D) 301 or two-dimensional (2D) rotation in literatures, respectively. In the former 302 case, the energy level of J=0 becomes equal to that of J=1 (M=0). On flat 303 surfaces, however, such a large $V_2(z)$ value is unlikely to exist. 304

The rotational sublevel splitting was investigated both experimentally and theoretically. Selective adsorption experiments directly identified the level splitting and evaluated the potential anisotropy to be 1 - 2 meV on the basis of Eq. 7 [70, 71, 72]. Although the resolution was not sufficiently good for the observation of the splitting, an EELS study, by applying a particu-

lar scattering geometry, analyzed the sublevel splitting of 1.4 meV for the 310 J=2 state on Cu(100) [73]. The anisotropy of the van der Waals interaction 311 originates from the asymmetry of the molecular polarizability, which was 312 theoretically evaluated to be as small as 7 % of the isotropic term because of 313 the image charge effect [67]. On the other hand, the anisotropy of the repul-314 sive part might be larger and dominant for physisorption. The physisorption 315 energy and potential anisotropy of H_2 are theoretically analyzed on surfaces 316 by Multi-reference-configuration-interaction calculations [74, 75] and van der 317 Waals density functional theory [65]. On a more corrugated surface such as 318 inner walls of nanopores, furthermore, the azimuthal dependence of the po-319 tential is significant, which lifts the degeneracy of the $M = \pm 1$ states in Fig. 320 9(b) It is noted that pure-rotational spectroscopy detailed in Sec. 4 allows 32 for experimental analysis of the rotational-sublevel splitting and potential 322 anisotropy [76, 77, 78]. 323

324 2.2.2. Physisorption energy of ortho and para species

One important consequence of the potential anisotropy and resulting ro-325 tational sublevel splitting is the difference of the adsorption energy between 326 the ortho and para species. The physisorption energy is determined by the 327 sum of the isotropic term and anisotropic term of the potential. The isotropic 328 term is the same for both ortho and para species. As clearly demonstrated 329 in Fig. 9(b), on the other hand, some of the ortho- H_2 level are lowered due 330 to the anisotropic term, the other being raised, while the $p-H_2$ level remains 331 unchanged. This indicates that $o-H_2$ can be more strongly bound to surfaces 332 under an anisotropic potential. 333

The adsorption energy difference was first recognized as the vapor pres-334 sure difference between the ortho and para species [79, 80]. This effect has 335 been used for separation of $o-H_2$ and $p-H_2$: Under a fixed $n-H_2$ pressure, 336 the o-p ratio in the physisorption state is higher than the gas phase value 33 of 3 because of the larger adsorption energy for $o-H_2$, which is called the 338 separation coefficient [13, 81]. The separation coefficient was theoretically 339 analyzed in terms of the rotational-energy level splitting in the physisorp-340 tion well [11, 12, 82, 69, 83, 84]. On these bases, highly purified $o-H_2$ and 34 $p-D_2$ can be obtained with several condensation steps. The adsorption en-342 ergy difference was also applied to Chromatographic separation, where more 343 strongly bound o- H_2 is released later than p- H_2 [85, 86, 87]. 344

The o-p separation was often performed on ionic surfaces like Al_2O_3 . The origin of the potential anisotropy was, therefore, considered to be the

electrostatic energy originating from the hydrogen polarization induced by 347 the electric field on an ionic-material surface [88]. From the chromatographic 348 data, the potential barrier for rotation was experimentally evaluated to be 349 C=0.46 kcal/mol for the anisotropy of $C\cos^2\Theta$ [89]. In addition to the 350 out-of-plane anisotropy, the in-plane anisotropy was also analyzed [90], and 35 the theory was applied to the heat capacity, entropy and o-p separation 352 coefficient [91, 92]. The difference of the adsorption energy was also examined 353 on a Graphon surface by a low-temperature calorimetric method: the heat 354 of adsorption was estimated to be 0.91 and 0.96 kcal/mol for $p-H_2$ and $o-H_2$, 355 respectively [93]. 356

The adsorption energy difference can also be evaluated by probing the 357 thermal stability of the ortho and para species on a surface. As detailed 358 in Sec. 4, infrared absorption spectroscopy (IRAS) and resonance-enhanced 359 multiphoton ionization (REMPI) combined with desorption techniques have 360 been applied to distinguish the rotational state of molecular hydrogen on 361 surfaces. By using IRAS, o-H₂ was shown to be more strongly adsorbed than 362 $p-H_2$ on NaCl [94]. TPD spectroscopy combined with REMPI was performed 363 on activated Al_2O_3 [24], amorphous ice [26, 95] and Ag surfaces [96]. Figure 364 12 shows the TPD spectra taken for H_2 from activated Al_2O_3 . After exposure 365 of the sample to n-H₂ of 3×10^3 L at 14 K, TPD of H₂ in the J=0 and 1 366 states was recorded. The peak temperature of the desorption signal in J=136 is clearly higher than that of J=0 suggesting that H_2 in J=1 is more strongly 368 bound than J=0. Similar experiments were performed for D_2 on amorphous 369 ice surfaces [26], which is shown in Fig. 13. Dominant desorption species are 370 in the J=0 and 1 states, and the desorption temperature of J=1 is obviously 371 higher than that of J=0 indicating stronger binding of D_2 in J=1. 372

373 2.2.3. Sticking probability

When $o-H_2$ undergoes o-p conversion in a physisorption well, the initial 374 step is the sticking of $o-H_2$ to surfaces from gas phase. The overall conversion 375 probability therefore depends on the sticking probability. The sticking prob-376 ability of light molecules like hydrogen on a substrate with heavier elements 377 is usually small, typically ~ 0.1 because the energy dissipation is inefficient. 378 When the surface is covered by H_2 , the sticking probability is reported to 379 be 1 [97], because the kinetic energy of the incident molecule is readily ac-380 commodated by preadsorbed molecules. In an o-p conversion experiment, 38 furthermore, the relative coverage of $o-H_2$ and $p-H_2$ is often discussed. Then, 382 the difference of the sticking probability onto surfaces of these species has 383



Figure 12: Rotational-state-selective temperature-programmed desorption spectra of H_2 from activated Al_2O_3 powder measured with a heating rate of 0.1 K/s. H_2 in J=0 and 1was probed by REMPI after exposure of the sample to n-H₂ of 3×10^3 L at 14 K. The activated Al_2O_3 powder with a specific area of 270 m²/g was heated at 520 K for 10 h prior to the experiments. The dip at 42 K for the J=0 spectrum is due to the REMPI laser instability, and the dashed curve is a guide for eyes. Reproduced by permission from [24].



Figure 13: Rotational-state-selective temperature-programmed desorption spectra of D_2 from amorphous ice surfaces. D_2 of 1.5 monolayer (ML) was adsorbed on porous amorphous ice of 10 ML at 10 K. D_2 is mostly desorbed in J=0 and 1 states, and the desorption temperature for J=1 is higher than that of J=0. Reproduced by permission from [26].

to be carefully examined. Historically, when EELS studies first observed a higher intensity of the loss feature corresponding to $p-H_2$ as compared to that of $o-H_2$, the sticking probability of $p-H_2$ was suspected to be higher than $o-H_2$. Such possibility was later discussed to be small. In this subsection, previous studies on the sticking probability of ortho and para hydrogen are briefly described.

For molecular adsorption on solid surfaces, an important factor governing 390 the sticking probability is the dissipation of the incident kinetic energy and 391 adsorption energy. In the physisorption potential of Fig. 9(a), the impinging 392 molecule is in a continuum state with a positive energy, and comes from larger 393 z toward the potential minimum. Sticking of a molecule is regarded as the 394 transition from a continuum state with a positive energy to a bound state 395 with a negative energy in the potential by dissipating the energy between 396 the initial and final states to other degrees of freedom. Two important paths 39 of the energy dissipation are phonon excitation and electronic excitation of 398 surfaces. In the case of chemisorption, electronic excitation is important (e.g. 399 [98]). As for the weak physisorption of light molecules, on the other hand, 400 coupling to electronic excitation is expected to be weak, and sticking occurs 401 via energy transfer to the substrate phonon system [53]. Other important 402

factors for molecular sticking are temporal trapping in the physisorption well 403 via the energy transfer of the kinetic energy of an impinging molecule in 404 the surface normal direction to the kinetic energy in the parallel direction 405 or the internal energy. This process corresponds to the selective adsorption 406 resonance. As the internal-vibrational energy of molecular hydrogen is as 407 high as 500 meV, the accessible internal mode is the rotational motion of the 408 molecule. Since the total energy of the trapped molecule is still positive, this 409 corresponds to a quasibound state, and further energy dissipation is required 410 for the transition of the molecule into the true stuck state with a total energy 411 of a negative value. 412

Experimentally, Andersson and Harris investigated the sticking probabil-413 ity of n-H₂ and p-H₂ with EELS and workfunction change ($\Delta \phi$) measure-414 ments to clarify the effect of the rotational motion on the sticking [99]. They 415 first evaluated the EELS cross section for $J=0 \rightarrow 2$ and $J=1 \rightarrow 3$ transi-416 tions to be the same as those in the gas phase, then precisely determined the 417 ortho and para H_2 coverages from the EELS intensities. On the basis of the 418 data, the sticking probability ratio for $p-H_2$ to $n-H_2$ was evaluated to be 1.5. 419 They proposed that the trapping of H_2 via translational to rotational energy 420 transfer is significant. 421

In later publications, the sticking probability of H_2 and D_2 was investi-422 gated in a more sophisticated way [100, 101, 53]. By changing the incident 423 energy and angle of molecular beams, both the sticking and reflection of in-424 cident molecules were experimentally measured in detail. Figure 14 shows 425 the initial sticking coefficient, S_0 , of n-H₂ and p-H₂ on Cu(100) at 15 K as a 426 function of the molecule energy. The sticking probability revealed a smooth 427 behavior falling off around 30 - 50 meV. This energy reflects the substrate 428 phonon density of states, which points to the phonon-mediated dissipation 429 [100, 53, 102]. This broad feature extending up to ~ 50 meV cannot be 430 explained by classical theory, and quantum treatment was shown to be nec-431 essary to reproduce the sticking behavior [53]. Overlapped with this smooth 432 behavior, an enhancement of S_0 at an energy corresponding to rotational 433 excitation of $J=0 \rightarrow 2$ was observed as indicated in Fig. 14. After more de-434 tailed and careful experiments and theoretical analysis, all features observed 435 in the sticking curves were attributed to either corrugation-mediated selec-436 tive adsorption resonances, rotation-mediated selective adsorption resonances 437 or combined resonances involving both rotational excitation and diffraction 438 [52, 53].439

For the energy transfer from the translational motion to rotational mo-



Figure 14: Sticking probability of n-H₂ and p-H₂ as a function of the incident energy of H₂ on Cu(100) at 15 K. Both reveal a broad feature extending up to \sim 50 meV, and the curve for p-H₂ shows a hump corresponding to the rotational excitation of $J=0 \rightarrow 2$. Reproduced by permission from [100].

tion, there must be coupling between these modes; while the two motions are 441 independent under an isotropic potential, they are coupled by the anisotropic 442 term of the potential described in Sec. 2.2. The energy transfer from the 443 normal motion to parallel motion, on the other hand, can be achieved by the 444 potential corrugation on the surface, i.e. potential difference as a function 445 of the molecule position in the surface parallel direction. When further in-446 teraction with the substrate phonon is taken into consideration, relaxation 447 from the quasibound state to the bound state can be calculated. According 448 to semiclassical trajectory calculations [103] and one-phonon distorted-wave 449 Born-approximation calculations [104], the selective adsorption resonances 450 are weak and dependence of the sticking probability on the rotational state 451 is small. Sticking and associated energy dissipation are obviously important 452 and continuing issues in Surface Science [105, 106]. 453

454 3. Ortho-para conversion – theoretical aspect

As described in previous sections, the wavefunctions of the ortho and para hydrogen are orthogonal with each other without coupling between the

nuclear-spin and rotational motion. Other molecules such as NH_3 , CH_4 and 457 H_2O are also classified into nuclear-spin isomers. The theory of the ortho-458 para conversion of polyatomic molecules was well developed [107, 108], and 459 has been applied to e.g. CH_3F [109], C_2H_4 [110], H_2CO [111], CH_4 [112], C_2H_3 460 [113] and H_2O [114]. The theory is essentially based on the intramolecular 461 mixing of the ortho and para states. Because of the spin-spin and spin-462 rotation interaction, there possibly occurs mixing between the ortho and para 463 states. Although these coupling terms are generally small and negligible, the 464 mixing becomes significant when the energy levels of the ortho and para states 465 are near-degenerate. This may be accidentally realized because polyatomic 466 molecules have some rotational modes with different rotational constants. 46 The mechanism of the o-p conversion of polyatomic molecules was discussed 468 that thermal excitation to the near-degenerate levels results in finite mixing 469 between the ortho and para states leading to o-p conversion [108]. 470

⁴⁷¹ Compared to the polyatomic molecules, because of the small moment ⁴⁷² of inertia, hydrogen molecules have a large energy separation between the ⁴⁷³ rotational states. In the physisorption state at low temperature dealt with ⁴⁷⁴ in this article, therefore, the rotational states accessible by thermal excitation ⁴⁷⁵ are restricted to low-J sates, only J=1 and 0. Under this condition, mixing ⁴⁷⁶ of the ortho and para states is extremely small and spontaneous conversion ⁴⁷⁷ probability can be neglected as described below.

When there is a perturbation H' due to interaction with surfaces and the matrix element between the ortho and para states has a non-zero value, the o-p transition probability becomes significant. In this section, fundamental concepts for the o-p conversion are reviewed.

482 3.1. Gas phase conversion mechanism

483 3.1.1. Radiative transition

In an isolated state, radiative transition occurs through interaction of a molecule with the electromagnetic field. Following the ordinary procedure, the dominant perturbation term is the electric dipole described with the position vectors of electrons $(\vec{r_1}, \vec{r_2})$ and nuclei $(\vec{R_a}, \vec{R_b})$ and the electric field \vec{E} :

$$H' = e(\vec{r_1} + \vec{r_2} - \vec{R_a} - \vec{R_b}) \cdot \vec{E}.$$
(9)

It is easy to see that the dipole moment apparently remains unchanged by exchange of the position vectors of \vec{R}_a and \vec{R}_b . Since the symmetry of the ⁴⁹¹ wavefunction with respect to the two-nuclei position exchange is different for ⁴⁹² the ortho and para species, either symmetric or antisymmetric, the dipole ⁴⁹³ moment between the ortho and para states is identically zero. This indicates ⁴⁹⁴ the transition between the ortho and para states is a forbidden process. Con-⁴⁹⁵ sidering the spin part, because the dipole moment does not contain the spin ⁴⁹⁶ vector, the ortho and para states with the nuclear-spin triplet and singlet ⁴⁹⁷ states are also orthogonal.

The radiative ortho-para conversion probability was first discussed the-498 oretically by considering the hyperfine interaction strength and mixing of 499 the electronically excited states [115]. The transition probability was evalu-500 ated to be $\sim 10^{-10}$ s⁻¹ [115, 1, 116]. In their discussion, however, only the 50 mixing with the ${}^{3}\Sigma_{u}^{+}$ state was taken into consideration, which means that 502 the selection rule for the rotational and spin states was not considered cor-503 rectly. Therefore, the value so obtained was an overestimation. Later, the 504 o-p conversion probability was more accurately evaluated on the basis of the 505 relativistic spin-orbit interaction (nuclear spin and electron orbital momenta) 506 [117, 118, 119, 120], and the formula for the probability of the $J \rightarrow J - 1$ 507 transition was obtained. According to this formula, the $J = 1 \rightarrow J = 0$ 508 transition probability can be calculated to be 7×10^{-20} s⁻¹. A more accurate 509 and complete calculation was recently performed by Pachucki and Komasa 510 by taking account of a nonadiabatic correction for H_2 and relativistic correc-511 tions, the value obtained being $6.2 \times 10^{-14} \text{ yr}^{-1} = 2.0 \times 10^{-21} \text{ s}^{-1}$ [121]. Note 512 that this time scale for conversion (inverse of the conversion probability) is 513 longer than the age of the universe $(1.37 \times 10^{10} \text{ yr})$. 514

515 3.1.2. Proton exchange

The o-p conversion of H_2 in an isolated state does not take place practically. When H_2 encounters other atoms or ions, the o-p conversion might be induced via proton exchange [122]:

$$\begin{array}{rcl} \mathrm{o}-\mathrm{H}_2+\mathrm{H} & \rightarrow & \mathrm{p}-\mathrm{H}_2+\mathrm{H} \\ \mathrm{o}-\mathrm{H}_2+\mathrm{H}^+ & \rightarrow & \mathrm{p}-\mathrm{H}_2+\mathrm{H}^+. \end{array}$$

These processes have been investigated in gas phase. Of the two reactions, there is an activation barrier of ~0.3 eV for the H₂ + H reaction [123]. The conversion cross section was estimated to be less than 10^{-16} cm³ s⁻¹ at 300 K [123, 124]. It should be noted that the H atom possibly causes ortho-para conversion of H₂ through magnetic interaction on the basis of

the Wigner model described in Sec. 3.3.2 because H possesses an unpaired 524 electron spin. As to the latter reaction, on the other hand, the reaction rate 525 for $H_2 + D^+ \rightarrow HD + H^+$ was experimentally estimated to be 1–2 ×10⁻⁹ 526 cm^3s^{-1} [125, 126]. Theoretically, the o-p conversion probability was initially 52 evaluated to be about 10^{-10} cm³ s⁻¹ [124]. Later, the conversion rate for J=1528 $\rightarrow 0$ was more accurately evaluated to be 2.2×10^{-10} cm³ s⁻¹ by statistical 529 theory [127], and $J=0 \rightarrow 1,3$ excitation probability was calculated in terms of 530 S-matrix [120]. Recent studies by Honvault et al. employed a fully quantum 531 time-independent approach combined with a high accuracy ab initio potential 532 energy surfaces (PES) to derive the cross section as a function of the collision 533 energy [128, 129]. The rate coefficient is calculated to be $4.15 \times 10^{-10} \text{ cm}^3 \text{s}^{-1}$ 534 at 10 K. 535

As discussed in Sec. 6, H and H⁺ are dominant species in an astronomical environment, which induce o-p conversion via proton exchange as described above. In addition to H and H⁺, H₂ might encounter H₃⁺ or o-H₂ in space. While H₃⁺ would also induce o-p conversion via proton exchange [126, 130], o-H₂ + o-H₂ might temporarily form H₄ [131], which would eventually lead to two p-H₂ molecules. The proton exchange processes might be significant on surfaces when co-adsorbed with these species.

543 3.2. Basic interaction inducing o-p conversion

The nuclear wavefunctions of the ortho and para hydrogen are as shown 544 in Eqs. 1 and 2 described by the products of the nuclear-spin function and 545 rotational-vibrational wavefunction, which are functions of the spin coordi-546 nates $(\vec{i_a} \text{ and } \vec{i_b})$ and position coordinates $(\vec{R_a} \text{ and } \vec{R_b})$ defined in Fig. 1, 54 respectively. Although there is a coupling between the nuclear spin and ro-548 tational angular momentum [132], the spin and rotational parts are described 549 in a separated form here, because the coupling term is small and the physical 550 picture for transition is better understandable. For the o-p conversion of H_2 , 551 the spin state is changed from the triplet state to the singlet state, while the 552 rotational state is changed from a J= odd state to a J= even state. As shown 553 below, H' must contain \vec{i} ($\vec{i}=\vec{i_a}-\vec{i_b}$) and \vec{R} ($\vec{R}=\vec{R_a}-\vec{R_b}$) for the spin-state 554 transition and rotational-state transition, respectively. Here, the calculation 555 is shown for the H_2 case below, yet the results are essentially the same for 556 D_2 . 557

Any perturbation Hamiltonian containing operators $\vec{i_a}$ and $\vec{i_b}$ is generally expressed by the following formula with coefficients A and B, and can be described by functions of $(\vec{i_a} + \vec{i_b})$ and $(\vec{i_a} - \vec{i_b})$:

$$H' = A\vec{i_a} + B\vec{i_b} = \frac{A+B}{2}(\vec{i_a} + \vec{i_b}) + \frac{A-B}{2}(\vec{i_a} - \vec{i_b}).$$
 (10)

We consider the matrix element of H' between the nuclear-spin functions χ_n in Fig. 3. Since χ_n 's are the eigenstates of the total nuclear spin $\vec{I} (=\vec{i_a} + \vec{i_b})$, the matrix element of $(\vec{i_a} + \vec{i_b})$ between $\chi_n(I = 1)$ and $\chi_n(I = 0)$ is zero. On the other hand, the matrix element of \vec{i} between $\chi_n(I = 1)$ and $\chi_n(I = 0)$ is non-zero:

$$<\chi_n(I=0)|i^{\pm}|\chi_n(I=1, I_z=\mp 1)> = \mp\sqrt{2} <\chi_n(I=0)|i^{z}|\chi_n(I=1, I_z=0)> = 1,$$
(11)

where i^{\pm} denotes the raising and lowering operators divided by \hbar . Therefore, the perturbation of $\vec{i}=\vec{i_a}-\vec{i_b}$ mixes the ortho and para states leading to finite transition probability (see Appendix A).

On the other hand, the position coordinates \vec{R} $(\vec{R}=(x,y,z)=(R,\theta,\phi))$ are expressed in the spherical expression as

$$R^{\pm 1} = \mp \frac{x \pm iy}{\sqrt{2}} = R \sqrt{\frac{4\pi}{3}} Y_{1,\pm 1}$$
$$R^{0} = z = R \sqrt{\frac{4\pi}{3}} Y_{1,0}.$$
(12)

In a similar way to the spin functions, the matrix element of $R^{\pm 1,0}$ between the rotational states of ortho and para H₂, $Y_{1,M}$ and $Y_{0,0}$, is described as [68]

$$< Y_{0,0}|R^{\pm 1}|Y_{1,\mp 1}> = -\frac{R}{\sqrt{3}}$$

 $< Y_{0,0}|R^{0}|Y_{1,0}> = \frac{R}{\sqrt{3}}.$ (13)

Although the above formula is given for the transition between the J=1 and J=0 states, it is easily generalized for arbitrary J states.

Finally, the matrix element of Eq. 13 between the vibrational wavefunctions $(L_v(R))$ is evaluated with the equilibrium interatomic distance of R_e as

$$< L_{v'}(R)|R|L_v(R) > = < L_{v'}(R)|R_e|L_v(R) > + < L_{v'}(R)|(R-R_e)|L_v(R) > .(14)$$

Here, the first and second terms of Eq. 14 give the selection rule of $\Delta v=0$ and ± 1 , respectively, for a harmonic oscillator.



Figure 15: Schematic illustration of a hydrogen molecule under an inhomogeneous magnetic field \vec{h} expressed by dashed curves. $\vec{i_a}$ and $\vec{i_b}$ denote the nuclear spins of nuclei a and b, respectively. \vec{R} and $\vec{R_c}$ represent the relative coordinate vector and the center of mass coordinate vector, respectively, i.e. $\vec{R}_a = \vec{R}_c + \vec{R}/2$ and $\vec{R}_b = \vec{R}_c - \vec{R}/2$. The inhomogeneous magnetic field is expressed by a dipole field originating from a magnetic dipole moment $\vec{\mu}_a$.

- 3.3. First-order perturbation 580
- 3.3.1. Perturbation Hamiltonian 58

Magnetic dipole interaction 582

When there is an inhomogeneous magnetic field as schematically shown 583 in Fig. 15, the perturbation is expressed as 584

$$H' = g_n \beta_n (\vec{i_a} \cdot \vec{h_a} + \vec{i_b} \cdot \vec{h_b}) = \frac{g_n \beta_n}{2} [(\vec{i_a} + \vec{i_b}) \cdot (\vec{h_a} + \vec{h_b}) + (\vec{i_a} - \vec{i_b}) \cdot (\vec{h_a} - \vec{h_b})]$$
(15)

where $\vec{h_a}$ and $\vec{h_b}$ denote the magnetic field at the nuclei (a and b), and g_n and 585 β_n are the nuclear g-factor and nuclear magneton, respectively. Following the 586 discussion in Sec. 3.2, the second term contributes to the spin conversion. 587 Since $(\vec{h_a} - \vec{h_b}) \cong \vec{R} \cdot \nabla \vec{h}(\vec{R_c})$ as the first term of the Taylor expansion, it is 588 seen that this term contains the operator \vec{R} [21]. When the magnetic field is 589 inhomogeneous, i.e. $\nabla \vec{h}(\vec{R}_c) \neq 0$, this term is effective for o-p conversion. 590 591

In Fig. 15, the magnetic field is expressed by the dipolar field due to a

⁵⁹² point magnetic moment near H_2 as

$$\vec{h}(\vec{r}) = \frac{1}{4\pi\mu_0} \left(\frac{\vec{\mu_a}}{|\vec{r}|^3} - \frac{3\vec{\mu_a}\cdot\vec{r}}{|\vec{r}|^5}\vec{r} \right), \tag{16}$$

which indicates that the field can be inhomogeneous on an atomic scale. The magnetic moment originates from either spins or orbital angular momenta of electrons and nuclei. Since the electron Bohr magneton is $\sim 10^3$ as large as the nuclear magneton, the o-p conversion due to the nuclear spin tends to be $\sim 10^6$ as slow as that due to the electron spin. This interaction was considered for the o-p conversion due to interaction of magnetic impurities, magnetic materials, and solid H₂.

The magnetic dipole moment was usually treated as a point dipole. Electrons are however spatially extended as expressed by wavefunctions. The effect of the extended feature of electrons on the dipole interaction was theoretically examined, and was shown to possibly enhance the o-p conversion rate [133, 134].

605 Fermi contact interaction

When there is a finite electron density at the nucleus position, there occurs the Fermi contact interaction. With the electron spin density operator $(\vec{s}(\vec{r}))$ and one electron density operator $(\rho(\vec{r}))$, the Fermi contact interaction of the electron with the two nuclei of H₂ is described as

$$H' = \frac{8\pi}{3} \frac{\mu_0}{4\pi} g_n \beta_n g_e \beta_e \{ \vec{s}(\vec{R}_a) \cdot \vec{i}_a + \vec{s}(\vec{R}_b) \cdot \vec{i}_b \}$$

$$\cong \frac{2\mu_0}{3} g_n \beta_n g_e \beta_e [\{ \vec{s} \cdot (\vec{i}_a + \vec{i}_b) \} \{ \rho(\vec{R}_a) + \rho(\vec{R}_b) \}$$

$$+ \{ \vec{s} \cdot (\vec{i}_a - \vec{i}_b) \} \{ \rho(\vec{R}_a) - \rho(\vec{R}_b) \}], \qquad (17)$$

where g_e and β_e are the electron g-factor and Bohr magneton, respectively 610 [21]. Since the surface electron wavefunction spills over from the surface ex-611 ponentially as schematically shown in Fig. 8, H_2 in the physisorption well 612 suffers from the Fermi contact interaction. As the magnetic dipole interac-613 tion, the electron density difference in the second term of the Hamiltonian 614 can be approximated as the first term of the Taylor expansion with respect to 615 the position vector as $\vec{R} \cdot \nabla \rho(\vec{R_c})$ [21, 135]. Then, it is shown that the pertur-616 bation is significant when the electron density is inhomogeneous containing 617 both \vec{R} and $\vec{i_a} - \vec{i_b}$, and mixes the ortho and para states. 618

In addition to the direct Fermi contact interaction, the overlap between the surface electron orbital and H_2 orbital was shown to be significant for the Fermi contact interaction leading to enhancement of the o-p conversion rate [135, 133]. Furthermore, the intramolecular Fermi contact interaction becomes important when considering higher-order perturbation through interaction with surfaces, which will be described in Sec. 3.4.

625 Electron motion

If the electron motion is taken into consideration, there occurs an interaction between the nuclear spin and the magnetic field originating from the electron motion. As is usual the case for the interaction with the electromagnetic field, this interaction can be described as

$$H' = \frac{e}{2m_e} \sum_{i=a,b} [\vec{p} \cdot \vec{A}(\vec{r} - \vec{R}_i) + \vec{A}(\vec{r} - \vec{R}_i) \cdot \vec{p}]$$
(18)

where $\vec{A}(\vec{r})$ is the vector potential originating from the magnetic field pro-630 duced by hydrogen nuclear spins, and \vec{p} and m_e denote the electron momen-631 tum and mass, respectively [21]. Since $\vec{A}(\vec{r})$ is described by the outer product 632 of the magnetic moment and the position vector, the perturbation Hamilto-633 nian has a term proportional to $[(\vec{i_a} - \vec{i_b}) \times (\vec{R} \cdot \nabla)\vec{D}] \cdot \vec{p}$ with $\vec{D} = (\vec{r} - \vec{R})/|\vec{r} - \vec{R}|^3$ 634 [21]. As in the cases above, it mixes the ortho and para states. This interac-635 tion was named "spin-orbit interaction" in the paper by Yucel [21], because 636 the interaction between the nuclear spin and electron orbital motion is taken 637 into account. Nevertheless, it should be noted that this is not the ordinary 638 spin-orbit interaction that appears as a relativistic effect [136]. 639

640 Quadrupole interaction

The deuteron nucleus has a quadrupole moment Q_{μ} as shown in Table 4, which interacts with the electric-field gradient V_{μ} [137, 138, 139]. The interaction Hamiltonian of D₂ with the electric-field gradient is expressed in the spherical representation as

$$H' = \sum_{\mu=0,\pm1,\pm2} [V_{\mu}(\vec{R_{a}})Q_{\mu}^{a} + V_{\mu}(\vec{R_{b}})Q_{\mu}^{b}]$$

$$= \frac{1}{2} \sum_{\mu=0,\pm1,\pm2} [V_{\mu}(\vec{R_{a}}) + V_{\mu}(\vec{R_{b}})](Q_{\mu}^{a} + Q_{\mu}^{b}) + [V_{\mu}(\vec{R_{a}}) - V_{\mu}(\vec{R_{b}})](Q_{\mu}^{a} - Q_{\mu}^{b})$$
(19)



Figure 16: Schematic illustration of the quadrupole interaction with an electric-field gradient. The electric-field gradient is expressed by four point charges, and the ellipsoidal charge distribution aligns the spin direction lowering the energy of the $i_z = \pm 1$ state.

Table 4: Nuclear g-factor and qu	iadrupole moment (Q	<i>Q</i>) of proton	and deuteron.
----------------------------------	------------------------	----------------------	---------------

	g- factor	$Q \; (e \cdot \mathrm{fm}^2)$
proton	+5.6	0
deuteron	+0.86	+0.28

As the above treatments, the second term in the summation is approximated as $[\vec{R} \cdot \nabla V_{\mu}(\vec{R}_c)](Q^a_{\mu} - Q^b_{\mu})$ in the first term of the Taylor expansion. Since Eq. 19 contains \vec{R} and $Q^a_{\mu} - Q^b_{\mu}$ has non-vanishing matrix elements between the $\chi_n(I=1)$ of p-D₂ and $\chi_n(I=2)$ of o-D₂ [139], the quadrupole interaction induces the para-ortho (p-o) conversion of D₂.

The quadrupole interaction is qualitatively understood by Fig. 16. While the nuclear-charge distribution is spherical for the proton, the deuteron has a non-spherical charge distribution and the spin direction is aligned to the major axis of the nucleus due to the relative motion of the proton and neutron in the nucleus. If there is an electric field gradient as represented by the charge distribution shown in Fig. 16, the electrostatic energy depends on the spin direction leading to lifting of the degeneracy.

657 3.3.2. Transition probability

658 Wigner law

⁶⁵⁹ Following the time-dependent perturbation theory, the amplitude of a

660 p-H₂ state, C_p , changes as [136]

$$i\hbar \frac{dC_p}{dt} = \langle p|H'|o\rangle \exp(-i\omega_{op}t).$$
⁽²⁰⁾

Here, $\hbar \omega_{op}$ represents the energy difference between the initial ortho $(|o\rangle)$ and final para $(|p\rangle)$ states, which is denoted by the o-p energy. By integrating the above equation, the probability to find the para state at time t can be expressed as

$$|C_p|^2 = \frac{|\langle p|H'|o\rangle|^2}{\hbar^2} \frac{\sin^2(\omega_{op}t/2)}{(\omega_{op}/2)^2}.$$
(21)

When H' is the dipole magnetic field due to a paramagnetic impurity with a magnetic moment of μ_a as schematically shown in Fig. 15, $H' \propto \nabla \vec{h}(\vec{R_c})$ is proportional to a^{-4} $(a = |\vec{R_c}|)$.

To evaluate the transition probability, it is assumed that the perturbation is present within a finite time. Wigner assumed that a corresponds to the interaction region around a paramagnetic impurity, and that the converting hydrogen molecule moves across this region with a thermal velocity of v [140]. The duration time t of the perturbation is then expressed as t = a/v, which leads to the transition probability P as

$$P = \frac{\mu_a^2 (g_n \beta_n)^2 K}{3h^2 \mu_0^2 a^6 M_v v^2}.$$
(22)

Here, $K = M_p R^2/2$ is the moment of inertia of the molecule, and the kinetic energy is taken as $M_p v^2 = 3k_B T$ with the proton mass of M_p and temperature of T. This conversion rate is proportional to μ_a^2/a^6 , which is referred to as the Wigner law. This was later applied to a heterogeneous physisorption system [141].

679 Spectral density

In the Wigner formalism, the o-p energy was simply assumed to be accommodated by the kinetic energy of the molecule. This energy dissipation/excitation can be more rigorously treated by considering the transition from an energetically discrete state to a continuum state with a density of states. The transition probability can be described on the basis of the timedependent perturbation theory as [136]

$$P = \frac{2\pi}{\hbar} |\langle f | H' | i \rangle|^2 J(\hbar \omega_{op}).$$
⁽²³⁾

Here, $J(\hbar\omega_{op})$ represents the density of states with an energy of $\hbar\omega_{op}$, and $|i\rangle$ and $|f\rangle$ are the initial ortho and final para states including the degrees of freedom for energy dissipation.

The density of states was theoretically treated by taking account of the 689 time evolution of the molecule position, where the correlation of the molecule 690 position gives rise to a spectral density. In addition to the translational 691 motion of the molecule, the paramagnetic spin correlation was also taken into 692 consideration implying that the spin system absorbs/emits the o-p energy 693 via spin waves. These were first discussed by Leffler [142] by fixing the 694 two-nuclei position without considering the rotational motion [143]. On the 695 same ground, the o-p conversion was dealt with in terms of the nuclear-69 spin relaxation [144]. However, the rotational motion is essential for the o-p 69 conversion, and must be taken into consideration. 698

The energy dissipation/excitation upon o-p conversion was later more 699 accurately discussed by Ilisca, and the relative importance of the translational 700 motion and spin waves was compared in detail [145, 146]. Petzinger and 701 Scalapino then gave the absolute conversion rate on the same ground [132]. 702 The idea of the o-p energy transfer to the spin system was further discussed 703 for the o-p conversion on magnetic surfaces. It was pointed out that the 704 spin relaxation time gets shorter above the phase transition temperature of 705 magnetic surfaces and becomes significant for the energy transfer thereby 706 enhancing the o-p conversion [147, 148, 14]. In relation to the o-p conversion 707 on magnetic surfaces, it is known that there are significant effects of the 708 external magnetic field on the o-p conversion, which acquired much attention 709 and was reviewed in the article by Ilisca [10]. After publication of the review 710 article, theory for the magnetic-field effect was developed in terms of the 71 efficiency of the o-p energy transfer into the electronic system [15, 16]. 712

On the basis of Eq. 23, Motizuki and co-workers calculated the probabil-713 ities of the o-p conversion of H_2 and p-o conversion of D_2 when the molecules 714 are solidified by taking account of the phonon density of states [149, 139]. 715 The perturbation Hamiltonian was expanded in powers of the displacement 716 of the molecule center of mass, which allowed evaluation of the phonon emis-717 sion/absorption matrix elements. By treating the lattice vibration within 718 the Debye approximation and calculating the phonon density of states, the 719 o-p transition probability was obtained, which corresponds to the o-p energy 720 dissipation into the lattice phonon. The calculated conversion rate for H_2 was 72 1.9 % per hour [149], which was in good agreement with the experimental 722 results [150]. The experimental p-o conversion rate of D_2 , on the other hand, 723

⁷²⁴ is 5.5×10^{-4} /h [151, 152]. The theoretical study showed that the conversion ⁷²⁵ rate due to the magnetic dipole interaction is the same order as the rate due ⁷²⁶ to the quadrupole interaction [139].

In a high pressure condition, the phonon density of states of solid hydrogen changes as compared to those at low pressures, and the rotational sublevel splitting is caused by the anisotropic potential due to the electric quadrupole– quadrupole interaction [3]. These effects modify the energy dissipation channel significantly influencing the conversion rate [153, 154, 155, 156, 157]. The dependence of the conversion probability on the magnetic sublevel is discussed in relation to the NMR spectrum [158].

When H_2 is in a fixed position with a distance of *a* from a paramagnetic 734 spin, the o-p conversion rate is expected to be proportional to $|\langle p|H'|o\rangle|^2 \propto$ 735 a^{-8} in contrast to the Wigner law of a^{-6} where H₂ is assumed to move 736 across the interaction region of a paramagnetic impurity. This was recently 73 confirmed by a sophisticated experiment for H_2 encapsulated in C_{60} [159]. 738 Figure 17(a) shows a schematic of their samples. A paramagnetic spin cata-739 lyst was covalently linked to C_{60} with spacer molecules (distance a) between 740 the encapsulated H_2 and the paramagnetic molecule. For various spacer dis-74 tances, the para to ortho conversion time was measured at room temperature 742 with NMR after thermalized at 77 K, which showed a clear a^{-8} dependence 743 of the conversion time as shown in Fig. 17(b) [159]. The interaction time 744 was treated by the coherence time of the impurity spin [146]. 745

For the physisorbed H_2 , Yucel calculated the o-p conversion probability 746 by taking account of the two-dimensional phonon of the adsorbed H_2 layer as 747 the o-p energy dissipation channel [21]. As the perturbation, the magnetic 748 dipole, Fermi contact and orbital motion of metal surface electrons were con-749 sidered. The calculated conversion time is shown in Fig. 18 as a function of 750 the H_2 -surface distance. Here, the surface was assumed to be diamagnetic 751 without any paramagnetic ions. This study was motivated by the experi-752 mental findings of possible o-p conversion on diamagnetic metal surfaces of 753 Cu and Ag as detailed below. Although the surface is diamagnetic, the mag-754 netic dipole and Fermi contact interaction become significant due to thermal 755 fluctuation of the metal electron. 756

757 3.4. Second-order perturbation

When H₂ is in a physisorption well, the second-order perturbation can also be important as well as the first-order term. This was first proposed by Ilisca and Sugano [160, 161], who recognized the experimentally observed



^aBlue balls indicate a mixture of encapsulated H₂ and HD.



Figure 17: (a) Schematic illustration of H_2 encapsulated in C_{60} . A paramagnetic molecule of NO is linked to C_{60} at a distance of a with various spacer molecules. (b) Para to ortho conversion rate of H_2 in C_{60} induced by the paramagnetic molecule at room temperature. The conversion rate is shown to be proportional to a^{-8} . Adapted by permission from [159].



Figure 18: The ortho-para conversion times as a function of the metal-molecule distance z_0 on a Cu surface for different processes of the magnetic dipole (τ_D) , Fermi contact (τ_{FC}) and electron motion (τ_{SO}) calculated by Yucel. Reproduced by permission from [21].
conversion rate on oxide surfaces was around two orders of magnitude larger
than the value estimated on the basis of the first-order perturbation. According to the perturbation theory, the second-order transition probability
via intermediate states is expressed as [136]

$$P = \frac{2\pi}{\hbar} \left| \sum_{I} \frac{\langle f | H' | I \rangle \langle I | H' | i \rangle}{E_i - E_I} \right|^2 J(E_f = E_i).$$
(24)

⁷⁶⁵ Ilisca and Sugano considered two interactions: Coulomb interaction (H_C) ⁷⁶⁶ expressed as Eq. 25 and Fermi contact interaction (H_{FC}) described by Eq. ⁷⁶⁷ 17 as $H' = H_C + H_{FC}$.

$$H_C = \frac{e^2}{4\pi\epsilon} \left(\sum_i \sum_j \frac{Z_j}{|\vec{r_i} - \vec{R_j}|} + \sum_{i,j(i \neq j)} \frac{1}{|\vec{r_i} - \vec{r_j}|}\right).$$
 (25)

Here, $\vec{r_i}$ and $\vec{R_j}$ represent the electron and nucleus positions including sur-768 faces, and Z_j is the atomic number of the j nucleus. They showed that the 769 o-p conversion occurs as a two-step process following eq. 24 where H_C vir-770 tually induces excitation to an electronically excited state of the H_2 -surface 771 complex and H_{FC} causes nuclear-spin flip through interaction of the elec-772 tronic spin in the virtually excited state. Following the original paper by 773 Ilisca and Sugano, this second-order process is called "two-step process" in 774 this article. 775

Figure 19(a) schematically shows the electronic and spin state of the sur-776 face paramagnetic ion and hydrogen molecule in the o-p conversion process. 777 The paramagnetic ion is shown to possess a spin 1/2, which does not lose 778 any generality. Since the hydrogen molecule is physisorbed, the electronic 779 state of the molecule is assumed to be retained. The $1s\sigma_a$ and $1s\sigma_u$ orbitals 780 of H_2 are denoted by g and u in the figure. With the Coulomb interaction, an 781 electron in the $1s\sigma_g$ orbital is excited to the surface and the electron in the 782 surface is excited to $1s\sigma_u$ while the total electron spin of the system is kept at 783 1/2. This corresponds to excitation of the molecule to the first excited state 784 of $b^3 \Sigma_u^+$ in the intermediate state. Because the orbital wavefunction of the 785 $b^{3}\Sigma_{u}^{+}$ is ungerade, the matrix element between the initial and intermediate 786 states has an antisymmetric character with respect to the permutation of the 787 two nuclei as shown in Table 3. This leads to a selection rule of even-odd 788 transition (ΔJ =odd) [160]. 789

In the second step, on the other hand, the matrix element of the Fermi contact interaction over the intermediate $b^3 \Sigma_u^+$ and final $X^1 \Sigma_q^+$ states leads 792 to

$$\langle {}^{3}\Sigma_{u}^{+}|H_{FC}|{}^{1}\Sigma_{g}^{+}\rangle = \lambda_{IFC}(\vec{s}_{1}-\vec{s}_{2})\cdot(\vec{i}_{a}-\vec{i}_{b})$$

$$\lambda_{IFC} = \int \phi_{1s\sigma_{g}}(\vec{r})\delta(\vec{r}-\vec{R}_{a})\phi_{1s\sigma_{u}}(\vec{r})d\vec{r}.$$

$$(26)$$

Here, the coefficient λ_{IFC} is described by the integration over the molecular 793 orbitals (see Appendix B), therefore this is called the intra-molecular Fermi 794 contact interaction. As discussed in Sec. 3.2, the spin difference operator 795 mixes the spin triplet and singlet states, which means that the electron spin 796 triplet and nuclear-spin triplet state can be mixed with the electron spin 797 singlet and nuclear-spin singlet state. As an overall process, $\Delta J = 1$ in the 798 first step and $\Delta I = 1$ in the second step results in o-p conversion, which 799 is schematically shown in Fig. 19(a). This overall process is labeled the 800 XY process [160]. It should be noted that the o-p conversion is significant 80 only when H₂ is virtually excited to $b^3 \Sigma_n^+$ through interaction with surfaces 802 although the Fermi contact coupling is intramolecular. 803

An intriguing outcome of this study was that the proposed concept could 804 also be applied to diamagnetic metals [162, 22, 163]. Figure 19(b) shows the 805 electronic state from the initial to final states in the case of a diamagnetic 806 metal, where the surface electrons form a singlet state in the initial state. In 807 the intermediate state, the surface electrons form a triplet state along with 808 the transition of the molecule state from ${}^{1}\Sigma_{g}^{+}$ to ${}^{3}\Sigma_{u}^{+}$ while the total spin of 809 the system is kept at 0. Note $\Delta J = \text{odd}$ in this first step. In the second step, 810 the molecule undergoes the transition from the triplet to singlet states for 811 both electrons and nuclei implying o-p conversion. 812

An extended version of this model was also proposed by considering an-813 other intermediate state, which is shown in Fig. 19(c) [22]. Here, the inter-814 mediate state is assumed to be a charge-transferred state, where H_2 forms 815 a negative ion with an additional electron in the $1\sigma_u$ orbital. In the second 816 step from the intermediate to final states in the figure, the molecule un-817 dergoes nuclear-spin conversion while the electron spin flips as it goes back 818 to the surface through the Fermi contact interaction. The electron config-819 urations of the intermediate $(|^{1}I\rangle)$ and final $(|^{3}f\rangle)$ states are expressed by 820 $(1s\sigma_u)^1(k)^1$ and $(k)^1(k')^1$, respectively, with the substrate states of k and 821 k', and the orbital integration $\langle {}^{3}f|H_{FC}|{}^{1}I\rangle$ has a spin antisymmetric com-822 ponent of $\lambda_{SFC}(\vec{s}_1 - \vec{s}_2) \cdot (\vec{i}_a - \vec{i}_b)$ (see Appendix B). This was labeled the 823 UY process [22]. It should be noted that the electrons relevant to the Fermi 824

contact interaction in this process are not exclusively those in the molecule, implying that the interaction is not an intramolecular one. There are image charge states on metal surfaces, which are furthermore shown to enhance the electron transfer thereby accelerating o-p conversion [164].

The conversion times calculated on the basis of these models are shown in Fig. 20 in comparison to those calculated in the first-order (one-step) processes, which obviously shows that the two-step models of XY and UYare more effective than the one-step mechanisms [22].

The electron transfer between H_2 and the substrate and the magnetic 833 dipole and direct Fermi contact interactions are obviously dependent on the 834 molecule orientation as well as the molecule distance from the surface as 835 intuitively understood in Fig. 8. Dependence of the o-p conversion on the 83 molecular orientation was theoretically examined on the basis of the one-step 837 process via the direct Fermi contact [165, 166] and two-step processes for the 838 system of paramagnetic impurities adsorbed on a surface [167, 168, 169, 170] 839 and metal surfaces [171, 172]. It was shown that the o-p conversion is more 840 efficient when the molecular axis is aligned along the surface normal direction. 841 This indicates that the M=0 state of J=1, which is often designated the 842 cartwheel-like rotation, is preferred for the o-p conversion compared to the 843 $M = \pm 1$ states designated the helicopter-like rotation. When the adsorption 844 potential is anisotropic, the energy level is split depending on M as shown in 845 Fig. 9. When $V_2 < 0$, H₂ molecules tend to occupy the M = 0 state because 846 the energy level is lower, then the o-p conversion would be enhanced as 847 compared with the case of an isotropic potential [173]. When the anisotropic 848 potential is larger and the rotational motion is expressed by the hindered 849 rotation, the orientational dependence will be more pronounced because of 850 the distortion of the rotational wavefunction. 851

852 3.5. Higher-order perturbation

Although the theory based on the second-order perturbation was success-853 ful, this concept may not be straightforwardly applied to diamagnetic insula-854 tors. In Fig. 19(b) and (c), the final state of the substrate is an electron spin 855 triplet state. If the substrate were a diamagnetic insulator, transition to the 856 triplet state requires a band-gap energy, which is much larger than the o-p 85 energy. This problem was shown to be circumvented by considering further 858 higher-order perturbation including the Stark coupling and spin-orbit cou-859 pling [33], which is called a multi-step process in analogy with the two-step 860 process. 861



Figure 19: Schematic illustration of the electronic excitation for the two-step o-p conversion model. From left to right, the initial, intermediate and final electronic states are shown, each consisting of the substrate and H₂ electronic levels. (a) XY process on a paramagnetic surface, (b) XY process on a diamagnetic metal, and (c) UY process on a diamagnetic metal. The initial o-H₂ (I=1) undergoes charge exchange with surfaces labeled C in the first step from the initial to intermediate state, and in the second step nuclear-spin conversion of $I = 1 \rightarrow I = 0$ through the Fermi contact interaction labeled Y.



Figure 20: The ortho-para conversion times as a function of the metal-molecule distance d (a. u.) for different processes on the basis of one-step (O: electron motion, W: magnetic dipole, Y_d: direct Fermi contact, and Y_{ind}: indirect Fermi contact via molecular orbitals) and two-step (UY and XY) processes calculated for an Ag surface by Ilisca. Reproduced by permission from [22].



Figure 21: Schematic illustration of mixing between o-H₂ and p-H₂ in the multi-step process. (a) Intrinsic mixing of $a^{3}\Sigma_{g}^{+}$ and $B^{1}\Sigma_{u}^{+}$ with intramolecular Fermi contact interaction, and enhanced spin-orbit coupling between ³II and ¹II. (b) Stark mixing between the gerade and ungerade states due to intense surface electric fields leads to an overall mixing between o-H₂ and p-H₂.

Figure 21(a) shows the excited states of H₂ relevant to such a multistep process. As discussed in the two-step process, the intramolecular Fermi contact interaction mixes the ${}^{3}\Sigma_{u}^{+}$ of o-H₂ and ${}^{1}\Sigma_{g}^{+}$ of p-H₂ (or vice versa). This also holds in the case of ${}^{3}\Sigma_{g}^{+}$ and ${}^{1}\Sigma_{u}^{+}$. Since the energy levels of $a^{3}\Sigma_{g}^{+}$ and $B^{1}\Sigma_{u}^{+}$ are closer than those of $b^{3}\Sigma_{u}^{+}$ and $X^{1}\Sigma_{g}^{+}$, the mixing coefficient is larger for $a^{3}\Sigma_{q}^{+}$ and $B^{1}\Sigma_{u}^{+}$ [174].

Another intramolecular interaction to be considered is spin-orbit coupling between ${}^{3}\Pi$ and ${}^{1}\Pi$ with the same electron configuration as shown in Fig. 22(a). The spin-orbit interaction Hamiltonian for a hydrogen molecule with two electrons is described as

$$H_{SOC} = \xi(\ell_1^z s_1^z + \ell_2^z s_2^z) \tag{27}$$

where ℓ_i^z and s_i^z denote the axial components of the angular momentum and spin of *i* electron, respectively, and ξ represents the spin-orbit coupling coefficient. Note that only the axial component of the electron angular momentum is well-defined in the diatomic molecule. This perturbation allows mixing between the total electron spin (S) singlet and triplet states with a selection rule of $\Delta S=1$ as detailed in Appendix C. The mixing coefficient is described as $\langle {}^{3}\Pi | H_{SOC} | {}^{1}\Pi \rangle / \delta_{t-s}$ with the energy difference δ_{t-s} between the ${}^{3}\Pi$ and ${}^{1}\Pi$



Figure 22: (a) Electron configuration of ${}^{1}\Pi$ and ${}^{3}\Pi$ states. (b) Adiabatic potentials for ${}^{c^{3}}\Pi$ and ${}^{C^{1}}\Pi$ for $(1s\sigma_{g})(2p\pi_{u})$, and ${}^{r^{3}}\Pi$ and ${}^{R^{1}}\Pi$ for $(1s\sigma_{g})(4d\pi_{g})$. The r and R states are energetically almost degenerate due to weak electron exchange-correlation.

states. This indicates that the mixing is larger as δ_{t-s} gets smaller.

The δ_{t-s} value is essentially determined by the Coulomb repulsion of the 880 two electrons. In a high Rydberg state where the electron orbital has an 88 extended feature, δ_{t-s} is expected to be small, i.e. weak electron exchange-882 correlation enhances the spin-orbit coupling. As shown in Fig. 22(b), the 883 adiabatic potentials of the $\mathbf{r}^3\Pi_g$ and $\mathbf{R}^1\Pi_g$ states are almost degenerate with 884 an energy difference of $\sim 0.1 \text{ meV}$ [175, 176]. Although the spin-orbit interac-885 tion strength is small for light elements like H, the mixing can be significant 886 in highly excited states due to the small energy difference. 887

These mixing properties are schematically shown in Fig. 21(a). The mixing of $a^{3}\Sigma_{g}^{+}$ and $B^{1}\Sigma_{u}^{+}$ with IFC and $r^{3}\Pi_{g}$ and $R^{1}\Pi_{g}$ with SOC is an inherent property of the molecule. Nevertheless, the o-p conversion is not induced without external perturbation because the ground state $X^{1}\Sigma_{g}^{+}$ is not mixed with the relevant excited states.

When H_2 is under an electric field, there occurs Stark coupling between 893 the electronically excited states of H_2 . The perturbation Hamiltonian (H_S) 894 is equivalent to Eq. 9. The matrix element of this perturbation is non-895 zero between gerade and ungerade states. Simultaneously, the molecular 896 rotational states of even and odd are mixed as the Coulomb interaction in 897 the two-step process. This situation is displayed in Fig. 21(b), where the 898 ortho and para states are mixed via intrinsically coupled B – a and r –R 899 states. The overall selection rule becomes $\Delta S = 0$, $\Delta J = 1$, and $\Delta I = 1$, 900

⁹⁰¹ and o-p conversion can be an allowed process [33].

902 3.6. Isotope effect

The isotopes of H_2 and D_2 have particular differences for the nuclear-spin conversions. This section summarizes the key factors governing the isotope dependence of the nuclear-spin conversion. While the electronic structure of H_2 and D_2 is essentially the same, the nuclei of proton and deuteron have different masses, nuclear spins, nuclear g factors, and nuclear quadrupole moments as shown in Table 4.

The mass difference leads to the difference in the rotational and vibra-909 tional energies. Whereas the rotational energy is different by a factor of 2, 910 the difference in the vibrational energy is $\sqrt{2}$. For the nuclear-spin conver-911 sion, the rotational-energy has to be dissipated into other degrees of freedom. 912 Since the rotational energy of D_2 is smaller than that of H_2 , the rotational-913 energy dissipation of D_2 might be more efficient than H_2 (Sec. 5.1.2). For 914 the conversion of solid hydrogen, in fact, the rotational energy of D_2 is con-915 sidered to be dissipated into the phonon system as a one-phonon process, 916 whereas that of H_2 requires two-phonon creation [149, 139]. In addition to 917 the efficiency of the rotational-energy dissipation, the rovibrational energy 918 is important for the level matching. In the two-step and multi-step conver-919 sion processes discussed in Secs. 3.4 and 3.5, intra-molecular coupling in the 920 excited state is important, and the coupling strength is enhanced when the 921 energy levels of the relevant states are near-degenerate. On the same adi-922 abatic potential, the level matching might be different between H_2 and D_2 923 because of the difference in the rovibrational energy. 924

The nuclear g factor, on the other hand, affects the magnetic interaction. 925 As shown in Eqs. 15 and 17, the magnetic interaction Hamiltonian is propor-926 tional to the nuclear g factor (g_n) , which is a common feature for both the 927 magnetic dipole and Fermi contact interaction. Since the total nuclear-spin 928 quantum number is different for H_2 and D_2 , the degeneracy of the initial and 929 final states in the nuclear-spin conversion is different for the two molecules 930 as detailed in Figs. 3 and 4. The matrix element between the initial and 931 final states, furthermore, depends on the spin quantum number and the z932 component of the spin. The square of the matrix element, when averaged 933 over the initial states and summed over the final states, is calculated to be 934 proportional to $\frac{g_n^2}{3}$ and $2g_n^2$ for H₂ and D₂, respectively, reflecting the initial 935 and final state degeneracy [21]. Considering $g_n = 5.59$ for proton and 0.857 936 for deuteron, both magnetic dipole and hyperfine contact interactions are 937

⁹³⁸ larger for H_2 than D_2 by a factor of 7.1, and the conversion time is expected ⁹³⁹ to be shorter for H_2 . In addition to the magnetic g factor, the quadrupole ⁹⁴⁰ interaction is significant for D_2 as deuteron has a non-zero quadrupole mo-⁹⁴¹ ment.

When intermolecular interaction is significant, the magnetic field due to 942 the nuclear spin and rotational moment of molecules contributes to the mag-943 netic interaction. While o-H₂ (I = 1, J = 1) has both nuclear-spin and 944 rotational moments, p-H₂ (I = 0, J = 0) has no magnetic moments. On the 945 other hand, both p-D₂ (I = 1, J = 1) and o-D₂ (I = 2, J = 0) contribute to 946 the magnetic dipole interaction. This difference leads to the difference of the 947 conversion kinetics: the H₂ o-p conversion follows the second-order kinetics, 948 while the D_2 p-o conversion is expressed by the first-order kinetics as demon-949 strated for solid hydrogen where the intermolecular magnetic interaction is 950 responsible [149, 139]. 95

⁹⁵² 4. Surface-sensitive experimental technique

The experimental technique that is capable of distinguishing $o-H_2$ and 953 $p-H_2$ is classified into two classes, one probing the nuclear spin of H_2 and 954 the other probing the rotational state of H_2 . The only technique of the 955 former class is the nuclear magnetic resonance (NMR) [177]. While only 956 $o-H_2$ with I=1 can be detected by NMR, the spin-lattice relaxation time is 95 dependent on the ortho concentration in a H_2 solid [178, 179], which allows 958 for estimation of the o-p ratio of samples. The problem at the present stage 959 is its sensitivity: the number of hydrogen atoms on ordinary well-defined 960 surfaces is 10^{15} , which is too small for the NMR measurement. 961

All other methods used for the distinction of the ortho and para species 962 are based on the fact that the two species have different rotational states 963 shown in Fig. 6. A traditional method used for o-p separation and conver-964 sion studies was the thermal conductivity measurement [1]. A wire in the 965 measurement cell was heated to 100 - 180 K, and the wire resistance was 966 measured at a certain H_2 gas pressure. Since the specific heat derived from 96 the partition function of Eq. 5 is different between the ortho and para species, 968 which is intuitively understood from the different rotational levels shown in 969 Fig. 6, the thermal conductivity is different between the two species, and 970 the resistance data gives the ortho concentration in the sample gas. This 97 type of measurement can be performed with Pirani gauges. [180, 181]. A 972 typical procedure of the experiments is as follows: H_2 gas is introduced in a 973

⁹⁷⁴ reaction chamber containing sample surfaces. From time to time, a part of ⁹⁷⁵ the H_2 gas is extracted from the reaction chamber and introduced into the ⁹⁷⁶ measurement cell, where the o-p ratio is evaluated by the thermal conductiv-⁹⁷⁷ ity measurement. Here, a large amount of H_2 is required for the study and a ⁹⁷⁸ large amount of surface area as the sample is necessary, which is not suited ⁹⁷⁹ to the study on well-defined single-crystal surfaces.

As described in Eq. 3, the internal energy of molecular hydrogen orig-980 inates from the electronic, vibrational and rotational degrees of freedom. 981 Since there is coupling of the rotational motion with the vibrational and 982 electronic states, rotational spectroscopy can be achieved by vibrational 983 excitation and electronic excitation accompanied by ro-vibrational excita-984 tion as well as pure rotational excitation. As H_2 has no permanent electric 98 dipole moment, pure rotational and vibrational spectroscopy can be carried 986 out by inelastic scattering of light and particles, such as Raman scattering 987 spectroscopy (Raman), Inelastic Neutron scattering spectroscopy (INS), and 988 Electron energy loss spectroscopy (EELS). In a particular situation, induced 989 infrared absorption spectroscopy (IRAS) can be used as exemplified below. 990

The schemes for rotational and vibrational excitations are schematically 991 shown in Fig. 23(a) and (b), respectively, and the excitation energies for 992 H_2 and D_2 are listed in Table 5. While $\Delta J=2$ transitions are observed in 993 pure rotational excitation of Raman and EELS because of the total-nuclear-994 spin conservation, INS can probe $\Delta J=1$ excitation, because a neutron has 995 a nuclear spin 1/2 and spin-flip scattering occurs through the nuclear force 996 [182]. In the vibrational excitation, on the other hand, $\Delta v=1$ excitation is 997 observed. Since the rotational constant depends slightly on the vibrational 998 level because of the stretch of the mean intra-molecular distance, the excita-999 tion energy is different between the J states as schematically shown in Fig. 1000 23(b), which allows us to distinguish ortho and para species. The excita-1001 tion energy difference is, however, small as listed in Table 5, therefore the 1002 vibrational spectroscopy is achieved by Raman and induced IRAS with high 1003 resolution. In EELS, on the other hand, vibrational excitation accompanied 1004 by rotational excitation is observed, which distinguishes the ortho and para 1005 species. 1006

As to the sensitivity, the scattering cross section of Raman and INS is small and the probing depth is long, hence these techniques are not suited to the study on well-defined surfaces. INS has been applied to bulk or powder/porous materials such as oxides [183, 77, 78], carbon [184, 185, 186], metal-organic framework [187] to date. Raman has also been applied to bulk



Figure 23: Schematic energy diagram of the rotational-energy levels in (a) the $X^1\Sigma_g^+$ state (v=0), (b) v=1 and 0 states of $X^1\Sigma_g^+$, and (c) $X^1\Sigma_g^+$ (v=0) and $E,F^1\Sigma_g^+$ (v=0) states. Solid (blue) and dashed (green) arrows denote excitations of the J=0 and J=1 states used in rotational-state spectroscopy. $\Delta J=0$ transitions are shown in (b) and (c). (c) is used for (2+1)REMPI (two-photon resonant excitation followed by one-photon ionization).

Table 5: Energy (meV) for pure rotational (J) and pure vibrational (v) excitations of H₂ and D₂ calculated by Eq. 3 using the constants listed in Tables 1 and 2. Wavelengths (μm) for corresponding photon energies are also shown in parentheses.

v	0 - 0		0 - 1	
J	0 - 2	1 - 3	0 -0	1 - 1
H_2	43.9	72.7	515.7	514.9
	(28.2)	(17.0)	(2.41)	(2.41)
D_2	22.2	36.9	370.9	370.7
	(55.9)	(33.6)	(3.34)	(3.35)

or powder/porous materials such as Si [188, 189], oxides [190] and carbon 1012 [191, 192, 193]. In EELS in a low-energy regime at a primary energy of 1– 1013 10 eV, electrons reflected from surfaces are detected after energy analysis. 1014 Hence, compared with the rather long probing depth of INS and Raman, 1015 EELS detects only surface-adsorbed molecules, which allows for rotational 1016 spectroscopy on well-defined surfaces. In the gas phase electron scattering 1017 experiments, it is known that there occurs a temporal negative-ion resonance 1018 at a primary electron energy of 3 eV, where the scattering cross section shows 1019 a maximum [194]. A similar resonance feature was also observed on the ph-1020 ysisorbed H_2 [195], therefore, the EELS experiments for H_2 on surfaces have 1021 often been done at a primary energy of about 3 eV. 1022

It should be noted that $\Delta J=2$ and $\Delta v=1$ radiative transitions are induced via the quadrupole moment of molecular hydrogen, which can be realized as infrared-light absorption and emission. Since the transition probability is small, however, IRAS due to the quadrupole transition is not utilized for detection of molecular hydrogen in laboratory experiments.

On solid surfaces, however, molecular hydrogen might be polarized par-1028 ticularly on ionic substrates owing to the strong electric field present on 1029 the surface. This induces a dipole moment in H_2 , which causes infrared 1030 light absorption due to vibrational excitation as depicted in Fig. 23(b). 1031 This was observed on porous NaCl, ice and glass having a large surface area 1032 [196, 197, 198, 199, 94], and has been applied to porous and bulk materials 1033 such as carbon materials [200, 201, 202], Si [203], metal-organic framework 1034 [204, 205, 206], and oxides [207]. An experimental progress has enabled to 1035 probe $o-H_2$ and $p-H_2$ physisorbed on single-crystal surfaces [208, 209]. 1036

The other type of techniques capable of rotational-state-resolved mea-1037 surements utilizes electronic excitation of molecules as shown in Fig. 23(c). 1038 Here, excitation to the E ${}^{1}\Sigma_{q}^{+}$ state is shown as an example. Similarly to the 1039 vibrational excitation in Fig. 23(b), the rotational constant in the E ${}^{1}\Sigma_{a}^{+}$ 1040 state is different from that in the $X^1\Sigma_q^+$ state as shown in Tables 1 and 2. 1041 Then, the excitation energy corresponding to the J=0 and J=1 molecules 1042 is slightly different, thereby allowing the rotational-state-resolved detection. 1043 Two experimental approaches to probe such electronically excited molecules 1044 are ionization and fluorescence, which are designated as resonance-enhanced 1045 multiphoton ionization (REMPI) and laser-induced fluorescence (LIF). The 1046 typically used resonant excitation is $X^1\Sigma_a^+ \to E, F^1\Sigma_a^+$ and $X^1\Sigma_a^+ \to B^1\Sigma_u^+$. 1047 The excitation, ionization and fluorescence processes are shown by arrows in 1048 the adiabatic potential (Fig. 24). 1049



Figure 24: Resonant excitation, ionization and fluorescence processes used for REMPI and LIF. (a) Two-photon excitation to E,F ${}^{1}\Sigma_{g}^{+}$ and subsequent ionization are shown by solid arrows, whereas the dashed arrow denotes fluorescence due to E,F ${}^{1}\Sigma_{g}^{+} \rightarrow B^{1}\Sigma_{u}^{+}$. (b) One photon excitation to $B^{1}\Sigma_{u}^{+}$ and subsequent ionization are shown by solid arrows, and fluorescence due to deexcitation to $X^{1}\Sigma_{g}^{+}$ is denoted by the dashed arrow.

¹⁰⁵⁰ Due to the symmetry restriction, excitation to the E,F ${}^{1}\Sigma_{g}^{+}$ state by one-¹⁰⁵¹ photon absorption is forbidden, and therefore can be achieved by two-photon ¹⁰⁵² absorption. In this excitation scheme, the selection rule for the rotational ¹⁰⁵³ state is $\Delta J=0$, ± 2 [210, 211, 212]. On the other hand, the B¹ Σ_{u}^{+} state is ¹⁰⁵⁴ accessible by a single-photon absorption due to its dipole-allowed nature with ¹⁰⁵⁵ respect to the X¹ Σ_{g}^{+} ground state [213, 214], and the selection rule for the ¹⁰⁵⁶ rotational state is $\Delta J=\pm 1$.

The excited molecule in these E, F ${}^{1}\Sigma_{q}^{+}$ and B ${}^{1}\Sigma_{u}^{+}$ states can be ionized by 1057 absorption of another photon [210, 211, 212, 213, 214]. While fluorescence 1058 due to the $B^1\Sigma_u^+ \to X^1\Sigma_g^+$ transition is dipole-allowed [215], the E,F¹ Σ_g^+ 1059 $\rightarrow X^1\Sigma_q^+$ transition is dipole-forbidden. However, the E, $F^1\Sigma_g^+$ state can be 1060 quenched through a dipole transition to the low-lying $B^1\Sigma_u^+$ state [216, 217, 1061 218], and the rotational-state resolved $E,F \rightarrow B$ fluorescence was recently 1062 achieved [37, 219]. REMPI and LIF have been applied to dynamics studies 1063 of hydrogen on surfaces as reviewed in Refs. [220, 221, 222]. 1064

A typical LIF data via the E state taken for H₂ at 300 K is shown in Fig. 25. Four distinct maxima are observed in the excitation wavelength range of about 202 nm. By comparing the energy levels for the X and E states calculated following Eq. 3, these maxima are assigned to the J=0, 1, 2, and 3 states.

¹⁰⁷⁰ The REMPI and LIF are essentially gas-phase spectroscopic techniques. ¹⁰⁷¹ REMPI can actually detect H_2 in the pressure range as low as 10^{-8} Pa. ¹⁰⁷² However, both techniques have not been directly applied to H_2 physisorbed ¹⁰⁷³ on a surface, because the electronic level of H_2 is modified and substantially ¹⁰⁷⁴ broadened in a physisorption well, which hampers the resonant excitation in ¹⁰⁷⁵ an internal-state-resolved manner.

To compromise this situation, the REMPI technique was combined with 1076 desorption methods where adsorbed H_2 was desorbed from the surface fol-1077 lowed by REMPI detection. Two desorption schemes were applied, photo-1078 stimulated desorption (PSD) [25, 223, 224, 225], and TPD [26, 33, 226]. 1079 The experimental setup of our group is schematically shown in Fig. 26. 1080 Well-defined sample surfaces prepared in UHV are exposed to H_2 or D_2 , and 1081 REMPI detects molecules upon TPD or PSD with another laser beam of 1082 2.3 - 6.4 eV (not shown). An important factor for this experiment is the 1083 desorption probability of the ortho and para species. Figure 27 shows the 1084 time of flight spectra of H_2 photodesorbed from Ag, which was obtained by 1085 changing the delay time of the REMPI laser from the pump desorption laser 1086 [25]. The spectra of J = 1 and J = 0 both reveal a maximum at a time 1087



Figure 25: Typical LIF data measured for H_2 via the $X \to E$ two-photon excitation followed by fluorescence via $E, F \to B$ [37]. The fluorescence intensity is plotted as a function of the excitation laser wavelength, and the four maxima correspond to resonant excitation of H_2 in the J=0, 1, 2, and 3 states.



Figure 26: Schematic illustration of the experimental setup for REMPI detection used in the group of University of Tokyo. The molecules either photodesorbed or thermally desorbed from the sample surface are ionized with the REMPI laser focused in front of the sample, and detected by a microchannel plate (MCP) through an ion lens. Tunable laser light for (2+1) REMPI via E at ~201 nm is generated by second-harmonic generation (SHG) and sum-frequency generation (SFG) of YAG-pumped dye laser at ~603 nm (line width: 0.15 cm⁻¹).

of flight of 22 μ s. It was argued that the desorption mechanism is the same 1088 independent of the rotational state, thus the desorption intensities probed by 1089 REMPI-PSD are expected to be proportional to the H₂ coverages on surfaces 1090 [25, 223]. Figure 28 shows typical REMPI-TPD results [33]: thermally des-1091 orbed H_2 was rotational-state selectively measured during sample heating, 1092 which corresponds to the sample temperature of 10 - 30 K. By integrating 1093 the desorption signals, the coverage of the ortho and para hydrogen can be 1094 estimated. In the PSD method, only a small portion of adsorbed H_2 is pho-1095 todesorbed followed by REMPI detection, which allows coverage estimation 1096 of the ortho and para species in real time. In the TPD method, on the other 1097 hand, the total amount of adsorbed H_2 is thermally desorbed from surfaces, 1098 hence repreparation of the sample is necessary to examine the time evolution 1099 of the ortho and para coverages. 1100



Figure 27: Time of flight spectra of H₂ in (a) J = 0 and (b) J = 1 states photodesorbed from an Ag surface at a pump laser wavelength of 193 nm (6.4 eV) as measured by REMPI-PSD. The spectra were collected by varying the delay of the probe REMPI laser pulse from the pump desorption laser. The solid curves denote fits with the Maxwell-Boltzmann distribution. Reproduced by permission from [25].



Figure 28: REMPI signals for H₂ during adsorption and desorption from amorphous ice surfaces in the J = 0 (blue open circle) and J = 1 (red cross) states. (a) n-H₂ was introduced at a pressure of 1.5×10^{-7} Pa into the chamber at t = 0 - 14 s, and subsequently the sample temperature was raised from t = 35 - 40 s. (b) the sample temperature was raised at t = 598 s after the n-H₂ dosage. Adapted by permission from [33].

¹¹⁰¹ 5. Physisorption and ortho-para conversion on various surfaces: ¹¹⁰² experimental studies

When hydrogen molecules are physisorbed on a surface, as shown in Fig. 1103 9, the molecule is trapped in the well with a binding energy of about 30 meV, 1104 which is realized at a low temperature of about 10 K, and the rotational level 1105 is split due to the potential anisotropy. At this low temperature, all $o-H_2$ 1106 $(p-D_2)$ molecules are in the ground state of the J=1 state, while all $p-H_2$ 1107 $(o-D_2)$ molecules are in the J=0 state. While trapped in the physisorption 1108 well, $o-H_2$ (p-D₂) interacts with the electronic and nuclear (phonon and spin) 1109 systems of surfaces and neighboring molecules, and undergoes o-p (p-o) con-1110 version relaxing to the J=0 state of p-H₂ (o-D₂), which is accompanied by the 1111 nuclear spin flip ($\Delta I=1$) and rotational-state transition ($\Delta J=1$). The ques-1112 tions on the conversion mechanism are (1) the origin of the perturbation and 1113 (2) the energy dissipation channel. As described in Sec. 3, possible origins 1114 of the magnetic interaction are the magnetic dipole, external and internal 1115 Fermi contact and electron motion. In combination with these interactions, 1116 electron Coulomb interaction, Stark and Spin-orbit interaction may be taken 1117 into consideration when two- or multi-step processes are considered. On vari-1118

ous surfaces, the origin of the perturbation has been discussed on the basis of 1119 surface electronic properties. On the other hand, possible energy dissipation 1120 channels are the substrate phonon/electron, overlayer H_2 phonon, localized 1121 H_2 vibration in the z direction, kinetic energy in the parallel direction (sur-1122 face diffusion), and magnon. While theoretical studies examine the energy 1123 dissipation channel so as to fit the experimental conversion time as exem-1124 plified for solid H_2 [149, 154, 157], little experimental information has been 1125 reported on the energy dissipation mechanism, which should be the subject 1126 in the future. 1127

As introduced in Sec. 4, EELS and REMPI studies have clarified the 1128 physisorption and o-p conversion of hydrogen mainly on clean and impurity-1129 adsorbed diamagnetic surfaces. The experimental studies on these surfaces 1130 are reviewed in this section. As compared to diamagnetic surfaces, fewer 1131 papers have been published on well-defined magnetic surfaces prepared in 1132 UHV with an exception of $Cr_2O_3(0001)$ [227], although a number of stud-1133 ies by using traditional methods were reported on the o-p conversion on 1134 magnetic materials until 60's, which are comprehensively reviewed by Ilisca 1135 [10]. One reason for this might be that the conversion mechanism was be-1136 lieved to be understood by the Wigner model. Nevertheless, the surface 1137 magnetic structure of magnetic materials has only been clarified by the re-1138 cently developed surface-sensitive probes such as spin-polarized photoemis-1139 sion, spin-polarized STM, magnetic circular dichroism, spin-polarized LEEM 1140 and grazing-incidence nuclear resonant scattering. Therefore, the o-p con-1141 version on various spin structures will be an interesting subject in the future. 1142

¹¹⁴³ 5.1. Diamagnetic metal surface

1144 5.1.1. Clean Cu surface

And ersson and Harris performed EELS experiments for H_2 , D_2 and HD 1145 physisorbed on a single-crystal Cu(100) surface, and succeeded in observing 1146 the rotational and ro-vibrational excitation of molecular hydrogen [18]. The 1147 Cu(100) surface was cleaned by surface-science-based techniques in UHV, and 1148 exposed to normal H_2 , D_2 or HD at 12 K. Figure 29 shows the EELS spectra 1149 taken for H_2 , HD and D_2 at several experimental conditions. Taking account 1150 of the EELS cross section, the adsorbed hydrogen was considered to be in 1151 a monolayer regime. In the EELS spectra, in addition to the substrate Cu 1152 phonon at 28 meV (later assigned to contamination of water), loss features 1153 were observed at 45 and 72 meV for H_2 , 22 meV and 37 meV for D_2 , and 33 1154 meV for HD. The features for H₂ were assigned to the transitions of $J=0 \rightarrow$ 1155



Figure 29: EELS spectra for H_2 , HD and D_2 physisorbed on Cu(100) at low temperature. In addition to the elastic peak at a loss energy of 0 eV, loss features due to rotational excitation are observed at 45 meV ($J=0\rightarrow 2$) and 72 meV ($J=1\rightarrow 3$) for H_2 , 22 meV ($J=0\rightarrow 2$) and 37 meV ($J=1\rightarrow 3$) for D_2 , and 33 meV ($J=0\rightarrow 2$) for HD. Reproduced by permission from [18].

2 of p-H₂ and $J=1 \rightarrow 3$ of o-H₂. Furthermore, vibrational excitations of v=01156 \rightarrow 1 were recognized at around 500 meV for H₂ and 350 meV for D₂ as shown 1157 in Fig. 30. In addition to the pure vibrational excitation at 518 meV for 1158 H_2 and 372 meV for D_2 , combined excitations of vibration and rotation were 1159 also observed in the spectra. It is noted that the ortho and para species were 1160 not resolved in the pure vibrational excitation peak because the difference of 1161 the vibrational excitation energy between the ortho and para H_2 is smaller 1162 than the EELS resolution (Fig. 23(b) and Table 5). 1163

The rotational excitation energies observed by EELS are close to the gas phase values (Table 5). If the potential is anisotropic, the rotational level degeneracy is lifted as discussed in Fig. 9, which is reflected in the EELS spectrum. Svensson and Andersson analyzed the spectral shape of the J=0 $\rightarrow 2$ feature in detail to decompose it into two components as shown in



7L D2

12 K

Figure 30: EELS spectra for H₂ and D₂ physisorbed on Cu(100) in the loss energy range for vibrational excitation. Loss features are observed at 518 $(v=0\rightarrow1)$, 560 $((v=0\rightarrow1)+(J=0\rightarrow2))$, and 587 meV $((v=0\rightarrow1)+(J=1\rightarrow3))$ for H₂, and 372 $(v=0\rightarrow1)$, 393 $((v=0\rightarrow1)+(J=0\rightarrow2))$, and 407 meV $((v=0\rightarrow1)+(J=1\rightarrow3))$ for D₂. Reproduced by permission from [18].



Figure 31: Analysis of the EELS spectra for H₂ and HD physisorbed on Cu(100). The loss peaks at 45 and 33 meV ($J=0\rightarrow 2$) are decomposed into two components corresponding to the rotational sublevels of $M = \pm 2$ and M = 0 in J = 2. Reproduced by permission from [73].

Fig. 31, which were ascribed to excitation to the (J=2, M=0) and (J=2, M=1) and (J=2, M=1) states [73], the latter being more stable. From the obtained values, the anisotropic potential was evaluated to be +1.4 meV.

The EELS results definitely showed that both ortho and para species 1172 are present on the surface. The observed intensity ratio of the loss features 1173 corresponding to ortho and para H₂ obviously deviates from the gas-phase 1174 degeneracy of 3. This was initially attributed to the J-dependent sticking 1175 probability [99]. Nevertheless, the J-dependence of the sticking probability 1176 was later found to be small by a molecular-beam resonant sticking study 1177 [100, 53] and theoretical studies [104, 103]. The other point that should be 1178 clarified is the EELS cross section. The EELS cross section for the ortho and 1179 para H_2 was examined in detail by changing the relative coverage of ortho 1180 and para hydrogen on Cu(001) [73]. The cross section difference was then 1181

¹¹⁸² not sufficient to explain the deviation of the J=1/J=0 intensity ratio from ¹¹⁸³ 3 observed in EELS. Therefore, it was concluded that o-p conversion takes ¹¹⁸⁴ place on the Cu(001) surface.

In the first publication by Andersson's group, it was reported that the intensity ratio corresponding to J=1 and J=0 does not change on the experimental timescale suggesting that the o-p conversion rate is smaller than 1 %/min. In a later publication, on the other hand, the conversion time scale is reported to be 1 ML in 5 min [73]. They suspected that some unidentified non-controllable active sites for o-p conversion are present on the surface.

In recent work, Svensson and Andersson investigated o-p conversion of H₂ 1191 on a stepped Cu(510) surface and compared with a flat surface of Cu(100)1192 [228]. Figure 32 shows the EELS spectra taken for H_2 adsorbed on Cu(510) 1193 at different temperatures and H_2 background pressures. On this stepped 1194 surface, two adsorption states are clearly present as identified by the EELS 1195 features at 31 and 46 meV in Fig. 32(a). These loss features were attributed 1196 to the $J=0 \rightarrow 2$ transition of p-H₂ in a strongly hindered-rotor state at a 1197 step site and nearly-free-rotor state at a flat terrace [59, 60]. Excitations 1198 corresponding to $o-H_2$ in the two rotor states are also observed at 61 and 79 1199 meV in Fig. 32(b). Although they are designated 2D and 3D rotor states in 1200 the paper, respectively, we call them hindered-rotor and nearly-free-rotor in 1201 this article as mentioned in Sec. 2.2.1. 1202

As seen in Fig. 32(b), the loss intensity of $o-H_2$ was enhanced as the 1203 temperature and H_2 pressure were increased. By assuming that adsorption 1204 from gas phase, desorption from the specific site of the surface and o-p con-1205 version were in equilibrium, the o-p conversion time of H_2 at the step site 1206 was estimated to be as short as 1 s. Compared with a flat terrace, H_2 at a 1207 step site is more strongly bound at a shorter distance with a large potential 1208 anisotropy. The reason for the fast conversion was discussed that the short 1209 H_2 -Cu distance enhances the interaction between H_2 and substrate electrons. 1210 They also extended their analysis to H_2 on a flat terrace and argued that the 1211 conversion mainly takes place at a special active site, which is likely to be 1212 step atoms. 1213

Since Cu is a diamagnetic metal, no apparent magnetic moments are present on the surface except for nuclear spins. Since the magnetic moment due to nuclear spins is small, the conversion time due to nuclear spins is expected to be $\sim 10^5$ s. Yucel considered fluctuation of the electron spin density in the surface, and theoretically calculated the o-p conversion time of H₂ on Cu within a first-order perturbation on the basis of the magnetic

dipole, Fermi contact and electron orbital motion, which is shown in Fig. 18 1220 [21]. The calculated conversion time is longer than 10 h at an adsorption 1221 distance of 3 A[75, 173], which does not seem to be consistent with the 1222 experimental results. On the other hand, the conversion time calculated on 1223 the basis of the two-step XY and UY processes is 100 - 1000 s at a distance 1224 of ~ 2.5 Å(Fig. 20), which might explain the experimental results. The 1225 conversion time of 1 s at step sites is, however, too short as the XY and UY 1226 processes, and further discussion is necessary for elucidation of the conversion 1227 mechanism on Cu. 1228

1229 5.1.2. Clean Ag surface

The EELS study for H₂ and D₂ physisorbed on Ag surfaces was performed 1230 by Avouris et al. [17]. Two samples, single-crystal Ag(111) and an *in-situ* 1231 evaporated Ag film, were prepared in UHV, and EELS spectra were taken 1232 after various dosages of H_2 and D_2 at ~10 K, which are shown in Fig. 33. 1233 Three distinct energy loss features were observed at 49, 518 and 562 meV 1234 after H₂ dosage of 1 L, which were assigned to the rotational excitation of J=01235 \rightarrow 2, pure vibrational excitation of $v=0 \rightarrow 1$, and rovibrational excitation 1236 of $(v=0 \rightarrow 1)+(J=0 \rightarrow 2)$, respectively. With increasing H₂ dosage, these 1237 peaks were found to slightly shift to lower energies, and another loss feature 1238 was observed at 70 meV, which was attributed to the $J=1 \rightarrow 3$ rotational 1239 excitation of $o-H_2$. These values are close to those of the gas phase H_2 (Table 1240 5). It was also shown that the spectra taken for the Ag film and Ag(111) 124 are similar. As for D_2 , energy loss features were observed at 25, 370, 395, 1242 and 410 meV. These values are similar to the excitation energies of $J=0 \rightarrow$ 1243 2, $v=0 \rightarrow 1$, $(v=0 \rightarrow 1)+(J=0 \rightarrow 2)$, and $(v=0 \rightarrow 1)+(J=1 \rightarrow 3)$ of D₂, 1244 respectively. On this basis, it was argued that both H_2 and D_2 adsorbed 1245 on Ag surfaces are in the nearly-free rotor state. On the other hand, the 1246 loss feature observed in the EELS spectra was broader than 10 meV. It was 1247 argued that this broadening is caused by rotational sublevel splitting [229], 1248 which suggests potential anisotropy as shown in Fig. 9. The broadening, 1249 however, might be caused by combined excitation of rotational states with 1250 the surface-molecule vibration $(n=0 \rightarrow 1 \text{ in Fig. } 9(a) \text{ and Fig. } 10)$. 1251

In the work by Avouris et al., the sticking probability of molecular hydrogen was considered to be unity, and a condensed multilayer was considered to be formed at a larger H₂ dosage. However, the sticking probability of molecular hydrogen is generally low as discussed in Sec. 2.2.3. Considering the vapor pressure of H₂ at 10 K [230, 231], furthermore, multilayer forma-



Figure 32: EELS spectra for H₂ physisorbed on Cu(510) taken at temperatures and H₂ pressures of (a) 14 K and 1.5×10^{-8} Torr, (b) 16 K and 1.5×10^{-7} Torr, (c) 17 K and 5×10^{-7} Torr, and (d) 14 K and 4×10^{-7} Torr. Two adsorption states are indicated by 2D and 3D, which correspond to the hindered rotor and nearly-free-rotor physisorbed at step and flat terrace sites, respectively. The rotational transitions relevant to the loss features are also shown in the figure. The relative intensities of the J = 0 and J = 1 states depend on the pressure and temperature reflecting the gas-phase exchange and o-p conversion kinetics. From the experimental data, the o-p conversion time is analyzed to be 1 s at step sites. Reproduced with permission from [228].

tion is unlikely to occur at 10 K in the UHV condition. Therefore, the EELS
spectra correspond to those in the monolayer region. On the basis of the
rotational excitation energy, the possibility of the contraction of the interatomic distance was discussed. Recent theoretical studies, however, showed
that contraction of the interatomic distance is small in the physisorbed state
[75, 173].

Similarly to Cu, both J = 0 and J = 1 corresponding to para and ortho species are present on Ag surfaces. Nevertheless, the intensity of $J=1 \rightarrow 3$ is considerably smaller than that of $J=0 \rightarrow 2$, indicating that the adsorbed H₂ is mostly in p-H₂. Avouris et al. suggested that o-p conversion takes place within 1–2 min on the surface or upon adsorption. Since the o-H₂ intensity is only recognized at a larger H₂ dosage, the o-p conversion time might be coverage-dependent [17].

The o-p conversion was later investigated by the REMPI-PSD method 1270 [25, 223, 224, 225]. Normal H₂ and D₂ were adsorbed on an Ag film surface 1271 prepared in UHV at 7 K, and the evolution of the ortho and para coverages 1272 were monitored by REMPI-PSD. Figure 34(a) shows the intensity ratio of 1273 H_2 in J = 1 and J = 0 as a function of time after H_2 dosage. This reflects 1274 the coverage loss of $o-H_2$ due to o-p conversion [25]. The intensity ratio 1275 decreased almost exponentially with increasing time, which suggests that o-1276 p conversion takes place on the surface as a first-order reaction. From the 1277 analysis of the data, the o-p conversion time of H_2 and p-o conversion time 1278 of D_2 on Ag were evaluated to be ~700 and ~1000 s, respectively [25, 224]. 1279

It was furthermore recognized that there are photoexcitation effects on 1280 the o-p conversion, which is displayed in Fig. 34(b) and (c). Here, the o-p 1281 ratio was monitored at a higher PSD laser fluence. As the PSD laser fluence 1282 of 193 nm (6.4 eV) was increased, the (J = 1)/(J = 0) ratio decreased 1283 faster. This suggests that photon irradiation accelerates the o-p conversion. 1284 Wavelength dependence of this photoexcitation effect is demonstrated in Fig. 1285 35. While irradiation of photons at 6.4 eV accelerated the o-p conversion of 1286 H_2 on Ag (Fig. 35(b)), no significant effects were observed with irradiation 1287 of 2.3 eV photons (Fig. 35(c)) [224]. 1288

The surface of deposited Ag films contains defects such as vacancies and steps, which might affect the o-p conversion kinetics. As described above, however, the spectral features and intensities of EELS representing o-H₂ and p-H₂ were similar between on a single-crystal surface of Ag(111) and an Ag film surface [17]. The o-p conversion time measured for the Ag(111) surface with REMPI-TPD also showed the conversion time of about 700 s in



Figure 33: EELS spectra for H₂ physisorbed on Ag surfaces at ~10 K after various H₂ dosages. Loss features develop at 49 ($J=0\rightarrow 2$), 70 ($J=1\rightarrow 3$), 518 ($v=0\rightarrow 1$), and 562 meV (($v=0\rightarrow 1$)+($J=0\rightarrow 2$)) with increasing H₂ dosage. The loss intensity corresponding to J=1 is apparently smaller than that of J=0. Reproduced with permission from [17].



Figure 34: The (J=1)/(J=0) ratio of H₂ on an Ag surface at 7 K as a function of time probed by REMPI-PSD with a pump laser wavelength of 193 nm at laser fluences of (a) 10, (b) 150, and (c) 250 μ J /cm² / pulse. The (J=1)/(J=0) ratio decreases exponentially due to the o-p conversion, and the solid curves are fits with an exponential function. Filled squares denote the (J=1)/(J=0) ratios measured 600, 1200, and 2400 s after the H₂ dosage on a nonirradiated sample to avoid the laser irradiation effect. Reproduced with permission from [25].



Figure 35: The (J=1)/(J=0) ratio of H₂ on an Ag surface at 6 K as a function of time probed by REMPI-PSD. (a) Probed every 250 s, (b) probed under laser irradiation at 193 nm with a fluence of 120 μ J/cm²/pulse and a repetition rate of 10 Hz, (c) probed under laser irradiation at 532 nm with a fluence of 120 μ J/cm²/pulse and a repetition rate of 10 Hz. While the laser irradiation of 193 nm accelerates the o-p conversion, that of 532 nm has little effect on the o-p conversion. Reproduced with permission from [224].



Figure 36: Electronic structure of H₂ physisorbed on Ag, and electronic processes for the two-step UY process. An electron in Ag is virtually excited to the $1\sigma_u$ orbital through the Coulomb interaction (C) forming H₂⁻. The electron is transferred back to Ag through the Fermi contact interaction (Y) along with the nuclear-spin conversion.

agreement with the value obtained for the deposited film [96]. These results suggest that the defect sites formed on an Ag film surface do not play a major role in the o-p conversion on Ag surfaces, which is also supported by the experiment on Xe-covered Ag(111) as mentioned in the next section. This is in contrast to the fact that the step site is active for the conversion on Cu surfaces [228].

On the basis of the experimental data, the o-p conversion mechanism 1301 was discussed in terms of the two-step process (Sec. 3.4). Since Ag is an 1302 sp-metal with a diamagnetic character as Cu, where the d-band is fully filled 1303 with the Fermi level located within the sp-band, the o-p conversion time 1304 can be compared with the values calculated for Cu on the basis of the one-1305 step process shown in Fig. 18 [21]. The values calculated on the basis 1306 of the interaction with substrate electrons are considerably larger than the 1307 experimentally obtained conversion time of ~ 700 s by REMPI-PSD. On the 1308 other hand, the conversion times predicted by the two-step model are close 1309 to the experimental values as shown in Fig. 20 [162, 22]. 1310

The excitation energy for the intermediate state in the XY process (Fig. 19) is ~10.5 eV considering the lowest neutral ${}^{3}\Sigma_{u}^{+}$ state of H₂ is ${}^{3}\Sigma_{u}^{+}$ [232]. The excitation energy for the UY process, on the other hand, was estimated to be ~6 eV, because the Fermi level (E_f) of Ag is located at ~4.6 eV below the vacuum level [233] and the affinity level of H₂ is lowered from 2.3 eV ¹³¹⁶ above the vacuum level in gas phase [194] by about 1 eV due to the image ¹³¹⁷ charge effect [234, 10], which is shown in Fig. 36.

Of these two processes, the two-step UY mechanism was furthermore 1318 rationalized by the results of photoacceleration. The experimental results 1319 showed that the o-p conversion is accelerated by 6.4 eV photons, but not at 1320 2.3 eV. As discussed above, the intermediate state for the UY process (the H_2^- 1321 state) can be formed at a photon energy of 6.4 eV, while it is not accessible 1322 by photons at 2.3 eV. The neutral ${}^{3}\Sigma_{u}^{+}$ state in the XY process, on the other 1323 hand, can be excited by photons at neither 6.4 eV nor 2.3 eV. Therefore, the 1324 intermediate state responsible for the o-p conversion was argued to be the 1325 ionic H_2^- state. 1326

The isotope difference between H_2 and D_2 was also discussed. The con-1327 version time of D_2 is longer than that of H_2 by a factor of 1.7 [224]. The elec-1328 tronic structure is essentially the same for these two isotopes, which means 1329 the electron transfer process in Fig. 36 is expected to be similar for H_2 and 1330 D_2 . On the other hand, the Fermi contact (Y) process for H_2 is more effi-1331 cient than that of D_2 by a factor of 7.1, as described in Sec. 3.6, which is 1332 significantly larger than the experimental result of 1.7. This apparent dis-1333 crepancy is discussed to be ascribed to the efficiency of the rotational-energy 1334 dissipation. The rotational energy of 7.1 meV for D_2 can be more efficiently 1335 accommodated by other degrees of freedom as compared to 14.7 meV for H₂ 1336 [224].1337

An important discussion was given on the energy dissipation mechanism. 1338 Whereas the o-p conversion time was derived by the REMPI-PSD method, 1339 the results showed no increase of the $p-H_2$ on the surface even after o-p 1340 conversion [25]. This suggested that the converted H_2 was desorbed from 134 the surface upon o-p conversion. As shown in the previous section, the o-p 1342 conversion is accompanied by the rotational-energy release. A possible inter-1343 pretation was that the rotational energy was transferred to the motion of the 1344 molecule center of mass, either the H₂-substrate vibration or H₂ translational 1345 motion in the surface parallel direction, and that desorption of the molecule 1346 was subsequently promoted. It is noticed, however, the rotational energy of 134 $o-H_2$ in J=1 is smaller than the typical physisorption energy, which means 1348 the rotational-energy transfer is not sufficient to induce desorption of H_2 . 1349

For such energy transfer, coupling between the rotational and translational (vibrational) motion is necessary, which originates from the potential anisotropy of physisorption. A recent theoretical study showed that the rotational sublevel splitting due to the anisotropy imposes steric effects and accelerated o-p H_2 conversion [173, 172]. In the original two-step model, on the other hand, the rotational energy is expected to flow into the electronic system of the substrate causing electronic excitation. The mechanism of the o-p energy release is obviously the subject of further experimental and theoretical studies.

1359 5.1.3. Impurity-adsorbed surface

1360 Xe

Sakurai et al. investigated H_2 adsorbed on Xe-covered Ag(111) with 1361 EELS [20]. The substrate was kept at 6 K and exposed to a certain pressure 1362 of H_2 while measuring the EELS spectrum. Thus, the adsorbed H_2 layer 1363 at the surface was expected to be in equilibrium with the gas phase H_2 , 1364 where the desorption rate is balanced with adsorption rate from the gas 1365 The EELS spectra revealed energy loss features corresponding to phase. 1366 rotational excitations and rovibrational excitations of H₂. From the analysis 136 of the experimental data on clean Ag(111), the o-p ratio was estimated to 1368 be 0.2 - 0.4 at a low H₂ pressure and increase to 1.3 - 1.5 at a higher 1369 H_2 pressure. On the basis of this observation, it was concluded that o-p 1370 conversion takes place on Ag(111). Since the o-p conversion was suspected 1371 to occur on local active sites such as defects, the o-p ratio was examined 1372 on Xe-covered Ag(111) where Xe atoms were expected to occupy the active 1373 sites prior to H_2 adsorption. However, the o-p ratio did not change with and 1374 without the Xe preadsorption, which would rule against the o-p conversion 1375 mechanism originating from defect sites. The results suggest that the o-p 1376 conversion on Ag(111) takes place via the direct interaction with the flat 1377 substrate Ag. 1378

1379 \mathbf{N}_2

The rotational spectrum was observed with EELS for H_2 co-adsorbed 1380 with N_2 on Ag(111) at a temperature of 20 K by Gruyters and Jacobi [235]. 1381 While H_2 was not adsorbed at 20 K on bare Ag(111), an appreciable amount 1382 of H_2 was adsorbed on the N_2 -precovered surface. The loss peaks due to 1383 rotational excitations of 44 and 73 meV were observed in the spectra for H_2 1384 of 2 L adsorbed on N₂-precovered Ag(111). The loss peak intensity at 73 1385 meV corresponding to J=1 of o-H₂ was about a quarter of the J=0 intensity 1386 of $p-H_2$, which suggested o-p conversion on Ag(111). Although the precise 1.38 conversion time was not given in the work, the o-p conversion seems to have 1388

¹³³⁹ proceeded with the time scale comparable with the EELS acquisition time. It
¹³⁹⁰ is also suggested that the o-p conversion takes place via the direct interaction
¹³⁹¹ with the substrate Ag as discussed for Xe-covered Ag.

1392 O_2

The electronic term of the ground state O_2 is ${}^{3}\Sigma_{q}^{-}$, which means that 1393 O_2 is paramagnetic with an electron spin of 1 in contrast to Xe and N_2 . 1394 When O_2 is chemisorbed on a surface, possible electron transfer between 1395 O_2 and the substrate changes the magnetic property of O_2 . When O_2 is 1396 in the physisorption state, however, little electron transfer occurs, and the 1397 electronic spin state is expected to be unchanged. Such an electronic state 1398 of O₂ on metallic surfaces was identified on Cu and Pt by Near-edge X-ray 1399 absorption fine structure (NEXAFS) spectroscopy [236, 237]. Although the 1400 electronic spin of O_2 is retained on metal surfaces, the Kondo effect could 1401 screen the spin at low temperature [238]. On Ag surfaces, O_2 is physisorbed at 1402 low temperature [239], and the adsorption structure and magnetic property 1403 of O_2 were studied experimentally and theoretically [240, 241, 242]. Even 1404 in the chemisorbed state, dissociated O might have a magnetic moment as 1405 pointed out by a recent theoretical study [243]. 1406

Sakurai et al. investigated the effect of O_2 on the o-p ratio of H_2 physisorbed on Ag(111) in equilibrium with gas phase by EELS [20]. The o-p ratio of H_2 measured on $O_2/Ag(111)$ was smaller than that on clean Ag(111), which was attributed to the magnetic effect of adsorbed O_2 .

The effect of adsorbed O_2 on the o-p conversion of H_2 and p-o conver-1411 sion of D_2 was investigated on Ag surfaces at 7 K by REMPI-PSD [225]. 1412 Figure 37 shows the change of the (J=1)/(J=0) ratio as a function of time 1413 on O_2 -adsorbed Ag surfaces with various O_2 coverages. As the O_2 coverage 1414 was increased, the decay rate of the (J=1)/(J=0) ratio was found to be en-1415 hanced. A similar tendency was also observed for D_2 . Adsorbed O_2 obviously 1416 promoted the H_2 o-p conversion and D_2 p-o conversion. Figure 38 shows the 1417 variations of the decay time constants for H_2 and D_2 as a function of the O_2 1418 coverage. The time constants decreased with increasing O_2 coverage. 1419

Because the O₂ coverage was in a submonolayer regime in the experi-¹⁴²¹ments, the o-p (p-o) conversion of adsorbed H₂ (D₂) was postulated to take ¹⁴²²place via two independent mechanisms: one near the paramagnetic O₂ and ¹⁴²³the other on the bare Ag surface. The conversion mechanism on Ag is, as ¹⁴²⁴described in Sec. 5.1.2, likely to be the two-step UY process with conversion ¹⁴²⁵times of ~700 and ~1000 s for H₂ and D₂, respectively. In the vicinity of



Figure 37: Time evolution of the (J=1)/(J=0) ratio of H₂ physisorbed on O₂-adsorbed Ag surfaces probed by REMPI-PSD at 7 K. With increasing O₂ coverage from 0 (open circles) to 0.01 ML (open triangles), the (J=1)/(J=0) ratio decreases faster, which indicates the o-p conversion is accelerated by adsorbed O₂. Reproduced with permission from [225].

the adsorbed O_2 , on the other hand, the o-p conversion can be promoted by 1426 the O_2 paramagnetic spin. When such two mechanisms are operative, the 142 entire conversion kinetics is governed by the diffusion of hydrogen molecules 1428 on the surface as was also discussed for o-p conversion in solid H_2 contain-1429 ing O_2 impurities [244, 245, 246]. The diffusion might proceed via quantum 1430 tunneling as well as thermal activation [247]. With the aid of Monte Carlo 1431 simulations, the hydrogen diffusion time between the adsorbed O_2 and bare 1432 Ag sites was found to be much faster than the o-p conversion time at either 1433 bare Ag site and in the vicinity of O_2 on the basis of the single-exponential 1434 feature observed by the experiments in Fig. 37. Figure 38 demonstrates the 1435 change of the relative contribution of the two conversion mechanisms. From 1436 the analysis of the data, the conversion times in the vicinity of adsorbed O_2 1437 was estimated to be 8.3 ± 1.2 s and 53.4 ± 9.3 s for H₂ and D₂, respectively 1438 [225].1439

As discussed in the beginning of this section, adsorbed O_2 is expected to



Figure 38: Decay time of the (J=1)/(J=0) ratio (nuclear-spin conversion time) of (a) H₂ and (b) D₂ on O₂-adsorbed Ag surfaces at 7 K estimated by REMPI-PSD as a function of the O₂ coverage. The o-p (p-o) conversion of H₂ (D₂) is promoted by adsorbed O₂. Reproduced with permission from [225].

possess paramagnetic electron spins. In the vicinity of O_2 , there exists an 1441 inhomogeneous magnetic field as illustrated in Fig. 15. The magnetic dipole 1442 interaction in Eq. 15 then yields the nuclear-spin and position difference 1443 operators, which induce o-p conversion as a one-step process. As well as 1444 the magnetic dipole interaction, the direct Fermi contact interaction might 1445 also be important for the o-p conversion, because adsorbed O_2 has a finite 1446 electron density at the H_2 position. The o-p conversion time due to the 1447 magnetic dipole and Fermi contact interactions was theoretically evaluated 1448 by a recent study by Kunisada et al. [248]. Figure 39 shows the o-p conversion 1449 time calculated as a function of the H_2 distance from the Ag surface at 1450 two adsorption sites of O_2 on-top and beside O_2 . The conversion time is 1451 strongly dependent on the H_2 -Ag distance at both adsorption sites, and the 1452 conversion time at an distance of 3.5 Å is about 10 s at the on-top site of 1453 O_2 for both interactions. This is consistent with the experimentally observed 1454 value, suggesting that the conversion mechanism in the vicinity of O_2 is the 1455 one-step process via the magnetic dipole and/or Fermi contact interactions. 1456 It should be noted, however, that the contribution of the two-step XY process 1457 via the Coulomb and Fermi contact interactions (Sec. 3.4) could be operative 1458 competing with the one-step process as pointed out by Ilisca and Sugano 1459 [160].1460

Compared with the conversion through direct interaction with Ag (Sec. 1461 5.1.2), the isotope dependence of the conversion time in the vicinity of ad-1462 sorbed O_2 is larger with a relative ratio of 6.4. As described in Sec. 3.6, the 1463 magnetic interaction, either magnetic dipole or Fermi contact, is larger for 1464 H_2 than D_2 by a factor of 7.1, which is similar to the experimental result 1465 of 6.4. In contrast to the conversion on bare Ag, the overall conversion rate 1466 is mainly governed by the strength of the magnetic interaction and the effi-1467 ciency of the rotational-energy dissipation seems to be similar for H_2 and D_2 1468 |225|. 1469

1470 5.1.4. Graphite

¹⁴⁷¹Compared with metallic surfaces, the graphite surface is inert and is likely ¹⁴⁷²to be free from contaminations as demonstrated by STM atomic imaging ¹⁴⁷³in an atmospheric condition. Moreover, the graphite sample can be easily ¹⁴⁷⁴cleaved providing a large surface area. Well-defined graphite surfaces are, ¹⁴⁷⁵therefore, expected to be prepared in an ordinary vacuum apparatus without ¹⁴⁷⁶UHV, and traditional non-surface-sensitive methods may be used for the o-p ¹⁴⁷⁷conversion study. The o-p conversion data on charcoal surfaces obtained in


Figure 39: O-p conversion time calculated for H₂ on O₂-adsorbed Ag(111) as a function of the H₂ distance from Ag(111). O₂ is adsorbed at the bridge site of Ag with its molecular axis parallel to the surface, and the O atom is calculated to have a magnetic moment of 0.52 β_e (β_e : Bohr magneton). The o-p conversion time is evaluated on the basis of the one-step process with the magnetic dipole and Fermi contact interactions at (a) on-top of O₂ and (b) near O₂ sites. At the on-top O₂ site, the conversion time is about 10 s. Reproduced with permission from [248].

¹⁴⁷⁸ a traditional manner are available in literatures [1], which may be compared ¹⁴⁷⁹ with the studies described below.

Kubik et al. investigated the H_2 layer physisorbed on graphite in order 1480 to examine the rotational ordering [249]. They prepared a stack of Grafoil 1481 sheets sandwiched with Teflon and Cu foils to ensure cooling of the sample, 1482 on which enriched $o-H_2$ and $p-D_2$ with a concentration of c>0.98 were ad-1483 sorbed, and the relative ortho concentration on the sample was measured by 1484 NMR below 4.2 K. The absolute concentration $o-H_2$ and $p-D_2$ before adsorp-1485 tion was determined with Raman by probing the $J=1 \rightarrow 3$ and $J=2 \rightarrow 4$ 1486 transitions. The o-p conversion of H_2 was found to be a first-order process 1487 and the conversion rate was as slow as 0.4 %/h, although the raw data are not 1488 presented in the paper. The D_2 conversion was much slower than that of H_2 , 1489 and the conversion rate was estimated to be $0.069 \ \%/h$ on the assumption 1490 that the conversion is the first-order process. It is also reported that the o-p 1491 conversion during the adsorption process is much faster than these conversion 1492 rates. The authors suspect that there are some paramagnetic sites, which 1493 can be active for o-p conversion when H_2 is mobile on the surface during 1494 adsorption. According to the phase diagram of H_2 on graphite, a fluid phase 1495 exists above 10 - 20 K, where H₂ is mobile on the surface [250, 251]. 1496



Figure 40: EELS spectrum for H₂ physisorbed on highly ordered pyrolitic graphite at 10 K. Loss features are observed at 47 ($J=0\rightarrow 2$), 513 ($v=0\rightarrow 1$), and 565 (($v=0\rightarrow 1$)+($J=0\rightarrow 2$)). No features originating from J = 1 (o-H₂) are observed. Reproduced by permission from [19].

Palmer and Willis investigated H_2 on highly ordered pyrolitic graphite 1497 (HOPG) at 10 K by EELS in UHV, which is shown in Fig. 40. Prior to the 1498 measurements, the sample was heated to 1300 K to remove contaminants. 1499 After adsorption of n-H₂, a loss feature was observed at 47 meV correspond-1500 ing to $J=0 \rightarrow 2$ of p-H₂ and no loss feature corresponding to o-H₂ was 1501 identified. The shoulder at 80-100 meV was assigned to double excitation 1502 of $J=0 \rightarrow 2$, and the loss features at 513 and 555 meV were attributed to 1503 the pure vibrational excitation of $v=0 \rightarrow 1$ and ro-vibrational excitation of 1504 $(v=0 \rightarrow 1) + (J=0 \rightarrow 2)$. On the basis of the data, it was concluded that 1505 H_2 is in a nearly-free-rotor state and that the o-p conversion occurs on the 1506 graphite surface on a time scale of 1 min. The authors argued that there are 1507 paramagnetic spins on the surface and the mechanism is the one-step process 1508 via the magnetic dipole interaction as the O_2 impurity on Ag. 1509

Yucel et al. investigated p-o conversion of D_2 on Grafoil at various temperatures by the thermal-conductivity method [252]. The grafoil disks were

sandwiched in copper foils to efficiently cool the sample. The $p-D_2$ concentra-1512 tion decreased exponentially against time indicating that the p-o conversion 1513 proceeds as a first-order reaction. The measured conversion time was almost 1514 constant up to a certain coverage followed by a linear increase with increas-1515 ing D_2 coverage. This coverage was regarded as completion of a monolayer. 1516 In a monolayer regime, D_2 forms either a dense incommensurate layer or a 1517 relatively nondense commensurate layer. The temperature dependence of the 1518 conversion time was measured for these two phases, which are shown in Fig. 1519 41. In the temperature range of 10 - 20 K, the conversion time was almost 1520 constant, and they increased below 10 K. No significant conversion was ob-1521 served at 6.4 K. They argued that there are paramagnetic sites of less than 1522 0.01 and the increase of the conversion time is due to limited diffusion of D_2 1523 on the surface. Furthermore, the conversion time was theoretically estimated 1524 in terms of the energy dissipation into two-dimensional solid phonon and lat-1525 eral vibration in the two-dimensional lattice gas, which were in a reasonable 1526 agreement with the experimental results. No unusual conversion during the 1527 adsorption process was observed in this work [252]. 1528

The three studies on graphite shown above all suggested presence of para-1529 magnetic impurity sites on the surface promoting o-p conversion of H₂ or p-o 1530 conversion of D₂. Compared with metal surfaces, a graphite surface possi-1531 bly has an edge site. The edge state of graphite was theoretically shown to 1532 possess a localized magnetic moment [253], which is being confirmed experi-1533 mentally [254]. Such edge states could produce an inhomogeneous magnetic 1534 field and promote o-p conversion as O_2 on Ag. The o-p conversion data might 1535 provide information about the magnetism of the edge state of graphite. On 1536 the other hand, graphite is a zero-gap semiconductor, and there is a finite 153 density of states near the Fermi level within the o-p energy [255]. This in-1538 dicates that the two-step XY or UY mechanism may be operative as Cu 1539 and Ag. Nevertheless, the slow conversion rate when H_2 is solidified on the 1540 surface at lower temperatures suggests that the two-step mechanism is not 1541 effective possibly because of the low density of states near the Fermi level. 1542 It is noted that the electron density near the Fermi level of the graphene 1543 (single layer of graphite) can be controlled by the gate bias as demonstrated 1544 by the electrical conductivity measurement shown in Fig. 42 [255]. The band 1545 structure of graphene is described by the Dirac cone and the electron density 1546 at the Fermi level increases with increasing and decreasing bias with respect 1547 to the Dirac point. This suggests the two-step conversion process can be 1548 enhanced by applying the gate bias enabling us to electrically control the o-p 1549



Figure 41: The p-o conversion time of D_2 on Grafoil measured by the thermal-conductivity method as a function of sample temperature. (a) High H_2 coverage incommensurate phase (open and solid symbols: first and second series) and (b) low H_2 coverage commensurate phase (circle and triangle symbols: low and high coverages). Reproduced by permission from [252].



Figure 42: Electric conductivity of a graphene as a function of the gate bias voltage. Insets show the Dirac cone band structure and change of the Fermi level at different bias voltages. Reproduced by permission from [255].

1550 conversion.

1551 5.2. Diamagnetic insulator surface

1552 5.2.1. Ice surfaces

The o-p conversion proceeds on diamagnetic metal surfaces via the two-1553 step process accompanied by virtual electronic excitation. On diamagnetic 1554 insulator surfaces, on the other hand, the two-step process seems not oper-1555 ative. As illustrated in Fig. 19(b) and (c), the final state of the surface in 1556 the two-step process is a spin triplet state. Whereas a diamagnetic metal 1557 requires no excitation energy for the transition from the electron spin singlet 1558 to triplet, a large energy corresponding to the band gap is necessary for the 1559 transition on a diamagnetic insulator. O-p conversion was investigated on 1560 some oxides and a D_2O -covered oxide in a traditional manner, where the 1561 conversion was clearly recognized [256]. While the fast conversion on oxide 1562 surfaces was attributed to paramagnetic impurities due to oxygen vacancies, 1563 the origin of the conversion on a D_2O surface was discussed to originate from 1564 an induced dipole moment, although detailed physical mechanism was not 1565 given [256]. 1566

As stated in Sec. 6, amorphous solid water (ASW) is a dominant inter-1567 stellar medium, and a diamagnetic insulator. H_2 on ASW was investigated 1568 with IRAS by Hixson et al. [198]. Microporous amorphous ice with a thick-1569 ness of $2-4 \mu m$ was formed on a substrate at 12 K, and H₂ was admitted in 1570 the ASW sample either during the water deposition or by exposure of ASW 157 to H_2 gas after ASW formation. Figure 43 shows the IR spectra taken for 1572 H_2 adsorbed on amorphous D_2O at 12 K under a high vacuum environment. 1573 Two absorption features were observed at 4132 and 4141 $\rm cm^{-1}$. After care-1574 ful examination of the spectra, the two features were ascribed to vibrational 1575 excitation of o-H₂ and p-H₂, respectively, which are slightly shifted from the 1576 gas phase values of 4153 and 4159 $\rm cm^{-1}$ listed in Table 5. As stated in Sec. 1577 4, these signals are induced absorption due to the surface electric field. As 1578 the adsorption time was increased from 10 to 600 min, the absorption in-1579 tensity corresponding to $o-H_2$ decreased whereas that due to $p-H_2$ increased 1580 noticeably. This suggests o-p conversion on the ice surface, and the effective 1581 conversion half-life time was estimated to be ~ 40 min. It was also noticed 1582 that the absorption intensity due to $o-H_2$ was recovered upon annealing at 1583 16 K, which was ascribed to enhanced mobility of H_2 in ASW. It was argued 1584 that there are several binding sites for H_2 , some of which are active for in-1585 duced IR absorption [198]. Upon annealing, unconverted $o-H_2$ migrated to 1586 such active sites. Hixson et al. also theoretically examined the binding site 1587 of H_2 in the amorphous ice. There were several binding geometries identified 1588 and the estimated vibrational frequencies of H_2 were in agreement with the 1589 experimental observation. As to the kinetics of the o-p conversion in ASW, 1590 the purity of the sample was questioned in this work, and the conversion 1591 was suspected to be induced by some paramagnetic impurities incorporated 1592 during the ice sample preparation [198]. 1593

The o-p conversion on ASW was recently investigated by the REMPI 1594 technique. While the p-o conversion rate of D_2 was found to be slow and 1595 negligible [26], Watanabe et al. suggested o-p conversion of H_2 taking place 1596 on an ASW surface [257]. Sugimoto and Fukutani investigated the o-p con-1597 version of H_2 and p-o conversion of D_2 on ASW in a UHV condition. ASW 1598 films of sixty layers were grown on a clean Ag(111) single-crystal surface at 1599 13 K by dosing D_2O molecules. Prior to the conversion experiments, the 1600 ice film was annealed up to 55 K to avoid any structural changes during the 1601 conversion experiments and to confirm the sample is free from O_2 , which des-1602 orbs below 50 K [258]. After exposing the ASW film at 10 K to $n-H_2$ (D₂), 1603 the coverages of the ortho and para species were measured by REMPI-TPD. 1604



Figure 43: Infrared absorption spectra of H_2 on porous amorphous ice at 12 K as a function of time after H_2 adsorption from 10 to 600 min. The absorption features at 4132 and 4141 cm⁻¹ correspond to vibrational excitation of o-H₂ and p-H₂, respectively, and the absorption intensities change as a function of time. Reproduced by permission from [198].

Figure 44 shows the time evolution of the H₂ and D₂ in J=0 and J=1. As the 1605 residence time on the ASW surface increased, the intensity of J=1 decreased 1606 and that of J=0 increased. Fitting a single exponential function to the J=11607 data gave decay time constants of 370^{+340}_{-140} s for H₂ and 1220^{+2980}_{-580} s for D₂ 1608 (the numbers in the subscripts and superscripts denote the upper and lower 1609 limits of uncertainty). In contrast to the conversion on a flat Ag surface [25], 1610 the sum of the intensities of these two species was constant within the ex-1611 perimental accuracy. This is probably because the ASW sample was porous: 1612 even if the converted H_2 was desorbed from the surface of ASW pores, the 1613 H_2 was readsorbed on the pore surface. The conversion time of D_2 is also 1614 reported to be as slow as 15 %/h [226], which is in rough agreement with the 1615 above value. 1616

A possible mechanism for the o-p conversion of H_2 and p-o conversion 1617 of D_2 was discussed in detail [33]. On an ice surface, there exists a strong 1618 electric field originating from the electric dipole moment of H_2O (1.85 D). 1619 When water molecules are condensed in a solid, the dipole moment is possi-1620 bly enhanced to about 3 D [259, 260], which would also enhance the electric 1621 field on the surface of ASW. The electric field on an ice surface was simulated 1622 to be $\sim 10^{10}$ V/m [261, 262], which could produce an electric field gradient of 1623 10^{20-21} V/m². Since the deuteron nucleus has a quadrupole moment (Table 1624 4), the nuclear quadrupole interaction can be significant for the p-o con-1625 version of D_2 . Under an electric-field gradient, as discussed in Sec. 3, the 1626 spin-state transition and rotational-state transition can be induced as a one-1627 step process. With this electric field gradient, the quadrupole interaction 1628 energy leads to $\sim 10^{-11}$ eV. This value was compared with the theoretical 1629 estimation for solid D_2 [139]. In solid D_2 , there exists an electric field of 1630 $\sim 10^8$ V/m and field gradient of $\sim 10^{18}$ V/m² at a neighboring site due to the 1631 quadrupole moment of hydrogen molecules (0.48 au) [3]. With this electric 1632 field gradient, the p-o conversion time was theoretically evaluated to be $\sim 10^8$ 1633 s [139]. If the electric field gradient is 10^{20-21} V/m², the conversion time is 1634 expected to be $\sim 10^3$ s, which is in agreement with the experimental value on 1635 ASW. 1636

The effects of nuclear spins should be considered, as hydrogen (deuteron) of H_2O (D_2O) has a nuclear spin, which produces an inhomogeneous magnetic field causing the conversion as a one-step process. The conversion time due to the nuclear spins, however, was roughly evaluated to be the order of 10^6 and 10^8 s for H_2 and D_2 , respectively [149, 139]. Therefore, the magnetic interaction with nuclear spins of water is not significant for the H_2 o-p



Figure 44: Time evolution of the intensities of J=0 and J=1 and the sum of them measured by REMPI-TPD. (a)H₂ and (b) D₂ on the porous amorphous ice at 10K. The J=1 intensity decreases, while that of J=0 increases with the sum kept constant. Reproduced by permission from [33].

 $_{1643}$ conversion and D_2 p-o conversion on ice surfaces observed in Fig. 44.

Since proton has no quadrupole moment, on the other hand, the quadrupole 1644 interaction is not responsible for the o-p conversion of H_2 on ASW. As a pos-1645 sible conversion mechanism, a multi-step process has been proposed [33]. As 1646 mentioned in Sec. 3.3.2, different I states are intrinsically mixed with each 164 other through the intra-molecular Fermi contact interaction in the electron-1648 ically excited state. It should be noted that the mixed state is the electron 1649 spin triplet (S=1) state with the same rotational state. For the o-p con-1650 version, both the electron spin state and rotational states must be changed. 1651 When the spin-orbit coupling (SOC) is taken into consideration, the S=01652 state can be mixed with the S=1 state as shown in Fig. 21. 1653

Without external perturbation, the ortho and para states are not mixed 1654 with each other. Under the surface electric field, however, the Stark coupling 1655 mixes gerade and ungerade states leading to $\Delta J=1$. With these multi-step 1656 processes, the ortho and para states can be mixed. As described in Sec. 1657 3.3.2, SOC is considerably enhanced for the near-degenerate $r^3\Pi_a$ and $R^1\Pi_a$ 1658 states because of weak electron exchange-correlation. The intra-molecular 1659 Fermi contact strength was also shown to be enhanced when the vibrational 1660 level was taken into consideration for the $a^{3}\Sigma_{g}^{+}$ and $B^{1}\Sigma_{u}^{+}$ states: the energy 1661 levels of a (v=6) and B (v=8) are near-degenerate, which results in strong 1662 o-p coupling of these states [263]. On the basis of a possible electric field 1663 of as large as 10^{10-11} V/m on the ice surface, the overall conversion time 1664 of H_2 was roughly estimated to be 10^2 s, which were in agreement with the 1665 experimental results of REMPI-TPD. It was also pointed out that the SOC 1666 may be enhanced with an external electric field [33]. As discussed in Sec. 3.6, 1667 the magnetic interaction is weaker for D_2 because of the smaller nuclear g 1668 factor. It is discussed that the conversion time due to the multi-step process 1669 for D_2 is comparable or longer than the conversion time evaluated on the basis 1670 of the quadrupole interaction described above [33]. Therefore, the dominant 1671 channel of the D_2 p-o conversion on ASW might be the one-step process via 1672 the quadrupole interaction. 1673

As another form of a solid water, clathrate has ice-like frameworks composed of H_2O molecules having a regular array of cages, in which small molecules can be encapsulated. O-p conversion of H_2 encapsulated in clathrate was examined at 20 K by NMR [264]. The NMR signal due to o- H_2 was monitored as a function of time over 200 h, which revealed a slight decrease in intensity. Although the conversion kinetics was not unambiguously determined because of the small intensity change, the conversion rate was evaluated to

be 6×10^{-4} /h on the assumption of the first-order kinetics. O-p conversion 168 of H₂ in clathrate hydrates was also investigated with Raman at 20 K [265], 1682 which revealed the conversion rate of $\sim 7 \times 10^{-3}$ /day. These values are much 1683 smaller than the value reported on an ASW surface described above. Since 1684 the cage has a symmetrical structure, the electric field at the inner surface 1685 of the clathrate might be weakened, and the electric-field induced conversion 1686 mechanism is not operative. Dependence of the conversion on the ice surface 1687 structure will be an important future subject. 1688

1689 5.2.2. Si

Although the o-p conversion has not been studied on well-defined Si sur-1690 faces, recent studies with Raman [188, 266, 189] and IRAS [203] reported the 1691 o-p conversion in bulk Si, which is briefly described in this section. In the 1692 work by Hiller et al., H₂ was admitted into Si with the RF plasma treatment 1693 at 220 °C where H_2 was incorporated into the interstitial site of Si. The o-p 1694 conversion rate at 77 K was evaluated to be 0.015 /h, and p-o back conver-1695 sion time at 300 K was evaluated to be 0.18 /h. By comparing the samples 1696 with different oxygen concentrations and dopant concentrations, it was con-1697 cluded that these impurities were not responsible for the o-p conversion in 1698 Si. On the basis of the experimental findings, the origin of the o-p conversion 1699 was attributed to the nuclear spin of 4.7 % abundant ²⁹Si, and the conversion 1700 mechanism was considered to be the one-step process via the magnetic dipole 1701 interaction (Sec. 3.3). Since the H_2 diffusion at 77 K is negligibly slow, the 1702 o-p conversion kinetics was analyzed by considering the distance distribution 1703 of H_2 from the magnetic moment of ²⁹Si, and the o-p conversion time in the 1704 nearest neighbor site was evaluated to be 65 h [188]. On the other hand, 1705 Peng et al. investigated the o-p conversion of H_2 and p-o conversion of D_2 1706 with IRAS. While the o-p conversion time at 77 K and p-o back conversion 1707 at 300 K of H_2 were 229 h and 8.1 h, respectively, the p-o conversion time 1708 at 77 K and o-p back conversion time at 300 K of D_2 were 213 h and 25 1709 h, respectively. In contrast to the multi-component exponential decays ex-1710 pected from the mechanism proposed by Hiller et al., the conversion kinetics 1711 measured by IRAS revealed a single-exponential behavior. As discussed in 1712 Sec. 3.6, the magnetic interaction for H_2 is stronger than that of D_2 due to 1713 the difference of the magnetic g factor. The experiments, however, revealed 1714 that the conversion times for H_2 and D_2 are similar. The IRAS study pro-1715 posed a conversion mechanism where the coupling between the nuclear spins 1716 and rotational motion of the molecule center of mass induces ortho-to-para 171

¹⁷¹⁸ conversion [203], which needs further verification.

1719 5.2.3. Metal-organic framework

Metal-organic framework (MOF) compounds are a class of materials con-1720 sisting of inorganic clusters connected by organic ligands, which have a peri-1721 odic well-ordered structure. MOF's have micropores, in which H_2 is expected 1722 to be stored with a high binding energy. In relation to the hydrogen storage, 1723 the binding property of H_2 to the inner surface of MOF's and the o-p conver-1724 sion on the surface have recently been investigated with IRAS [267, 204, 205], 1725 although the pore surface is not characterized with surface science techniques. 1726 As MOF's without magnetic cations are non-magnetic insulators, the exper-1727 imental work performed for $Zn_2(C_8H_2O_6)$ named MOF-74 is shown in this 1728 section [204]. 1729

The MOF-74 sample has a periodic array of micropore with a diameter 1730 of 1.1 nm and parallel channels. The inner surface of the pore is lined with 1731 Zn^{2+} ions, and the Zn-O electric dipole is exposed at the pore surface. After 1732 the powder sample was fixed to a cold head and H₂ was introduced into the 1733 chamber at 30 - 50 K, the (ro)vibrational spectra for H₂ physisorbed on 1734 the sample were obtained with IRAS. Figure 45 shows the time evolution 1735 of the IRAS spectrum taken for H_2 on MOF-74. Two pairs (labeled 1° and 1736 2°) of absorption bands corresponding to $v=0\rightarrow 1$ are observed. Each pair 1737 consists of two distinct peaks, which were attributed to $Q(1)(J=1\rightarrow 1)$ of 1738 o-H₂ and Q(0)($J=0\rightarrow 0$) of p-H₂. As described in Sec. 4, o-H₂ in J=1 and p-1739 H_2 in J=0 are clearly distinguished. The two pairs with different vibrational 1740 frequencies are H_2 bound to different sites. The spectral change shown in 1741 the figure, furthermore, represents time evolution of the IRAS. Evidently, 1742 Q(1) of o-H₂ is reduced in intensity, while the intensity of p-H₂ increases 1743 with increasing time indicating o-p conversion. The o-p conversion time is 1744 estimated to be the order of minutes. As the o-p conversion mechanism, the 1745 authors suspect presence of magnetic impurities, which induce o-p conversion 1746 as the one-step process via the magnetic dipole or Fermi contact interaction 1747 (Sec. 3.3). The authors also point out that the o-p conversion time on a 1748 different type of MOF (MOF-5) is 25 %/h much slower than that on MOF-1749 74. No further analysis and discussion are given in the paper at present. It 1750 should be noted that the Zn-O electric dipole is exposed at the pore surface, 1751 which is similar to the amorphous ice surface. Therefore, the multi-step o-p 1752 conversion process as discussed for the ice surface might be operative on the 1753 MOF surface. 1754



Figure 45: IRAS spectra taken for H₂ physisorbed on the pore surfaces of MOF-74 at 30 K. The absorption bands labeled 1° and 2° are assigned to different adsorption sites. Q(1) and Q(0) correspond to o-H₂ and p-H₂, respectively. Traces from green to brown are collected at t=55, 95, 135, 175, 280, 600, and 2700 s after H₂ loading. Q(1) is reduced in intensity, while the intensity of Q(0) increases with increasing time. Reproduced by permission from [204].

The IRAS spectra also revealed absorption bands corresponding to the 1755 $S(0)(J=0\rightarrow 2)$ and $S(1)(J=1\rightarrow 3)$ transitions. These bands are observed to be 1756 split due to the lifting of the rotational-state degeneracy under the anisotropic 1757 potential. The number of the splitting and relative absorption intensity are 1758 not consistent with the theoretical formula of Eq. 7, which are ascribed to 1759 possible ΔM selection in the IRAS observation [204]. The IRAS data for 1760 MOF-5 are compared with the recent theoretical work based on the first-1761 principles calculations [65, 206]. 1762

1763 5.2.4. Impurity-adsorbed Ice surfaces

¹⁷⁶⁴ Chehrouri et al. investigated o-p conversion of H_2 and p-o conversion of ¹⁷⁶⁵ D_2 on an amorphous solid water (ASW) surface co-adsorbed with O_2 at 10 ¹⁷⁶⁶ K by the REMPI-TPD method [226]. ASW samples were prepared by two



Figure 46: Time evolution of D_2 in p- D_2 (J = 1) and o- D_2 (J = 0) on porous ASW with 0.012 ML of co-adsorbed O_2 measured by REMPI-TPD. The J=1 intensity decreases while the J=0 intensity increases with increasing time. Adapted by permission from [226].

¹⁷⁶⁷ steps: the initial condensation of 250 ML of water on a Cu block at 120 K ¹⁷⁶⁸ leading to compact non-porous ice film formation, and subsequent 12 ML ¹⁷⁶⁹ porous ASW deposition at 10 K. After exposing the sample to O_2 at 25 K ¹⁷⁷⁰ allowing sufficient diffusion of O_2 in porous ASW, n-H₂ was dosed to the ¹⁷⁷¹ sample at 10 K.

Figure 46 shows the time evolution of the J=0 and J=1 of D_2 on ASW 1772 with O_2 of 0.012 ML. The J=0 intensity increased while the J=1 inten-1773 sity decreased with increasing residence time. By fitting a function of A +1774 $B \exp(-t/\tau)$, the p-o conversion time of D₂ and o-p conversion time of H₂ at 1775 this O_2 coverage were estimated to be 600 ± 50 and 220 ± 50 s, respectively. 1776 Since the effective surface area of the ASW sample is larger than the flat 1777 surface by a factor of >5, the O₂ concentration was estimated to be as low 1778 as 0.2 %. The conversion times were similar to those observed on an O_2 -1779 adsorbed Ag surface [225], suggesting that the mechanism originating from 1780 the O_2 spin, as discussed in Sec 5.1.3, is operative on the O_2 -adsorbed ASW. 1781 It was pointed out that the asymptotic value (A) of J=1/J=0 for D_2 1782 might be higher than the thermal-equilibrium value in gas phase (3×10^{-4}) . 1783

The origin of this difference was discussed to be the rotational-sublevel split-1784 ting of the J=1 state as shown in Fig. 9. The energy separation between 1785 the J=1 and 0 states gets smaller in the adsorption state compared to that 1786 of the gas phase due to the potential anisotropy, which results in enhanced 1787 population in J = 1 in thermal equilibrium compared to the gas phase. This 1788 is an interesting suggestion, and more accurate estimation of the asymptotic 1789 value is strongly required. On the other hand, the asymptotic value for H_2 1790 was recognized to be too high and its possible origins were discussed to be 1791 either backconversion during TPD, presence of non-converting species, and 1792 background adsorption. At a low H_2 coverage, some molecules are trapped 1793 at a large-binding site because of a distribution of binding energies on ASW 1794 [55, 268], which suppresses the diffusion of H_2/D_2 to the neighboring site and 1795 hinders the spin conversion. 1796

1797 6. Astronomical relevance

Hydrogen is astronomically important, because it constitutes a significant 1798 fraction of the Universe, which includes stars and interstellar media (ISM). 1799 While the stars are hot and dense, ISM are cold and dilute with a typical 1800 density and temperature of 1 cm^{-3} and 100 K, respectively. The density and 1801 temperature of ISM are not uniform but dynamically changing spatially and 1802 temporally. Density fluctuation might cause further gravitational contraction 1803 eventually leading to star formation. Therefore, spatially-resolved observa-1804 tion of the density and temperature of ISM is of considerable importance for 1805 astronomical physics [27, 28]. 1806

1807 6.1. Hydrogen in Molecular clouds: importance of surface processes

ISM consists of gas and dust, and the original main constituent of gas is 1808 atomic hydrogen. Figure 47(a) shows the main three steps of the evolution 1809 of hydrogen. As the first step, atomic hydrogen undergoes recombinative 1810 reaction forming molecular hydrogen on dust surfaces. This eventually leads 1811 to formation of molecular clouds with a temperature of about 30 K as the 1812 second step, which are regarded as a precursor state of star formation [28, 1813 269]. Although the temperature of the molecular cloud is as low as 30 K, 1814 the surface part of the cloud is heated up to $\sim 10^3$ K by either radiation or 1815 particles incident to the molecular clouds in the astronomical environment 1816 as schematically shown in Fig. 47(c). 1817

Figure 47(b) schematically illustrates the molecular processes taking place 1818 on the dust surfaces. It is worth emphasizing that the reaction of $H + H \rightarrow H_2$ 1819 is slow in gas phase and only occurs on solid surfaces. Two possible reaction 1820 paths for H_2 formation are the Eley-Rideal (ER) reaction where an impinging 1821 H atom directly reacts with adsorbed H or the Langmuir-Hinshelwood (LH) 1822 reaction where a weakly physisorbed H atom diffuses on the surface recom-1823 bining with another H atom either physisorbed or chemisorbed on surfaces. 1824 H_2 molecules are vibrationally and rotationally excited upon H_2 formation 1825 because of the exothermic character of the reaction. An important question 1826 is the rotational-state distribution and OPR (ortho-para ratio) of nascent 1827 H_2 formed on the surfaces [270]. These would strongly depend on the dust 1828 surface structure, the adsorption site of H, the temperature, and so on. 1829

When the H_2 molecules are cooled, they possibly interact with dust sur-1830 faces and undergo o-p conversion, which modifies the OPR of the molecular 1831 cloud. The time scale of the o-p conversion should depend on the magnetic 1832 and electronic structure of the dust surfaces and the sticking probability and 1833 residence time of H_2 on the dust surfaces. During the heating process of the 1834 surface part of molecular clouds, furthermore, H₂ might undergo o-p conver-1835 sion again. During heating by radiation of either X-rays or far-ultraviolet 1836 radiation (FUV)), H_2 is once electronically excited and decays into the elec-1837 tronic ground state with simultaneous population of vibrationally and rota-1838 tionally excited states [271]. The electronic excitation sometimes leads to 1839 dissociation of H_2 , therefore this heated region is called photo-dissociated 1840 region or photon-dominated region (PDR). On the other hand, particles at a 1841 high velocity emitted from young stars and supernovae as called shocks are 1842 incident to molecular clouds and cause collisions with hydrogen molecules in-1843 ducing excitation. As described in Sec. 3, o-p conversion does not occur by 1844 simple collision with other H_2 or excitation by radiation. The o-p conversion 1845 only occurs through interaction with dust surfaces or H^+ (and possibly H_3^+) 1846 in gas phase (Sec. 3.1). It should be noted that the chemical o-p conver-1847 sion process on dust surfaces described in the introduction section could be 1848 important at high temperature as well as the physical o-p conversion. 1849

As described in Sec. 2.1, the rotational-state distribution and OPR in a thermal equilibrium are expressed by the partition functions. Since the line intensity reflects the population of each rovibrational state, the temperature of the gas can be inferred from the spectral observation. On the other hand, the populations in the ortho and para states are not necessarily in thermal equilibrium with the environment, because the timescale of the o-p conversion might be significantly longer than the timescale of the thermal change. The
OPR therefore reflects the thermal history of the astronomical object [27, 28].
Spectroscopic measurements of molecular clouds are therefore highly required
in the astronomical community, which are shown in the next section.

1860 6.2. Astronomical observation

1861 6.2.1. Light emission of H_2

The spectrum and line intensities of H_2 serve as a probe of the density and temperature. As detailed in Sec. 2.1, molecular hydrogen in the electronically ground state has no electric dipole moment. Therefore, light emission or absorption due to rotational and vibrational motion of hydrogen molecules is not induced by the electric dipole transition but originates from the electric quadrupole moment with a low oscillator strength. The quadrupole moment Q_m of an axially symmetric molecule is generally described as [272]

$$Q_m(R,\theta,\phi) = Q(R)Y_{2m}(\theta,\phi), \qquad (28)$$

where the angular part is expressed by the spherical harmonics Y_{2m} with 1869 respect to the angle of the molecular axis. By considering the matrix ele-1870 ment of Q_m with respect to two rotational states, the selection rule for the 1871 rotational transition is obtained as $\Delta J = 0, \pm 2$. When Q is expressed in the 1872 Taylor expansion of the intramolecular distance R, furthermore, the zero-th 1873 and first terms lead to the selection rule of $\Delta v = 0$ and ± 1 , respectively. The 1874 energies and wavelengths for these transitions are listed in Table 5. From the 1875 viewpoint of astronomical observation, the emission due to vibrational de-1876 excitation ($\Delta v = -1$) is transparent through air, whereas pure rotational 1877 transition $(\Delta v = 0, \Delta J = -2)$ is air-opaque. 1878

However, the rotational-excitation energy of molecular hydrogen is as 1879 large as 540 K (the $J=0 \rightarrow 2$ transition energy). Most of the H₂ molecules 1880 in molecular clouds are therefore in the rotationally ground states and no 1881 emission is expected at a typical molecular cloud temperature of 30 K. This 1882 is in contrast to atomic H, which has a hyperfine splitting due to coupling of 1883 the nuclear spin and electron spin. Transition between the hyperfine levels 1884 yields emission with a wavelength of 21 cm, which has allowed a variety 1885 of astronomical observation. Although H_2 in molecular clouds is generally 1886 invisible, the surface part of molecular clouds at a temperature of $\sim 10^3$ K 1887 shown in Fig. 47(c) yields emission of H₂. The emission spectra of H₂ 1888 originating from the rotational and vibrational levels have been reported 1889



Figure 47: Schematic illustration of (a) evolution of hydrogen, (b) cross-section of a molecular cloud and heating of the cloud surface and (c) molecular processes on dust surfaces, during formation of molecular clouds. In the first step, H_2 is formed from atomic hydrogen on dust surfaces via the either Eley-Rideal (ER) reaction or Langmuir-Hinshelwood (LH) reaction. In the next steps, H_2 is cooled to about 30 K forming molecular clouds, while the surface part of the clouds are heated by radiation or collision. The OPR in these processes are of interest and importance, the dust surfaces playing a role to induce the o-p conversion.

for a variety of heated regions of molecular clouds. The observation resultsrelevant to OPR are reviewed in the next subsection.

1892 6.2.2. Observational spectra

The rotational spectra of the vibrationally excited states were first re-1893 ported for a molecular cloud named the NGC2023 nebula [273]. The emis-1894 sion lines originating from the vibrational transitions, which lie in the air-1895 transparent region, were detected by the ground-based infrared telescope in 1896 Hawaii. As described in Sec. 2.1, the vibrational wavefunction of H_2 does 189 not change its sign with respect to the two-nuclei permutation, which means 1898 the ortho and para species can be defined in the same way as the vibrational 1899 ground state. Figure 48 shows the Boltzmann plot of the rovibrational-state 1900 intensities often designated as a (ro)vibrational-energy diagram, where the 1901 intensity divided by the nuclear-spin and rotational degeneracy is plotted as 1902 a function of the rovibrational energy. From the relative population of the 1903 v=1 and 2 states, the vibrational temperature of 3600 K was derived. In the 1904 v=2 state, while the relative populations of J=2 and 4 and J=3 and 5 reveal 1905 a rotational temperature of 1500 K, the ortho (J=odd) and para (J=even)1906 populations are evidently not in equilibrium with each other. The OPR was 1907 estimated to be 1.4 - 2.0, which corresponds to a nuclear-spin temperature 1908 of 95 - 120 K in Fig. 7. If the ortho and para states were in equilibrium 1909 with a temperature of 1500 K, the OPR should be 3. Following this work, 1910 the OPR value was investigated for several nebulas through observation of 1911 the S branches $(\Delta J = -2)$ of $v = 1 \rightarrow 0$, i.e. $J = 2 \rightarrow 0$ and $J = 3 \rightarrow 1$ 1912 transitions. The OPR values evaluated from the line intensities ranged from 1913 1 to 2.5 [274, 275, 276, 277, 278, 279, 280, 281]. 1914

Additionally, in a dense and cold region of a molecular cloud, absorption of the background light due to vibrational excitation (v=0-1) of J=0 (para) and J=1 (ortho) was observed toward a nebula of NGC2024. The J=1to J=0 ratio was smaller than 0.8, which is roughly consistent with the molecular temperature of ~ 30 K [282] (Fig. 7).

¹⁹²⁰ In later work, however, it was pointed out that the OPR in the vibra-¹⁹²¹ tionally excited state possibly deviates from the real OPR of the molecular ¹⁹²² cloud because of the different optical depth of the ortho and para species ¹⁹²³ [284].

To that end, the air-opaque IR emission corresponding to the pure rotational transitions has been observed to probe the OPR in the vibrationally ground state. The Orion bar region was first studied by a ground-based



Figure 48: Boltzmann plot of the rovibrational-level population of H_2 observed for the NGC2023 nebula by a ground-based infrared telescope. The rotational temperatures obtained from the dashed lines and vibrational temperatures evaluated from the solid lines are shown in the figure. The populations of o-H₂ and p-H₂ are not in thermal equilibrium with the rotational temperatures, and the OPR is 1.4 - 2.0. Adapted with permission from [273].



Figure 49: Boltzmann plot of the rotational-level population of H_2 in J=3-7 for a Herbig Hallo object of HH54 as observed by ISO. Straight lines show the fit for a gas temperature of 650 K and an OPR of 1.2 [283].

infrared telescope, and the ratio of J=3 and J=4 was reported to be 0.5 1927 -3.6 [285, 286] though the sensitivity was limited. Owing to the infrared 1928 astronomical satellite named the Infrared Space Observatory (ISO), which 1929 was launched in 1995, the pure rotational emission lines of H_2 in v=0 were 1930 successfully measured with a high sensitivity [287]. An overview of the data 1931 acquired by ISO for a plenty of heated regions of molecular clouds such as 1932 shocked regions and PDR's of nebulas are given in a review article by Habart 1933 et al. [28]. Some of the typical results are briefly shown in the following. 1934

Figure 49 shows the rotational-energy diagram observed for a Herbig 1935 Hallo object named HH54 [283]. Herbig Hallo objects are shock-heated ISM's 1936 near newly-born stars. The straight line represents a rotational temperature 1937 of 650 K. As clearly seen in the figure, the ortho and para state populations 1938 are not on the same line, indicating that the two species are not in thermal 1939 equilibrium with the rotational temperature. The OPR value was estimated 1940 to be 1.2, which corresponds to a nuclear-spin temperature of lower than 90 1941 K (Fig. 7). 1942

A comprehensive set of the rotational and rovibrational lines in the wave-1943 length range of $2.4 - 45 \ \mu m$ was observed for the Orion molecular cloud, 1944 OMC-1 [288], which is shown in Fig. 50(a). In addition to the H_2 lines for 1945 v=0-2 and J=1-25, emission lines due to other elements such as Si, S, 1946 and Ne were identified. Figure 50(b) shows the rovibrational energy diagram 1947 of H_2 , where the level population was well described by a sum of five Boltz-1948 mann distributions with temperatures of about 600 - 3000 K as shown by 1949 the dotted curve. Since all rovibrational states are on the same curve, the 1950 OPR of this cloud was concluded to be 3. 1951

While another shock-heated region of HH2 was found to have an OPR 1952 of 1.2–1.6 [289], most shocked regions are reported to have an OPR value of 1953 3 [290, 291, 292, 293, 294]. The H_2 spectra were also observed for a PDR 1954 named the ρ Ophiuchi molecular cloud [295]. The spectrum was measured 1955 in a spatial-resolved manner, and revealed a rotational temperature of about 1956 300 K. It was also shown that the OPR was significantly smaller than 3. 1957 Smaller OPR values were similarly observed for NGC7023 [296] and galactic 1958 center [297]. 1959

After ISO completed its mission, Spitzer Space Telescope (SST) was launched in 2003, and has reported H₂ spectral data with higher spatial resolution from various molecular clouds. Figure 51 shows the rotational-energy diagram observed towards the shock-heated regions of HH54 and HH7–11 [298]. The diagram revealed a zigzag feature indicating that the ortho and



Figure 50: (a) Spectrum in the wavelength range of $2.4 - 45 \ \mu m$ observed for the Orion molecular cloud of OMC-1 by ISO. (b) Boltzmann plot of the extinction-corrected rovibrational-level intensity obtained from (a). Rovibrational-energy positions for the v=0-4 and various J states are shown in the figure. The OPR is evaluated to be 3. Reproduced with permission from [288].

para states are not in equilibrium with the gas temperature. The rotational-1965 state distribution was well expressed by two components with temperatures 1966 of 400 - 550 K and OPR of 1.5 - 2. The rotational temperature and OPR 1967 value derived from the J=4-7 intensities were two-dimensionally mapped for 1968 the two directions as shown in Fig. 52(a). It is apparent that the rotational 1969 temperature and OPR value vary depending on the position in the object. 1970 While the rotational temperature is 600 - 1000 K, the OPR changes from 0.5 1971 3. Figure 52(b) displays the correlation between the OPR and rotational 1972 temperature obtained from two-component fits, where the OPR was plotted 1973 as a function of the rotational temperature. The OPR values tend to in-1974 crease as the rotational temperature increases. The orange curve represents 1975 the thermal equilibrium, and the black curve displays the simulated result 1976 on the assumption of the initial OPR of 0.4 (HH54) and 0.25 (HH7) followed 1977 by o-p conversion via the proton exchange mechanism (Sec. 3.1). 1978

HH 7-11 are Herbig-Haro objects in the Perseus molecular cloud, and 1979 associated with the NGC 1333 star-forming region. The NGC 1333 region 1980 was also investigated by Spitzer in detail [299]. The measured rotational-1981 energy diagrams were analyzed with two components. While the obtained 1982 rotational temperatures were 300 - 600 K and 1000 - 1500 K, the OPR's 1983 were 0.3 - 0.7 and 1.4 - 2, respectively. The rotational temperature and OPR 1984 were precisely mapped in detail similarly to Fig. 52. The outflows associated 1985 with star formation were also analyzed in detail. Whereas the OPR of L1448, 1986 BHR71, NGC2071, and L1157 [300, 301, 302] were significantly smaller than 1987 3, those of IRAS 16253-2429 protostar and HH211 were found to be 3 [303, 1988 304]1989

The rotational temperature and OPR were also investigated for photon-1990 heated PDR regions of molecular clouds. The PDR is characterized by the 1991 strength of the radiation field, and six galactic PDR's named L1721, Califor-1992 nia, N7023E, Horsehead, ρ Oph and N2023N at different excitation conditions 1993 were analyzed in detail [269]. As demonstrated for shock-heated regions, the 1994 rotational-energy diagram was expressed by two temperature components, 1995 where the temperatures for the low-J and high-J regions were 200 - 3001996 K and 300 - 700 K, respectively. The OPR value of the low-J and high-J 199 regions was 1 - 1.6 and 2.6, respectively. The rotational temperature was 1998 confirmed to be higher than the value expected from a model calculation, 1999 which was also observed for a reflection Nebula of NGC2023 with a low OPR 2000 of 1.68 [305]. The rotational spectrum was further investigated for a lower 2001 excitation condition [306]. The next generation Herschel Space Observatory 2002



Figure 51: Boltzmann plot of the rotational-level population observed toward different directions of HH54 and HH7. The solid lines show the two-component fit to the observations, and the OPR is estimated to be 1.5 - 2. Reproduced with permission from [298].



Figure 52: (a) Mapping of the rotational temperature and OPR towards HH54 and HH7. The horizontal and vertical axes are the right ascension and declination offsets in arcseconds relative to a certain direction. (b) Correlation of the OPR and rotational temperature. Triangles are from a two-component \Re^{0} to the rotational-level population shown in Fig. 51. Other data points apply to $5^{\circ} \times 5^{\circ}$ square subregions within the source regions shown in (a), green and red referring, respectively, to the warm and hot gas components. The orange curve represents the thermal equilibrium, and the black curve displays the simulated result on the assumption of the initial OPR of 0.4 (HH54) and 0.25 (HH7) followed by o-p conversion via the proton exchange mechanism. Reproduced with permission from [298].

²⁰⁰³ will provide us with new observation data.

$_{2004}$ 6.3. Ortho-para ratio of nascent H_2 and ortho-para conversion

As discussed for Fig. 47, the initial OPR value when H_2 is formed is of particular importance. Since the formation of H_2 molecules is an exothermic process, part of the energy released in the chemical bond formation is partitioned in the internal energy of nascent H_2 . In recent years, the rotationalstate distribution and OPR have been investigated by laboratory experiments and theoretical studies, which are reviewed along with discussion on possible o-p conversion in an astronomical environment.

2012 6.3.1. H_2 formation

Most H_2 is thought to be produced on surfaces of interstellar grains [307, 2013 308]. We first discuss why H₂ formation is slow in gas phase. In gas phase, two 2014 possible configurations when two hydrogen atoms come closer are the electron 2015 spin triplet and singlet states. The triplet state of $b^3 \Sigma_u^+$ is a repulsive state, 2016 and the radiative-transition to the ground state $X^1\Sigma_g^+$ is spin-forbidden with a lifetime of much longer (by a factor of $10^8 - 10^{10}$) than the ordinary dipole 2017 2018 allowed transition of $\sim 10^{-9}$ s [309, 307]. The spin singlet configuration, on 2019 the other hand, corresponds to a vibrationally continuum unbound state in 2020 the electronically ground state. As H_2 in the $X^1\Sigma_g^+$ state has no electric 2021 dipole, the radiative transition probability should be small. The vibrational 2022 period of $\sim 10^{-14}$ s is much shorter than the lifetime of these states, hence, 2023 the recombination efficiency becomes extremely small [310]. 2024

Recombinative desorption and its reverse process of dissociative chemisorp-2025 tion of H_2 on metal surfaces have been a topic of intensive studies in the last 2026 decades in Surface Science. Hydrogen atoms are chemisorbed on most metals 2027 with a typical chemisorption energy of 0.5 eV, and the dynamics of recom-2028 binative desorption dissociative adsorption was investigated on the basis of 2029 multi-dimensional potential energy surfaces. As detailed in the Surface Sci-2030 ence textbooks and several review articles [311, 312, 313, 314, 315, 316, 317], 2031 a considerable progress has been made for the understanding of the dynamics 2032 of desorption and adsorption on metal surfaces. Since the chemisorption en-2033 ergy is large, the recombinative desorption occurs at temperatures of higher 2034 than 100 K, which is not of direct relevance to the astronomical condition. 2035 When two H atoms encounter, four possible nuclear-spin configurations 2036

$$|\frac{1}{2}\rangle + |\frac{1}{2}\rangle, \ |-\frac{1}{2}\rangle + |-\frac{1}{2}\rangle, \ |\frac{1}{2}\rangle + |-\frac{1}{2}\rangle, \ |\frac{1}{2}\rangle + |-\frac{1}{2}\rangle.$$

Very naively, it is expected that the former two configurations lead to $o-H_2$ 2038 formation, whereas the latter two result in $o-H_2$ and $p-H_2$ with the same 2039 probability. When there is no particular J-selection in H_2 formation and 2040 the rotational-state is described by a Boltzmann distribution with a certain 2041 temperature, the OPR becomes in thermal equilibrium with the rotational 2042 temperature. If there were some tricks operative to polarize or select the 2043 proton nuclear-spin during H_2 formation, the OPR would be affected in favor 2044 of o-H₂ formation. It should be noted, however, that the electron spin of an H 2045 atom is coupled through hyperfine interaction with the proton spin forming 2046 either total spin singlet or triplet states. Hence, the H_2 formation process 2047 should be described as the four-spin dynamics. This might strongly depend 2048 on the adsorption state of H, because the H electron spin is retained in the 2049 physisorbed state while it tends to form a singlet state through chemical 2050 bond formation with substrate electrons in the chemisorbed state. 2051

As an experimental approach, the H_2 formation rate and its mechanism on the surfaces of ISM such as graphite, amorphous ice and silicate have been investigated as reviewed in the articles by Williams et al. [318] and Watanabe and Kouchi [319]. Recent studies on the OPR and internal-state distribution of H_2 from physisorbed states on surfaces are briefly reviewed below.

The OPR of the newly formed H_2 from spin-polarized H atoms was investigated by NMR. It is known that the NMR spin relaxation time is a function of the ortho concentration of solid H_2 , which allowed to measure the OPR of H_2 recombined on liquid He surfaces. According to this study, the OPR formed from H atoms with random spin orientation was estimated to be 0.2, while the spin-polarized H atoms yielded 100 % pure o- H_2 [320].

By applying the rotational-state-resolved techniques, the internal-energy 2064 distribution was measured for molecular hydrogen formed on graphite [321, 2065 322, 323. Atomic H and D beams were supplied to the surface at 15 - 502066 K, and resultant H_2 and HD molecules formed on the surface were state-2067 selectively detected. The product molecules were in the vibrationally excited 2068 states of v=1-7. For v=1 and 2, the rotational distribution in J=0-42069 revealed a rotational temperature of about 300 K, which was furthermore 2070 found to be independent of the surface temperature of 15 - 50 K [321]. The 2071 rotational temperature of higher v states was also examined experimentally, 2072 which were similar to those of v=1 and 2 states [322, 323]. In these stud-2073 ies, the rotational-state distribution seemed to be described by a Boltzmann 2074 distribution, which suggests no particular selection between even and odd J2075

 $_{2076}\,$ states. The OPR of H₂ formed on graphite, therefore, was considered to be $_{2077}\,$ 3.

The internal-state distribution of nascent H_2 was also investigated on 2078 amorphous ice surfaces [324, 325, 257, 326]. During atomic H dosage onto 2079 amorphous ice at 8 K, H_2 formed on the surface in J=1 and 0 was detected by 2080 REMPI, and the ratio of these species was consistent with an OPR value of 3 2081 [257]. The OPR was also examined for both H_2 and D_2 formed on amorphous 2082 ice at 10 K [326]. In this work, either atomic H or D was supplied to the 2083 sample surface and product H_2 or D_2 formed on the ice surface was detected 2084 by REMPI. By comparing the populations in J=1 and 0, the OPR of H₂ and 2085 D_2 was evaluated to be 2.91 \pm 1.23 and 1.44 \pm 0.43, respectively. Although 2086 the experimental uncertainty was large due to the background signal and 208 experimental difficulty, these OPR values were claimed to be consistent with 2088 the high-temperature limit of 3 (H_2) and 2 (D_2) . 2089

Theoretically, the rotational-state distribution upon H_2 formation was 2090 first simply assumed to be in thermal equilibrium with the surface tempera-2091 ture or the effective temperature corresponding to the energy released in the 2092 rotational degree of freedom [327, 328]. An interesting theoretical approach 2093 was application of the resonance theory for atom recombination. The highest 2094 vibrational level of H_2 is the v=14 state. At high vibrational levels, the to-2095 tal internal energy goes beyond the dissociation energy with high rotational 2096 states, referred to as quasibound states or orbiting resonances because of 2097 the bound nature within the centrifugal limit [329]. The orbiting resonance 2098 theory for atom recombination considers that two atoms are scattered into 2099 such quasibound states followed by rotational deexcitation to bound states 2100 [330, 331, 332]. This idea was also applied to hydrogen recombination for 2101 physisorbed systems on Xe crystal surfaces [333], where two hydrogen atoms 2102 were assumed to move freely on the surface and undergo collisions forming 2103 quasibound states. The surface serves to conserve the energy and momentum 2104 of the system while H₂ falls into a true bound state. The surface corrugation 2105 is important providing coupling between different degrees of freedom. Two 2106 dominant resonances are the (v=14, J=5) and (v=13, J=8) states, which 210 correspond to the ortho and para states. The resonance probabilities then 2108 determine the OPR of H_2 formed on surfaces. 2109

A more elaborate theoretical approach is being conducted in recent years. As successfully performed on metal surfaces, the hydrogen recombinative reaction is essentially governed by the multidimensional potential energy surface (PES) and dynamics of atoms on the PES. Description of PES and clas-

sical or quantum dynamics of hydrogen atoms have recently been performed 2114 on the surface of graphite. As an LH reaction illustrated in Fig. 47(b), PES 2115 for two physisorbed H atoms on graphite was described semiempirically or 2116 by first-principles calculations, and the H_2 formation probability and rovi-2117 brational state distribution in the final state were quantum-mechanically 2118 simulated as a function of the collision energy [334, 335, 336, 337]. It 2119 was shown that the product H_2 is strongly excited in the vibrational mo-2120 tion and the rotational motion is moderately excited with an OPR value 2121 of around 3. The reaction dynamics of the ER type was also investigated 2122 in detail. The PES was obtained by first-principles density functional the-2123 ory (DFT) [338, 339, 340, 341, 342, 343], and model potentials were occa-2124 sionally used by fitting to the DFT potential [344]. On the basis of these 2125 PES's, the reaction dynamics was investigated at a variety of conditions 2126 [345, 346, 339, 347, 348, 340, 349, 344, 350, 341, 351, 352, 353, 342], e.g.,2127 the initial state of the preadsorbed H, either chemisorbed [344], physisorbed 2128 [350], or paired with another H on graphite [348, 341], the substrate re-2129 laxation effect [340], relation with other reaction paths of collision-induced 2130 desorption and trapping as well as the ER reaction [349], and isotope effects 2131 [347]. All these studies showed the product H_2 is strongly excited vibra-2132 tionally, sometimes occupying a particular vibrational state. Although the 2133 rotational-state distribution has not been thoroughly investigated as vibra-2134 tional excitation, the reaction probability to form ortho and para species was 2135 similar in the ER reaction [345]. Further theoretical studies are expected to 2136 clarify the final-state distribution and resultant OPR on various ISM surfaces 2137 in the near future. 2138

2139 6.3.2. Ortho-para conversion

The H₂ molecules in molecular clouds are cooled to ~ 30 K from the nascent hot condition, yet the surface part of the clouds is heated by radiation or collision (Fig. 47). Without o-p conversion during cooling and heating processes, the OPR in the nascent condition would be retained. During the thermal processes, however, H₂ interacts with the surfaces of cosmic dusts consisting of carbon, water, silicates and so on, and possibly undergoes o-p conversion.

In considering the o-p conversion on dust surfaces, the collision rate γ , sticking coefficient S_t and residence time t_R of H₂ on the surfaces are important. Since the typical desorption temperature of H₂ is lower than 30 K, t_R might be shorter than the o-p conversion time τ . The number of the ²¹⁵¹ physisorbed o-H₂ decreases by either desorption from the dust surface with ²¹⁵² a probability of t_R^{-1} or o-p conversion with a probability of τ^{-1} . Therefore, ²¹⁵³ the overall conversion probability P is described as

$$P = S_t \gamma \frac{\tau^{-1}}{t_R^{-1} + \tau^{-1}}.$$
(29)

As discussed in Sec. 2.2.3, S_t is typically of the order of 0.1. On the other 2154 hand, t_R changes exponentially as $t_R = t_0 \exp(E_d/kT)$ ($t_0 \sim 10^{-11}$ s), and 2155 might be as short as 10^{-6} s at 30 K and a physisorption energy of 30 meV. 2156 As described in Sec. 5, the microscopic o-p conversion times of H_2 in the 2157 physisorption state of amorphous ice and graphite were obtained to be $\sim 10^3$ 2158 s or shorter (Secs. 5.1.4 and 5.2.1). Although the conversion data on oxides 2159 like silicates are not available to date, the electric-field induced multi-step 2160 mechanism could be operative on those surfaces. On the other hand, the 2161 collision rate is described by the product of the dust size σ , mean velocity of 2162 H₂ and dust density [37]. On the assumption of the dust size of $\sim 10^{-12}$ cm², 2163 the mean velocity of $\sim 10^5$ cm s⁻¹, and dust density of ~ 1 cm⁻³ [269], γ leads 2164 to 10^{-7} s⁻¹, and the overall o-p conversion probability is roughly estimated to 2165 be 10^{-17} s⁻¹ for ASW. If some magnetic impurities like O₂ and H are present 2166 on the surfaces, the overall o-p conversion probability could be enhanced 2167 to 10^{-14} s⁻¹. Since the physisorption energy depends on the local atomic 2168 structure, as discussed by Bourlot [29], the overall conversion probability 2169 is strongly temperature-dependent and might be further enhanced if H_2 is 2170 strongly bound and the residence time is longer. As these timescales are 2171 similar to the timescale of the thermal process of molecular clouds ($\sim 10^{15}$ s), 2172 the o-p conversion on surfaces might have significant effects on the OPR of 2173 the molecular clouds. 2174

Another important o-p conversion channel in the heated region is the H_2 + 2175 H^+ reaction via proton exchange described in Sec. 3.1. The o-p conversion 2176 rate via the proton exchange was roughly evaluated to be $\sim 10^{-16}$ s⁻¹ for 2177 a particular condition [354], and analysis has been performed for various 2178 astronomical parameters [354, 355, 356, 357]. Proper modeling on the basis 2179 of these time scales of the microscopic molecular processes would elucidate 2180 the origin of the astronomical data and dynamical evolution of molecular 2181 hydrogen. 2182

2183 7. Concluding remarks

We have reviewed recent progress on the physisorption and o-p conversion 2184 of molecular hydrogen along with basic concepts of the ortho-para conver-2185 sion of molecular hydrogen. In spite of the recent advances, experimental 2186 data on various surfaces are still missing and quantitative verification of the 2187 conversion theory including magnetic interaction and energy dissipation is 2188 strongly required. We have also tried to provide astronomical observation 2189 data relevant to ortho-para conversion, which will hopefully be understood 2190 on the basis of the microscopic knowledge acquired in Surface Science. 2191

2192 8. Acknowledgement

The authors thank T. Okano, K. Niki and K. Yamakawa for fruitful discussion, and support by a Grant-in-Aid for Scientific Research by Japan Society for the Promotion of Science.

²¹⁹⁶ Appendix A. Matrix element of \vec{i}

$$<\chi_{n}(I=0)| \quad i^{-} \quad |\chi_{n}(I=1, I_{z}=+1) >$$

$$= \frac{1}{\sqrt{2}}(<\frac{1}{2} - \frac{1}{2}| - < -\frac{1}{2} \frac{1}{2}|)(i_{a}^{-} - i_{b}^{-})|\frac{1}{2} \frac{1}{2} >$$

$$= \frac{1}{\sqrt{2}}(<\frac{1}{2} - \frac{1}{2}| - < -\frac{1}{2} \frac{1}{2}|)(|-\frac{1}{2} \frac{1}{2} > -|\frac{1}{2} - \frac{1}{2} >)$$

$$= -\sqrt{2}$$
(A.1)

$$<\chi_{n}(I=0)| \quad i^{z} \quad |\chi_{n}(I=1,I_{z}=0)> \\ = \frac{1}{2}(<\frac{1}{2} - \frac{1}{2}|-<-\frac{1}{2} \frac{1}{2}|)(i^{z}_{a} - i^{z}_{b})(|\frac{1}{2} - \frac{1}{2}>+|-\frac{1}{2} \frac{1}{2}>) \\ = \frac{1}{2}(<\frac{1}{2} - \frac{1}{2}|-<-\frac{1}{2} \frac{1}{2}|)(|\frac{1}{2} - \frac{1}{2}>-|-\frac{1}{2} \frac{1}{2}>) \\ = 1$$
 (A.2)

²¹⁹⁷ Appendix B. Fermi contact interaction

The matrix element of the Fermi contact interaction for the XY and UY processes is considered. In the XY process, the intermediate and final states are electron triplet $({}^{3}\Sigma_{u}^{+})$ and singlet $({}^{1}\Sigma_{g}^{+})$ states, which correspond to $(1s\sigma_{g})^{1}(2p\sigma_{u})^{1}$ and $(1s\sigma_{g})^{2}$ electron configurations, respectively. The orbital part of the wavefunction is described as

$$|^{1}\Sigma_{g}^{+}\rangle = \chi_{e}(S=0)\phi_{g}(\vec{r}_{1})\phi_{g}(\vec{r}_{2})$$

$$|^{3}\Sigma_{u}^{+}\rangle = \chi_{e}(S=1)\frac{1}{\sqrt{2}}(\phi_{g}(\vec{r}_{1})\phi_{u}(\vec{r}_{2}) - \phi_{g}(\vec{r}_{2})\phi_{u}(\vec{r}_{1})), \qquad (B.1)$$

where ϕ_g and ϕ_u represent the orbitals of $1s\sigma_g$ and $2p\sigma_u$, respectively, and $\chi_e(S=1)$ and $\chi_e(S=0)$ denote the electron spin triplet and singlet functions, respectively.

Since the H_2 molecule has two electrons, the orbital integral must be done for the two electron position vectors.

$$\langle {}^{3}\Sigma_{u}^{+}|H_{FC}|{}^{1}\Sigma_{g}^{+}\rangle = \langle \frac{1}{\sqrt{2}}(\phi_{g}(\vec{r}_{1})\phi_{u}(\vec{r}_{2}) - \phi_{g}(\vec{r}_{2})\phi_{u}(\vec{r}_{1}))| \\ \sum_{j=1,2} \sum_{K=a,b} \vec{s}_{j} \cdot \vec{i}_{K}\delta(\vec{r}_{j} - \vec{R}_{K})|\phi_{g}(\vec{r}_{1})\phi_{g}(\vec{r}_{2})\rangle$$
(B.2)

Among the four terms of the Fermi contact Hamiltonian, integration of $\vec{s}_1 \cdot \vec{i}_a \delta(\vec{r}_1 - \vec{R}_a)$ over \vec{r}_1 and \vec{r}_2 leads to λ_{IFC} by using the relation of $\langle \phi_g(\vec{r}) | \phi_u(\vec{r}) \rangle = 0$ as

$$\lambda_{IFC} = -\int \phi_g(\vec{r})\delta(\vec{r} - \vec{R_a})\phi_u(\vec{r})d\vec{r}$$

= $-\phi_u(\vec{R_a})\phi_g(\vec{R_a})$ (B.3)

²²¹¹ On the other hand, integration of the term $\vec{s}_1 \cdot \vec{i}_b \delta(\vec{r}_1 - \vec{R}_b)$ results in $-\phi_u(\vec{R}_b)\phi_g(\vec{R}_b)$. ²²¹² While $\phi_g(\vec{R}_a) = \phi_g(\vec{R}_b)$ because of the gelade character, $\phi_u(\vec{R}_a) = -\phi_u(\vec{R}_b)$ ²²¹³ due to the ungelade nature. Therefore, the sum of the two terms leads to ²²¹⁴ $\lambda_{IFC}\vec{s}_1 \cdot (\vec{i}_a - \vec{i}_b)$. Similarly, $-\lambda_{IFC}\vec{s}_2 \cdot (\vec{i}_a - \vec{i}_b)$ is obtained by calculation of ²²¹⁵ the other two terms, hence Eq. 26 is obtained.

In the UY process, the intermediate and final states are electron triplet and singlet states with configurations of $(1s\sigma_u)^1(k)^1$ and $(k)^1(k')^1$, respectively, where k and k' denote the substrate states. The orbital part of the ²²¹⁹ wavefunction is described as

$$|{}^{1}I\rangle = \chi_{e}(S=0)\frac{1}{\sqrt{2}}(\phi_{k}(\vec{r_{1}})\phi_{u}(\vec{r_{2}}) + \phi_{k}(\vec{r_{2}})\phi_{u}(\vec{r_{1}}))$$
 (B.4)

$$|{}^{3}f\rangle = \chi_{e}(S=1)\frac{1}{\sqrt{2}}(\phi_{k}(\vec{r_{1}})\phi_{k'}(\vec{r_{2}}) - \phi_{k}(\vec{r_{2}})\phi_{k'}(\vec{r_{1}})).$$
(B.5)

Although integration of the Fermi contact Hamiltonian is more complicated than the intramolecular case, the surface-molecule Fermi contact coefficient λ_{SFC} proportional to the $(\vec{s_1} - \vec{s_2}) \cdot (\vec{i_a} - \vec{i_b})$ term is given by

$$\lambda_{SFC} = -\phi_u^*(\vec{R}_a)(\phi_{k'}(\vec{R}_a) + \phi_{k'}(\vec{R}_b))/2 + \phi_u^*(\vec{R}_a)(\phi_k(\vec{R}_a) + \phi_k(\vec{R}_b))/2 \int \phi_k^*(\vec{r})\phi_{k'}(\vec{r})d\vec{r} + (|\phi_k(\vec{R}_a)|^2 - |\phi_k(\vec{R}_b)|^2)/2 \int \phi_u^*(\vec{r})\phi_{k'}(\vec{r})d\vec{r} - (\phi_k^*(\vec{R}_a)\phi_{k'}(\vec{R}_a) - \phi_k^*(\vec{R}_b)\phi_{k'}(\vec{R}_b))/2 \int \phi_u^*(\vec{r})\phi_k(\vec{r})d\vec{r}.$$
 (B.6)

When this term is integrated over the rotational wavefunctions, only the first and second terms in B.6 are significant by using the relation of $\langle Y_{J,m} | \phi_k(\vec{R}_a) | Y_{J,m} \rangle =$ $\langle Y_{J,m} | \phi_k(\vec{R}_b) | Y_{J,m} \rangle$ and $\langle Y_{J,m} | \phi_{k'}(\vec{R}_a) | Y_{J,m} \rangle = \langle Y_{J,m} | \phi_{k'}(\vec{R}_b) | Y_{J,m} \rangle.$

2226 Appendix C. Spin-orbit interaction

²²²⁷ When the two electrons in H₂ are in the $(1s\sigma_g)^1(nl\pi)^1$ configuration, the ²²²⁸ resulting state can be either a singlet $(|^1\Pi\rangle)$ or triplet $(|^3\Pi\rangle)$ state, of which ²²²⁹ orbital part of the wavefunctions are described in a first approximation as,

$$|^{1}\Pi\rangle = \frac{1}{\sqrt{2}}(\phi_{\sigma}(\vec{r}_{1})\phi_{\pi}(\vec{r}_{2}) + \phi_{\sigma}(\vec{r}_{2})\phi_{\pi}(\vec{r}_{1}))$$
$$|^{3}\Pi\rangle = \frac{1}{\sqrt{2}}(\phi_{\sigma}(\vec{r}_{1})\phi_{\pi}(\vec{r}_{2}) - \phi_{\sigma}(\vec{r}_{2})\phi_{\pi}(\vec{r}_{1}))$$
(C.1)

(C.2)

where $\phi_{\sigma}(\vec{r})$ and $\phi_{\pi}(\vec{r})$ represent the orbitals of $1s\sigma_g$ and $nl\pi$. Since $\phi_{\sigma}(\vec{r})$ and $\phi_{\pi}(\vec{r})$ have the axial component of the orbital angular momentum of 0 and 1, $\langle \phi_{\sigma} | \ell_z | \phi_{\sigma} \rangle = 0$ and $\langle \phi_{\pi} | \ell_z | \phi_{\pi} \rangle = 1$, which leads to

$$\langle {}^{3}\Pi | \ell_{1}^{z} s_{1}^{z} + \ell_{2}^{z} s_{2}^{z} | {}^{1}\Pi \rangle = s_{2}^{z} - s_{1}^{z}$$
(C.3)

(C.4)

 $_{\rm 2233}$ As described in Sec. 3.2,

$$\langle \chi_e(S=1, S_z=0) | s_2^z - s_1^z | \chi_e(S=0) \rangle = 1$$
 (C.5)

 $_{2234}$ $\,$ hence, the spin-orbit coupling mixes the electron singlet and triplet states.

2235 **References**

- [1] A. Farkas, Orthohydrogen, Parahydrogen, and Heavy Hydrogen, Cambridge University Press, Cambridge, 1935.
- [2] G. B. Field, W. B. Somerville, K. Dressler, Annu. Rev. Astro. Astrophys. 4 (1966) 207.
- ²²⁴⁰ [3] I. F. Silvera, Rev. Mod. Phys. 52 (1980) 39.
- ²²⁴¹ [4] D. M. Dennison, Proc. Roy. Soc. (London) 115 (1927) 483.
- ²²⁴² [5] F. Hund, Z. Phys. 40 (1927) 93.
- ²²⁴³ [6] W. Heisenberg, Z. Phys. 41 (1927) 239.
- ²²⁴⁴ [7] T. Hori, Z. Phys. 44 (1929) 834.
- ²²⁴⁵ [8] K. F. Bonhoeffer, P. Harteck, Naturewiss. 17 (1929) 182.
- [9] A. H. Larsen, F. E. Simon, C. A. Swenson, Rev. Sci. Instr. 19 (1948)
 2247 266.
- ²²⁴⁸ [10] E. Ilisca, Prog. Surf. Sci. 41 (1992) 217.
- ²²⁴⁹ [11] Y. L. Sandler, J. Phys. Chem. 58 (1954) 58.
- ²²⁵⁰ [12] Y. L. Sandler, J. Chem. Phys. 29 (1958) 97.
- [13] C. M. Cunningham, D. S. Chapin, H. L. Johnston, J. Am. Chem. Soc.
 80 (1958) 2382.
- ²²⁵³ [14] Y. Ishii, Prog. Surf. Sci. 21 (1986) 163.
- ²²⁵⁴ [15] E. Ilisca, S. Paris, Phys. Rev. Lett. 82 (1999) 1788.
- ²²⁵⁵ [16] S. Paris, E. Ilisca, J. Phys. Chem. A 103 (1999) 4964.
- [17] P. Avouris, D. Schmeisser, J. E. Demuth, Phys. Rev. Lett. 48 (1982)
 199.
- ²²⁵⁸ [18] S. Andersson, J. Harris, Phys. Rev. Lett. 48 (1982) 545.
- ²²⁵⁹ [19] R. Palmer, R. Willis, Surf. Sci. 179 (1987) L1.
- ²²⁶⁰ [20] M. Sakurai, T. Okano, Y. Tuzi, Appl. Surf. Sci. 33/34 (1988) 245.
- ²²⁶¹ [21] S. Yucel, Phys. Rev. B 39 (1989) 3104.
- ²²⁶² [22] E. Ilisca, J. Phys. I France 1 (1991) 1785.
- ²²⁶³ [23] H. Kasai, W. Diño, R. Muhida, Prog. Surf. Sci. 72 (2003) 53.
- ²²⁶⁴ [24] T. Magome, K. Fukutani, T. Okano, Shinku 42 (1999) 286.
- [25] K. Fukutani, K. Yoshida, M. Wilde, W. Diño, M. Matsumoto,
 T. Okano, Phys. Rev. Lett. 90 (2003) 096103.
- [26] L. Amiaud, A. Momeni, F. Dulieu, J.-H. Fillion, E. Matar, J. L.
 Lemaire, Phys. Rev. Lett. 100 (2008) 056101.
- [27] F. Combes, G. P. des Forêts (Eds.), Molecular hydrogen in space, Cambridge University Press, 2000.
- [28] E. Habart, M. Walmsleya, L. Verstraete, S. Cazaus, R. Maiolino,
 P. Cox, F. Boulanger, G. P. des Forêts, Space Sci. Rev. 119 (2005)
 71.
- ²²⁷⁴ [29] J. L. Bourlot, Astron. Astrophys. 360 (2000) 656.
- ²²⁷⁵ [30] W. F. Thi, Nature 409 (2001) 60.
- ²²⁷⁶ [31] T. Fouchet, E. Lellouch, H. Feuchtgruber, Icarus 161 (2003) 127.
- [32] L. Fletcher, P. Irwin, N. Teanby, G. Orton, P. Parrish, R. de Koka,
 C. Howett, S. Calcutt, N. Bowles, F. Taylor, Icarus 189 (2007) 457.
- ²²⁷⁹ [33] T. Sugimoto, K. Fukutani, Nature Phys. 7 (2011) 307.
- ²²⁸⁰ [34] T. E. Sharp, At. Data 2 (1971) 119.
- ²²⁸¹ [35] L. Wolniewicz, K. Dressler, J. Chem. Phys. 82 (1985) 3292.
- ²²⁸² [36] L. Wolniewicz, K. Dressler, J. Chem. Phys. 88 (1988) 3861.
- [37] K. Niki, M. Fujiwara, Y. Motoshima, T. Kawauchi, K. Fukutani, Chem.
 Phys. Lett. 504 (2011) 136.

- [38] G. Herzberg, Molecular spectra and molecular structure, I. Spectra of diatomic molecules, 2nd ed., Krieger, Florida, 1950.
- [39] K. P. Huber, G. Herzberg, Molecular spectra and molecular structure,
 IV. Constants of Diatomic Molecules, Van Nostrand Reinhold, New
 York, 1979.
- ²²⁹⁰ [40] K. Christmann, Surf. Sci. Rep. 9 (1988) 1.
- ²²⁹¹ [41] G. Vidali, G. Ihm, H.-Y. Kim, M. W. Cole, Surf. Sci. Rep. 12 (1991) ²²⁹² 133.
- ²²⁹³ [42] Y. G. Ptushinskii, Low Temp. Phys. 30 (1991) 1.
- [43] L. W. Bruch, R. D. Diehl, J. A. Venables, Rev. Mod. Phys. 79 (2007)
 1381.
- ²²⁹⁶ [44] K. Svensson, S. Andersson, Phys. Rev. Lett. 78 (1997) 2016.
- ²²⁹⁷ [45] P. Nordlander, C. Holmberg, J. Harris, Surf. Sci. 152 (1985) 702.
- ²²⁹⁸ [46] P. Nordlander, C. Holmberg, J. Harris, Surf. Sci. 175 (1986) L753.
- ²²⁹⁹ [47] H. Hoinkes, Rev. Mod. Phys. 52 (1980) 933.
- ²³⁰⁰ [48] J. Perreau, J. Lapujoulade, Surf. Sci. 122 (1982) 341.
- [49] C.-F. Yu, K. B. Whaley, C. S. Hogg, S. J. Sibener, Phys. Rev. Lett. 51 (1983) 2210.
- [50] C.-F. Yu, K. B. Whaley, C. S. Hogg, S. J. Sibener, J. Chem. Phys. 83
 (1985) 4217.
- ²³⁰⁵ [51] U. Harten, J. P. Toennies, C. Woll, J. Chem. Phys. 85 (1986) 2249.
- ²³⁰⁶ [52] S. Andersson, L. Wilzen, Phys. Rev. B 38 (1988) 2967.
- [53] S. Andersson, L. Wilzen, M. Persson, J. Harris, Phys. Rev. B 40 (1989)
 8146.
- ²³⁰⁹ [54] S. Andersson, M. Persson, Phys. Rev. B 48 (1993) 5685.
- [55] L. Amiaud, J.-H. Fillion, S. Baouche, F. Dulieu, A. Momeni, J. L.
 Lemaire, J. Chem. Phys. 124 (2006) 094702.

- [56] L. Amiaud, F. Dulieu, J.-H. Fillion, A. Momeni, J. L. Lemaire, J.
 Chem. Phys. 127 (2007) 144709.
- [57] S. Iwata, Y. Sato, K. Nakai, S. Ogura, T. Okano, M. Namura, A. Kasuya, K. Tohji, K. Fukutani, J. Phys. Chem. C 111 (2007) 14937.
- [58] A.-S. Mårtensson, C. Nyberg, S. Andersson, Phys. Rev. Lett. 57 (1986)
 2045.
- [59] K. Svensson, L. Bengtsson, J. Bellman, M. Hassel, M. Persson, S. Andersson, Phys. Rev. Lett. 83 (1999) 124.
- [60] L. Bengtsson, K. Svensson, M. Hassel, J. Bellman, M. Persson, S. Andersson, Phys. Rev. B 61 (2000) 16921.
- ²³²² [61] P. K. Schmidt, K. Christmann, G. Kresse, J. Hafner, M. Lischka,
 A. Groß, Phys. Rev. Lett. 87 (2001) 096103.
- [62] J. Wang, C. Y. Fan, Q. Sun, K. Reuter, K. Jacobi, M. Scheffler, G. Ertl,
 Ang. Chem. Int. Ed. 42 (2003) 2151.
- [63] K. Svensson, J. Bellman, A. Hellman, S. Andersson, Phys. Rev. B 71 (2005) 245402.
- [64] J. Bellman, K. Svensson, S. Andersson, J. Chem. Phys. 125 (2006)
 064704.
- [65] L. Kong, G. Román-Pérez, J. M. Soler, D. C. Langreth, Phys. Rev.
 Lett. 103 (2009) 096103.
- [66] K. Lee, A. K. Kelkkanen, K. Berland, S. Andersson, D. Langreth,
 E. Schroder, B. I. Lundqvist, P. Hyldgaard, Phys. Rev. B 84 (2011)
 193408.
- ²³³⁵ [67] J. Harris, P. Feibelman, Surf. Sci. 115 (1982) L133.
- ²³³⁶ [68] R. Zare, Angular Momentum, Wiley, New York, 1988.
- ²³³⁷ [69] D. White, E. N. Lasserttre, J. Chem. Phys. 32 (1960) 72.
- [70] M. Chiesa, L. Mattera, R. Musenich, C. Salvo, Surf. Sci. 151 (1985)
 L145.

- ²³⁴⁰ [71] R. Schinke, V. Engel, H. Voges, Chem. Phys. Lett. 104 (1984) 279.
- [72] L. Wilzen, F. Althoff, S. Andersson, M. Persson, Phys. Rev. B 43 (1991)
 7003.
- ²³⁴³ [73] K. Svensson, S. Andersson, Surf. Sci. 392 (1997) L40.
- ²³⁴⁴ [74] G. Cilpa, G. Chambaud, Surf. Sci. 601 (2007) 320.
- ²³⁴⁵ [75] G. Cilpa, M. Guitou, G. Chambaud, Surf. Sci. 602 (2008) 2894.
- ²³⁴⁶ [76] I. F. Silvera, M. Nielsen, Phys. Rev. Lett. 37 (1976) 1275.
- ²³⁴⁷ [77] A. J. Ramirez-Cuesta, P. C. H. Mitchell, Catal. Today 120 (2007) 368.
- [78] J. Z. Larese, T. Arnold, L. Frazier, R. J. Hinde, A. J. Ramirez-Cuesta,
 Phys. Rev. Lett. 101 (2008) 165302.
- [79] F. G. Brickwedde, R. B. Scott, H. S. Taylor, J. Phys. Chem. 3 (1935)
 653.
- [80] A. van Itterbeek, R. Hellemans, W. van Dael, Physica 30 (1964) 324.
- [81] H. Imao, K. Ishida, N. Kawamura, T. Matsuzaki, Y. Matsuda, A. Toyoda, P. Strasser, M. Iwasaki, K. Nagamine, Rev. Sci. Instrum. 79 (2008)
 053502.
- ²³⁵⁶ [82] A. A. Evett, J. Chem. Phys. 31 (1959) 565.
- ²³⁵⁷ [83] A. A. Evett, J. Chem. Phys. 33 (1960) 789.
- ²³⁵⁸ [84] A. Katorski, D. White, J. Chem. Phys. 40 (1964) 3183.
- ²³⁵⁹ [85] W. R. Moore, H. R. Ward, J. Am. Chem. Soc. 80 (1958) 2909.
- ²³⁶⁰ [86] P. L. Gant, K. Yang, Science 129 (1958) 1548.
- [87] W. J. Haubach, C. M. Knobler, A. Katorski, D. White, J. Phys. Chem.
 71 (1967) 1398.
- ²³⁶³ [88] J. King, S. W. Benson, J. Chem. Phys. 44 (1966) 1007.
- ²³⁶⁴ [89] E. M. Mortensen, H. Eyring, J. Phys. Chem. 64 (1960) 433.

- ²³⁶⁵ [90] T. B. MacRury, J. R. Sams, Mol. Phys. 19 (1970) 337.
- ²³⁶⁶ [91] T. B. MacRury, J. R. Sams, Mol. Phys. 19 (1970) 353.
- ²³⁶⁷ [92] T. B. MacRury, J. R. Sams, Mol. Phys. 20 (1971) 57.
- ²³⁶⁸ [93] E. L. Pace, A. R. Siebert, J. Phys. Chem. 63 (1959) 1398.
- [94] J. Heidberg, N. Gushanskaya, O. Schonekas, R. Schwarte, Surf. Sci.
 331-333 (1995) 1473.
- ²³⁷¹ [95] T. Sugimoto, K. Fukutani, in preparation.
- ²³⁷² [96] T. Sugimoto, K. Fukutani, submitted 00 (2012) 00.
- 2373 [97] A. Schutte, D. Bassi, F. Tommasini, A. Turelli, J. Chem. Phys. 64
 2374 (1976) 4135.
- ²³⁷⁵ [98] J. C. Tully, Annu. Rev. Phys. Chem. 51 (2000) 153.
- ²³⁷⁶ [99] S. Andersson, J. Harris, Phys. Rev. B 27 (1983) 9.
- ²³⁷⁷ [100] S. Andersson, L. Wilzen, J. Harris, Phys. Rev. Lett. 55 (1985) 2591.
- ²³⁷⁸ [101] S. Andersson, L. Wilzen, J. Harris, Phys. Rev. Lett. 57 (1986) 1603.
- ²³⁷⁹ [102] G. Armand, J. R. Manson, Phys. Rev. B 43 (1991) 14371.
- ²³⁸⁰ [103] A. E. DePristo, C.-Y. LEE, J. M. Hutson, Surf. Sci. 169 (1986) 451.
- ²³⁸¹ [104] M. D. Stiles, J. W. Wilkins, Phys. Rev. Lett. 54 (1985) 595.
- ²³⁸² [105] B. Gumhalter, Phys. Rep. 351 (2001) 1.
- ²³⁸³ [106] A. Sanz, S. Miret-Artés, Phys. Rep. 451 (2007) 37.
- [107] R. F. C. Jr., J. V. V. Kasper, K. S. Pitzer, J. Chem. Phys. 46 (1967)
 3220.
- [108] P. L. Chapovsky, L. J. F. Hermans, Annu. Rev. Phys. Chem. 50 (1999)
 315.
- ²³⁸⁸ [109] P. L. Chapovsky, Phys. Rev. A 43 (1999) 3624.

- ²³⁸⁹ [110] Z.-D. Sun, K. Takagi, F. Matsushima, Science 310 (2005) 1938.
- ²³⁹⁰ [111] C. Bechtela, E. Eliasb, B. F. Schramm, J. Mol. Struct. 741 (2005) 97.
- [112] Y. Miyamoto, M. Fushitani, D. Ando, T. Momose, J. Chem. Phys. 128
 (2008) 114502.
- [113] K. Tanaka, M. Hayashi, M. Ohtsuki, K. Harada, T. Tanaka, J. Chem.
 Phys. 131 (2009) 111101.
- ²³⁹⁵ [114] P. Cacciani, J. Cosleou, M. Khelkhal, Phys. Rev. A 85 (2012) 012521.
- ²³⁹⁶ [115] K. F. Bonhoeffer, P. Harteck, Z. Phys. Chem. B 4 (1929) 113.
- ²³⁹⁷ [116] D. E. Osterbrock, Astrophys. J. 136 (1962) 359.
- [117] H. A. Bethe, E. E. Salpeter, Quantum Mechanics of One- and Two Electron Atoms, Dover, New York, 2008.
- ²⁴⁰⁰ [118] J. C. Raich, J. R. H. Good, Astrophys. J. 139 (1964) 1004.
- ²⁴⁰¹ [119] S. Dodelson, J. Phys. B 19 (1986) 2871.
- ²⁴⁰² [120] Y. Sun, A. Dalgarno, Astrophys. J. 427 (1994) 1053.
- ²⁴⁰³ [121] K. Pachucki, J. Komasa, Phys. Rev. A 77 (2008) 030501(R).
- ²⁴⁰⁴ [122] D. L. Huestis, Planetary and Space Science 56 (2008) 1733.
- ²⁴⁰⁵ [123] D. G. Truhlar, R. E. Wyatt, Ann. Rev. Phys. Chem. 27 (1976) 1.
- ²⁴⁰⁶ [124] A. Dalgarno, J. H. Black, J. C. Weisheit, Astrophys. J. 14 (1973) L77.
- [125] F. C. Fehsenfeld, D. B. Dunkin, E. E. Ferguson, D. L. Albritton, Astrophys. J. 183 (1973) L25.
- ²⁴⁰⁹ [126] D. Smith, N. G. Adams, E. Alge, Astrophys. J. 263 (1982) 123.
- ²⁴¹⁰ [127] D. Gerlich, J. Chem. Phys. 92 (1990) 2377.
- [128] P. Honvault, M. Jorfi, T. González-Lezana, A. Faure, L. Pagani, Phys.
 Rev. Lett. 107 (2011) 023201.

- [129] P. Honvault, M. Jorfi, T. González-Lezana, A. Faure, L. Pagani, Phys.
 Chem. Chem. Phys. 13 (2011) 19089.
- 2415 [130] D. R. Flower, G. P. des Forêts, C. M. Walmsley, Astron. Astrophys.
 2416 449 (2006) 621.
- [131] A. I. Boothroyd, P. G. Martin, W. J. Keogh, M. J. Peterson, J. Chem.
 Phys. 116 (2002) 666.
- ²⁴¹⁹ [132] K. G. Petzinger, D. J. Scalapino, Phys. Rev. B 8 (1973) 266.
- [133] K. Makoshi, M. Rami, E. Ilisca, J. Phys. Condens. Matter 5 (1993)
 7325.
- ²⁴²² [134] M. Rami, K. Makoshi, E. Ilisca, Surf. Sci. 68 (1993) 197.
- ²⁴²³ [135] E. Ilisca, Chem. Phys. Lett. 168 (1990) 289.
- [136] L. D. Landau, E. M. Lifshitz, Quantum Mechanics- Nonrelativistic
 Theory, 2nd Edition, Pergamon, Oxford, 1965.
- ²⁴²⁶ [137] M. Hamermesh, Phys. Rev. 73 (1948) 638.
- ²⁴²⁷ [138] Y. L. Sandler, J. Chem. Phys. 20 (1952) 1050.
- ²⁴²⁸ [139] K. Motizuki, J. Phys. Soc. Jpn. 12 (1957) 163.
- ²⁴²⁹ [140] E. Wigner, Z. Phys. Chem. B 23 (1933) 28.
- ²⁴³⁰ [141] L. G. Harrison, C. A. McDowell, Proc. Roy. Soc. A 220 (1953) 77.
- ²⁴³¹ [142] A. J. Leffler, J. Chem. Phys. 43 (1965) 4410.
- [143] In ref. [142], it is pointed out that the indistinguishability principle
 of identical particles might be broken under spatially inhomogeneous
 potential. As stated in Sec. 2.2, however, the rotational motion is modified in an inisotropic potential with the eigen state described by linear
 combination of other rotational states, and the symmetry of the rotational wavefunction is strictly maintained with respect to the particle
 exchange.
- ²⁴³⁹ [144] C. F. Ng, Surf. Sci. 79 (1979) 470.

- ²⁴⁴⁰ [145] E. Ilisca, Phys. Rev. Lett. 24 (1970) 797.
- ²⁴⁴¹ [146] E. Ilisca, A. P. Legrand, Phys. Rev. B 5 (1972) 4994.
- ²⁴⁴² [147] Y. Ishii, S. Sugano, Surf. Sci. 127 (1983) 21.
- ²⁴⁴³ [148] Y. Ishii, S. Sugano, J. Phys. C. Solid State Phys. 17 (1984) 2094.
- ²⁴⁴⁴ [149] K. Motizuki, T. Nagamiya, J. Phys. Soc. Jpn. 11 (1956) 93.
- ²⁴⁴⁵ [150] F. Schmidt, Phys. Rev. B 10 (1974) 4480.
- ²⁴⁴⁶ [151] Y. Y. Milenko, R. M. Sibileva, Sov. J. Low Temp. Phys. 1 (1975) 382.
- [152] A. Honig, M. Lewis, Z.-Z. Yu, S. Yucel, Phys. Rev. Lett. 56 (1986)
 1866.
- [153] P. Pedroni, H. Meyer, F. Weinhaus, D. Haase, Solid State Commun.
 14 (1974) 279.
- ²⁴⁵¹ [154] A. Driessen, E. van der Poll, I. F. Silvera, Phys. Rev. B 30 (1984) 2517.
- ²⁴⁵² [155] M. G. Pravica, I. F. Silvera, Phys. Rev. Lett. 81 (1998) 4180.
- ²⁴⁵³ [156] M. A. Strzhemechny, R. J. Hemley, Phys. Rev. Lett. 85 (2000) 05595.
- [157] M. A. Strzhemechny, R. J. Hemley, H. k. Mao, A. F. Goncharov, J. H.
 Eggert, Phys. Rev. B 66 (2002) 014103.
- ²⁴⁵⁶ [158] A. J. Berlinsky, W. N. Hardy, Phys. Rev. B 8 (1973) 5013.
- [159] Y. Li, X. Lei, R. G. Lawler, Y. Murata, K. Komatsu, N. J. Turro, J. Phys. Chem. Lett. 2 (2011) 741.
- ²⁴⁵⁹ [160] E. Ilisca, S. Sugano, Phys. Rev. Lett. 57 (1986) 2590.
- ²⁴⁶⁰ [161] E. Ilisca, S. Paris, Surf. Sci. 363 (1996) 347.
- ²⁴⁶¹ [162] E. Ilisca, Phys. Rev. Lett. 66 (1991) 667.
- ²⁴⁶² [163] E. Ilisca, Surf. Sci. 242 (19916) 470.
- ²⁴⁶³ [164] E. Ilisca, Opt. Commun. 89 (1992) 399.

- [165] R. Muhida, H. Setiyanto, M. M. Rahman, W. Diño, H. Nakanishi,
 H. Kasai, K. Fukutani, T. Okano, Eur. Phys. J. D 38 (2006) 99.
- [166] R. Muhida, M. David, M. M. Rahman, W. Diño, H. Nakanishi, H. Ka sai, K. Fukutani, T. Okano, Eur. Phys. J. D 38 (2006) 99.
- [167] R. Muhida, W. Diño, A. Fukui, H. Kasai, H. Nakanishi, A. Okiji,
 K. Fukutani, T. Okano, Surf. Sci. 493 (2001) 285.
- ²⁴⁷⁰ [168] R. Muhida, W. Diño, A. Fukui, H. Kasai, H. Nakanishi, A. Okiji,
 ²⁴⁷¹ K. Fukutani, T. Okano, J. Phys. Soc. Jpn. 70 (2001) 3654.
- ²⁴⁷² [169] R. Muhida, W. Diño, Y. Miura, H. Kasai, H. Nakanishi, A. Okiji,
 ²⁴⁷³ K. Fukutani, T. Okano, Surf. Sci. 514 (2002) 273.
- ²⁴⁷⁴ [170] R. Muhida, W. Diño, Y. Miura, H. Kasai, H. Nakanishi, K. Fukutani,
 ²⁴⁷⁵ T. Okano, A. Okiji, J. Vac. Soc. Jpn. 45 (2002) 448.
- ²⁴⁷⁶ [171] R. Muhida, Y. Miura, W. Diño, H. Kasai, H. Nakanishi, A. Okiji,
 ²⁴⁷⁷ K. Fukutani, T. Okano, J. Appl. Phys. 93 (2003) 644.
- ²⁴⁷⁸ [172] Y. Kunisada, H. Kasai, J. Phys. Soc. Jpn. 82 (2013) 023601.
- [173] Y. Kunisada, H. Nakanishi, W. A. Diño, H. Kasai, J. Vac. Soc. Jpn.
 55 (2012) 115.
- ²⁴⁸¹ [174] E. Ilisca, S. Sugano, Chem. Phys. Lett. 149 (1988) 20.
- ²⁴⁸² [175] L. Wolniewicz, J. Mol. Spec. 169 (1995) 329.
- ²⁴⁸³ [176] W. Kolos, J. Rychlewski, Comp. Methods Sci. Technol. 5 (1999) 39.
- [177] A. Abragam, The principles of nuclear magnetism, Oxford University
 Press, Oxford, 1961.
- ²⁴⁸⁶ [178] A. B. Harris, Phys. Rev. B 2 (1970) 3495.
- ²⁴⁸⁷ [179] N. S. Sullivan, R. V. Pound, Phys. Rev. A 6 (1972) 1102.
- ²⁴⁸⁸ [180] E. R. Grilly, Rev. Sci. Instrum. 24 (1952) 72.
- ²⁴⁸⁹ [181] T. W. Bradshaw, J. O. W. Norris, Rev. Sci. Instrum. 58 (1986) 83.

- ²⁴⁹⁰ [182] M. Nielsen, Phys. Rev. B 7 (1973) 1626.
- [183] D. W. Brown, P. E. Sokol, S. A. FitzGerald, Phys. Rev. B 59 (1999)
 13258.
- [184] S. A. FitzGerald, T. Yildirim, L. J. Santodonato, D. A. Neumann,
 J. R. D. Copley, J. J. Rush, F. Trouw, Chem. Phys. Lett. 329 (2000)
 311.
- [185] C. M. Brown, T. Yildirim, D. A. Neumann, M. J. Heben, T. Gennett,
 A. C. Dillon, J. L. Alleman, J. E. Fischer, Chem. Phys. Lett. 329 (2000)
 311.
- [186] A. J. Horsewill, K. S. Panesar, S. Rols, M. R. Johnson, Y. Murata, K. Komatsu, S. Mamone, A. Danquigny, F. Cuda, S. Maltsev, M. C. Grossel, M. Carravetta, M. H. Levitt, Phys. Rev. Lett. 102 (2009) 013001.
- [187] K. Sumida, S. Horike, S. S. Kaye, Z. R. Herm, W. L. Queen, C. M.
 Brown, F. Grandjean, G. J. Long, A. Dailly, J. R. Long, Chem. Sci. 1
 (2010) 184.
- ²⁵⁰⁶ [188] M. Hiller, E. V. Lavrov, J. Weber, Phys. Rev. Lett. 98 (2007) 055504.
- ²⁵⁰⁷ [189] M. Hiller, E. V. Lavrov, J. Weber, Phys. Rev. B 80 (2009) 045306.
- [190] V. Plotnichenko, S. Vasiliev, A. Rybaltovskii, V. Koltashev, V. Sokolov,
 S. Klyamkin, O. Medvedkov, A. Rybaltovskii, A. Malosiev, E. Dianov,
 J. Non-Cryst. Solids 351 (2005) 3677.
- [191] K. A. Williams, B. K. Pradhan, P. C. Eklund, M. K. Kostov, M. W.
 Cole, Phys. Rev. Lett. 88 (2002) 165502.
- ²⁵¹³ [192] A. Centrone, L. Brambilla, G. Zerbi, Phys. Rev. B 71 (2005) 245406.
- ²⁵¹⁴ [193] A. Centrone, J. Raman Spectrosc. 42 (2011) 2165.
- ²⁵¹⁵ [194] G. J. Schulz, Rev. Mod. Phys. 45 (1973) 423.
- [195] J. E. Demuth, D. Schmeisser, P. Avouris, Phys. Rev. Lett. 47 (1981)
 1166.

- ²⁵¹⁸ [196] M. Folman, Y. Kozirovski, J. Colloid Int. Sci. 38 (1972) 51.
- ²⁵¹⁹ [197] T. E. Huber, C. A. Huber, Phys. Rev. Lett. 59 (1987) 1120.
- [198] H. G. Hixson, M. J. Wojcik, M. S. Devlin, J. P. Devlin, V. Buch, J.
 Chem. Phys. 97 (1992) 753.
- ²⁵²² [199] V. Buch, J. P. Devlin, J. Chem. Phys. 98 (1993) 4195.
- ²⁵²³ [200] S. A. FitzGerald, S. Forth, M. Rinkoski, Phys. Rev. B 65 (2002) 140302.
- [201] S. A. FitzGerald, H. O. H. Churchill, P. M. Korngut, C. B. Simmons,
 Y. E. Strangas, Phys. Rev. B 73 (2006) 155409.
- [202] S. Mamone, M. Ge, D. Hüvonen, U. Nagel, A. Danquigny, F. Cuda,
 M. C. Grossel, Y. Murata, K. Komatsu, M. Levitt, T. Rõõm, M. Carravetta, J. Chem. Phys. 130 (2009) 081103.
- [203] C. Peng, M. Stavola, W. B. Fowler, M. Lockwood, Phys. Rev. B 80 (2009) 125207.
- [204] S. A. FitzGerald, J. Hopkins, B. Burkholder, M. Friedman, Phys. Rev. B 81 (2010) 104305.
- [205] S. A. FitzGerald, B. Burkholder, M. Friedman, J. B. Hopkins, C. J.
 Pierce, J. M. Schloss, B. Thompson, J. L. C. Rowsell, J. Am. Chem.
 Soc. 133 (2011) 20310.
- ²⁵³⁶ [206] L. Kong, Y. J. Chabal, D. C. Langreth, Phys. Rev. B 83 (2011) 121402.
- [207] G. Ricchiardi, J. G. Vitillo, D. Cocina, E. N. Gribovc, A. Zecchinaa,
 Phys. Chem. Chem. Phys. 9 (2007) 2753.
- ²⁵³⁹ [208] M. Grunwald, G. E. Ewing, J. Chem. Phys. 109 (1998) 4990.
- ²⁵⁴⁰ [209] J. Heidberg, A. Vosberg, M. Hustedt, M. Thomas, J. Chem. Phys. 110 (1999) 2566.
- ²⁵⁴² [210] E. Marinero, C. Rettner, R. Zare, Phys. Rev. Lett. 48 (1982) 1323.
- ²⁵⁴³ [211] E. Marinero, R. Vasudev, R. Zare, J. Chem. Phys. 78 (1983) 692.
- ²⁵⁴⁴ [212] K.-D. Rinnen, M. Buntine, D. Kliner, J. Chem. Phys. 95 (1991) 214.

- ²⁵⁴⁵ [213] H. Rottke, K. H. Welge, Chem. Phys. Lett. 99 (1983) 456.
- ²⁵⁴⁶ [214] W. Meier, H. Rottke, H. Zacharias, K. H. Welge, J. Chem. Phys. 83
 ²⁵⁴⁷ (1985) 4360.
- [215] F. J. Northrup, J. C. Polanyi, S. C. Wallace, J. M. Williamson, Chem.
 Phys. Lett. 105 (1984) 34.
- ²⁵⁵⁰ [216] D. J. Kligler, C. Rhodes, Phys. Rev. Lett. 40 (1978) 309.
- ²⁵⁵¹ [217] D. Kligler, J. Bokor, C. Rhodes, Phys. Rev. A 21 (1980) 607.
- ²⁵⁵² [218] H. Pummer, H. Egger, T. Luk, T. Srinivasan, C. Rhodes, Phys. Rev.
 ²⁵⁵³ A 28 (1983) 795.
- ²⁵⁵⁴ [219] Q. Wang, T. Sugimoto, K. Fukutani, Seisankenkyu 64 (2012) 911.
- ²⁵⁵⁵ [220] H. Zacharias, Appl. Phys. A 47 (1988) 37.
- ²⁵⁵⁶ [221] G. Pozgainer, L. Windholz, A. Winkler, Meas. Sci. Technol. 5 (1994)
 ²⁵⁵⁷ 947.
- [222] J. S. A. Perry, J. M. Gingell, K. A. Newson, J. To, N. Watanabe, S. D.
 Price, Meas. Sci. Technol. 13 (2002) 1414.
- [223] K. Fukutani, K. Niki, T. Ito, H. Tashiro, M. Matsumoto, M. Wilde,
 T. Okano, W. Diño, H. Kasai, Surf. Sci. 593 (2005) 229.
- [224] K. Niki, T. Kawauchi, M. Matsumoto, K. Fukutani, T. Okano, Phys.
 Rev. B 77 (2008) 201403(R).
- [225] K. Niki, S. Ogura, M. Matsumoto, T. Okano, K. Fukutani, Phys. Rev. B 79 (2009) 085408.
- [226] M. Chehrouri, J.-H. Fillion, H. Chaabouni, H. Mokrane, E. Congiu,
 F. Dulieu, E. Matar, X. Michaut, J. Lemaire, Phys. Chem. Chem.
 Phys. 13 (2011) 2172.
- [227] M. Fujiwara, K. Niki, T. Okano, K. Fukutani, J. Phys. Conf. Ser. 200
 (2010) 022038.
- ²⁵⁷¹ [228] K. Svensson, S. Andersson, Phys. Rev. Lett. 98 (2007) 096105.

- ²⁵⁷² [229] A. J. Berlinsky, Phys. Rev. B 26 (1982) 443.
- ²⁵⁷³ [230] C. Benvenuti, R. S. Calder, G. Passardi, J. Vac. Sci. Technol. 13 (1976)
 ²⁵⁷⁴ 1172.
- ²⁵⁷⁵ [231] H. Schlichting, D. Menzel, Rev. Sci. Instrum. 64 (1993) 2013.
- ²⁵⁷⁶ [232] T. Sharp, At. Data 2 (1971) 119.
- ²⁵⁷⁷ [233] H. B. Michaelson, J. Appl. Phys. 48 (1977) 4729.
- [234] M.-C. Desjonquères, D. Spanjaard, Concepts in Surface Physics, 2nd
 Edition, Springer, Berlin, 1996.
- ²⁵⁸⁰ [235] M. Gruyters, K. Jacobi, Chem. Phys. Lett. 225 (1994) 309.
- [236] W. Wurth, J. Stöhr, P. Feulner, X. Pan, K. R. Bauchspiess, Y. Baba,
 E. Hudel, G. Rocker, D. Menzel, Phys. Rev. Lett. 65 (1990) 2426.
- ²⁵⁸³ [237] T. Yokoyama, D. Arvanitis, T. Lederer, M. Tischer, L. Tröger,
 ²⁵⁸⁴ K. Baberschke, G. Comelli, Phys. Rev. B 48 (1993) 15405.
- [238] Y. Jiang, Y. N. Zhang, J. X. Cao, R. Q. Wu, W. Ho, Science 333 (2011)
 324.
- ²⁵⁸⁷ [239] F. Besenbacher, J. K. Nørskov, Prog. Surf. Sci. 44 (1993) 5.
- ²⁵⁸⁸ [240] Y. Kazama, M. Matsumoto, T. Sugimoto, T. Okano, K. Fukutani,
 ²⁵⁸⁹ Phys. Rev. B 84 (2011) 064128.
- ²⁵⁹⁰ [241] Y. Kunisada, H. Nakanishi, H. Kasai, J. Phys. Soc. Jpn. 80 (2011)
 ²⁵⁹¹ 084605.
- ²⁵⁹² [242] M. Kurahashi, X. Sun, Y. Yamauchi, Phys. Rev. B 86 (2012) 245421.
- ²⁵⁹³ [243] Y. Kunisada, M. C. Escaño, H. Kasai, J. Phys.: Condens. Matter 23
 (2011) 394207.
- ²⁵⁹⁵ [244] K. Motizuki, T. Nagamiya, J. Phys. Soc. Jpn. 11 (1956) 654.
- ²⁵⁹⁶ [245] V. Shevtsov, A. Scherbakov, P. Malmi, E. Ylinen, M. Punkkinen, J.
 ²⁵⁹⁷ Low Temp. Phys. 104 (1996) 211.

- ²⁵⁹⁸ [246] V. Shevtsov, P. Malmi, E. Ylinen, M. Punkkinen, J. Low Temp. Phys.
 ²⁵⁹⁹ 114 (1999) 431.
- ²⁶⁰⁰ [247] V. Shevtsov, E. Ylinen, P. Malmi, M. Punkkinen, Phys. Rev. B 62
 (2000) 12386.
- ²⁶⁰² [248] Y. Kunisada, H. Nakanishi, W. A. Diño, H. Kasai, Curr. Appl. Phys.
 ²⁶⁰³ 12 (2012) S115.
- ²⁶⁰⁴ [249] P. R. Kubik, W. N. Hardy, H. Glattli, Can. J. Phys. 63 (1985) 605.
- ²⁶⁰⁵ [250] H. Freimuth, H. Wiechert, Surf. Sci. 162 (1985) 432.
- ²⁶⁰⁶ [251] H. Freimuth, H. Wiechert, Surf. Sci. 178 (1985) 716.
- ²⁶⁰⁷ [252] S. Yucel, N. Alexander, A. Honig, Phys. Rev. B 42 (1990) 820.
- [253] M. Fujita, K. Wakabayashi, K. Nakada, K. Kusakabe, J. Phys. Soc.
 Jpn. 65 (1996) 1920.
- [254] e.g. J. Cervenka, M. I. Katsnelson, C. F. J. Flipse, Nat. Phys. 5 (2009)
 840.
- ²⁶¹² [255] A. K. Geim, K. S. Novoselov, Nature Mater. 6 (2007) 183.
- ²⁶¹³ [256] Y. L. Sandler, J. Phys. Chem. 58 (1953) 54.
- [257] N. Watanabe, Y. Kimura, A. Kouchi, T. Chigai, T. Hama, V. Pir ronello, Astrophys. J. Lett. 714 (2010) 233.
- ²⁶¹⁶ [258] H. Mokrane, H. Chaabouni, M. Accolla, E. Congiu, F. Dulieu,
 ²⁶¹⁷ M. Chehrouri, J. L. Lemaire, Astrophys. J. 705 (2009) L195.
- ²⁶¹⁸ [259] D. J. Adams, Nature 293 (1981) 447.
- [260] E. R. Batista, S. S. Xantheas, H. Jonsson, J. Chem. Phys. 109 (1998)
 4546.
- ²⁶²¹ [261] C. Toubin, S. Picaud, C. Girardet, Chem. Phys. 244 (1999) 227.
- ²⁶²² [262] C. Girardet, C. Toubin, Surf. Sci. Rep. 44 (2001) 159.
- ²⁶²³ [263] T. Sugimoto, K. Fukutani, in preparation.

- ²⁶²⁴ [264] L. Senadheera, M. S. Conradi, J. Phys. Chem. A 112 (2008) 8303.
- ²⁶²⁵ [265] A. Giannasi, M. Celli, L. Ulivi, M. Zoppi, J. Chem. Phys. 129 (2008)
 084705.
- ²⁶²⁷ [266] M. Hiller, E. V. Lavrov, J. Weber, Phys. Rev. Lett. 99 (2007) 209901.
- [267] S. A. FitzGerald, K. Allen, P. Landerman, J. Hopkins, J. Matters,
 R. Myers, Phys. Rev. B 77 (2008) 224301.
- [263] J.-H. Fillion, L. Amiaud, E. Congiu, F. Dulieu, A. Momeni, J. L.
 Lemaire, Phys. Chem. Chem. Phys. 11 (2009) 4396.
- ²⁶³² [269] E. Habart, A. Abergel, F. Boulanger, C. Joblin, L. Verstraete,
 M. Compiègne, G. P. des Forêts, J. L. Bourlot, Astron. Astrophys.
 ²⁶³⁴ 527 (2011) A122.
- ²⁶³⁵ [270] J. Takahashi, Astrophys. J. 561 (2001) 254.
- ²⁶³⁶ [271] J. H. Black, E. F. van Dishoeck, Astrophys. J. 322 (1987) 412.
- ²⁶³⁷ [272] C. G. Gray, Can. J. Phys. 46 (1968) 135.
- [273] T. Hasegawa, I. Gatley, R. P. Garden, P. W. J. Brand, M. Ohishi,
 M. Hayashi, N. Kaifu, Astrophys. J. 318 (1987) L77.
- ²⁶⁴⁰ [274] M. Tanaka, T. Hasegawa, S. S. Hayashi, P. W. J. L. Brand, I. Gatley,
 ²⁶⁴¹ Astrophys. J. 336 (1989) 207.
- ²⁶⁴² [275] M. Tanaka, T. Hasegawa, I. Gatley, Astrophys. J. 374 (1991) 516.
- [276] M. D. Smith, C. J. Davis, A. Lioure, Astron. Astrophys. 327 (1997)
 1206.
- [277] S. K. Ramsay, A. Chrysostomou, T. R. Geballe, P. W. J. L. Brand,
 M. Mountain, Mon. Not. R. Astr. Soc. 263 (1993) 695.
- ²⁶⁴⁷ [278] J. L. Hora, W. B. Latter, Astrophys. J. 461 (1996) 288.
- [279] A. Chrysostomou, P. W. J. L. Brand, M. G. Burton, A. Moorhouse,
 Mon. Not. R. Astr. Soc. 265 (1993) 329.

- [280] D. L. Shupe, J. E. Larkin, R. A. Knop, L. Armus, K. Matthews, B. T.
 Soifer, Astrophys. J. 498 (1987) 267.
- ²⁶⁵² [281] L. E. Kristensen, T. L. Ravkilde, D. Field, J. L. Lemaire, G. P. des
 ²⁶⁵³ Forêts, Astron. Astrophys. 469 (2007) 561.
- ²⁶⁵⁴ [282] J. H. Lacy, R. Knacke, T. R. Geballe, A. T. Tokunaga, Astrophys. J.
 ²⁶⁵⁵ 428 (1994) L69.
- ²⁶⁵⁶ [283] D. A. Neufeld, G. J. Melnick, M. Harwit, Astrophys. J. 506 (1998) L75.
- ²⁶⁵⁷ [284] A. Sternberg, D. A. Neufeld, Astrophys. J. 516 (1999) 371.
- [285] P. S. Parmar, J. H. Lacy, J. M. Achtermann, Astrophys. J. 372 (1991)
 L25.
- [286] K. N. Allers, D. T. Jaffe, J. H. Lacy, B. T. Draine, M. J. Richter,
 Astrophys. J. 630 (2005) 368.
- [287] R. Timmermann, F. Bertoldi1, C. M. Wright, S. Drapatz, B. T. Draine,
 L.Haser, A. Sternberg, Astron. Astrophys. 315 (1996) L281.
- ²⁶⁶⁴ [288] D. Rosenthal, F. Bertoldi, S. Drapatz, Astron. Astrophys. 356 (2000)
 ²⁶⁶⁵ 705.
- [289] B. Lefloch, J. Cernicharo, S. Cabrit, A. Noriega-Crespo, A. Moro-Martín, D. Cesarsky, Astrophys. J. 590 (2003) L41.
- [290] C. M. Wright, S. Drapatz, R. Timmermann, P. van der Werf, R. Kat terloher, T. de Graauw, Astron. Astrophys. 315 (1996) L301.
- ²⁶⁷⁰ [291] D. Cesarsky, P. Cox, G. P. des Forêts, E. F. van Dishoeck, F. Boulanger,
 ²⁶⁷¹ C. M. Wright, Astron. Astrophys. 348 (1999) 945.
- ²⁶⁷² [292] M. Benedettini, T. Giannini, B. Nisini, E. Tommasi, D. Lorenzetti,
 ²⁶⁷³ A. D. Giorgio, P. Saraceno, H. Smith, G. White, Astron. Astrophys.
 ²⁶⁷⁴ 359 (2000) 148.
- [293] B. Nisini, M. Benedettini, T. Giannini, C. Codella, D. Lorenzetti, A. D.
 Giorgio, J. Richer, Astron. Astrophys. 360 (2000) 297.

- ²⁶⁷⁷ [294] S. Molinari, A. Noriega-Crespo, C. Ceccarelli, B. N. T. Giannini,
 D. Lorenzetti, E. Caux, P. Saraceno, G. J. White, Astrophys. J. 538
 ²⁶⁷⁹ (2000) 698.
- [295] E. Habart, F. Boulanger, L. Verstraete, G. P. des Forêts, E. Falgarone,
 A. Abergel, Astron. Astrophys. 397 (2003) 623.
- [296] A. Fuente, J. Martín-Pintado, N. Rodríguez-Fernández, A. Rodríguez-Franco, P. D. Vicente, Astrophys. J. 518 (1999) L45.
- ²⁶⁸⁴ [297] N. J. Rodríguez-Fernández, J. Martín-Pintado, P. D. Vicente,
 ²⁶⁸⁵ A. Fuente, S. Hüttemeister, T. L. Wilson, D. Kunze, Astron. Astro²⁶⁸⁶ phys. 356 (2000) 695.
- [298] D. A. Neufeld, G. J. Melnick, P. Sonnentrucker, E. A. Bergin, J. D.
 Green, K. H. Kim, D. M. Watson, W. J. Forrest, J. L. Pipher, Astrophys. J. 649 (2006) 816.
- [299] S. Maret, E. A. Bergin, D. A. Neufeld, J. D. Green, D. M. Watson,
 M. O. Harwit, L. E. Kristensen, G. J. Melnick, P. Sonnentrucker,
 V. Tolls, M. W. Werner, W. Michael, K. Willacy, Y. Yuan, Astrophys.
 J. 698 (2009) 1244.
- [300] D. Neufeld, B. Nisini, T. Giannini, G. J. Melnick, E. A. Bergin,
 Y. Yuan, S. Maret, V. Tolls, R. Güsten, M. J. Kaufman, Astrophys. J.
 706 (2009) 170.
- [301] T. Giannini, B. Nisini, D. Neufeld, Y. Yuan, S. Antoniucci, A. Gusdorf,
 Astrophys. J. 738 (2011) 80.
- [302] B. Nisini, T. Giannini, D. Neufeld, Y. Yuan, S. Antoniucci, E. A.
 Bergin, G. J. Melnick, Astrophys. J. 724 (2010) 69.
- [303] M. Barsony, G. A. Wolf-Chase, D. R. Ciardi, J. O'Linger, Astrophys.
 J. 720 (2010) 64.
- [304] O. Dionatos, B. Nisini, S. Cabrit, L. Kristensen, G. P. des Forêts,
 Astron. Astrophys. 521 (2010) A7.
- [305] Y. Sheffer, M. G. Wolfire1, D. J. Hollenbach, M. J. Kaufman,
 M. Cordier, Astrophys. J. 741 (2011) 45.

- [306] P. F. Goldsmith, T. Velusamy, D. Li, W. D. Langer, Astrophys. J. 715
 (2010) 1370.
- ²⁷⁰⁹ [307] R. J. Gould, E. E. Salpeter, Astrophys. J. 138 (1963) 393.
- ²⁷¹⁰ [308] D. Hollenbach, E. E. Salpeter, Astrophys. J. 163 (1971) 155.
- ²⁷¹¹ [309] W. Lichten, Phys. Rev. 120 (1960) 848.
- [310] S. Glasstone, K. J. Laidler, H. Eyring, The theory of rate processes: the kinetics of chemical reactions, viscosity, diffusion and electrochemical phenomena, McGraw-Hill, New York, 1941.
- ²⁷¹⁵ [311] A. Winkler, K. D. Rendulic, Int. Rev. Phys. Chem. 11 (1992) 101.
- ²⁷¹⁶ [312] K. Christmann, Prog. Surf. Sci. 48 (1995) 15.
- ²⁷¹⁷ [313] A. Groß, Surf. Sci. Rep. 32 (1998) 291.
- ²⁷¹⁸ [314] G.-J. Kroes, Prog. Surf. Sci. 60 (1999) 1.
- ²⁷¹⁹ [315] W. A. Dino, H. Kasai, A. Okiji, Prog. Surf. Sci. 63 (1998) 63.
- [316] G.-J. Kroes, A. Gross, E.-J. Baerends, M. Scheffler, D. A. Mccormack,
 Acc. Chem. Res. 35 (2002) 193.
- ²⁷²² [317] D. Farías, R. Miranda, Prog. Surf. Sci. 86 (2011) 222.
- [318] D. A. Williams, W. A. Brown, S. D. Price, J. M. C. Rawlings, S. Viti,
 Astron. Geophys. 48 (2007) 1.25.
- ²⁷²⁵ [319] N. Watanabe, A. Kouchi, Prog. Surf. Sci. 83 (2008) 439.
- [320] Y. M. Xiao, S. Buchman, L. Pollack, D. Kleppner, T. J. Greytak, J.
 Chem. Phys. 96 (1992) 4032.
- [321] S. C. Creighan, J. S. A. Perry, S. D. Price, J. Chem. Phys. 124 (2006)
 114701.
- ²⁷³⁰ [322] F. Islam, E. R. Latimer, S. D. Price, J. Chem. Phys. 127 (2007) 064701.
- ²⁷³¹ [323] E. R. Latimer, F. Islam, S. D. Price, Chem. Phys. Lett. 455 (2008) 174.

- [324] A. Yabushita, T. Hama, D. Iida, N. Kawanaka, M. Kawasaki,
 N. Watanabe, M. N. R. Ashfold, H.-P. Loock, Astrophys. J. 682 (2008)
 L69.
- [325] A. Yabushita, T. Hama, D. Iida, N. Kawanaka, M. Kawasaki, Astro phys. J. 129 (2008) 044501.
- [326] L. Gavilan, G. Vidali, J. L. Lemaire, M. Chehrouri, F. Dulieu, J.-H.
 Fillion, E. Congiu, H. Chaabouni, Astrophys. J. 760 (2012) 1.
- ²⁷³⁹ [327] K. Takayanagi, K. Sakimoto, K. Onda, Astrophys. J. 318 (1987) L81.
- ²⁷⁴⁰ [328] B. T. Draine, F. Bertoldi, Astrophys. J. 468 (1996) 269.
- ²⁷⁴¹ [329] T. G. Waech, R. B. Bernstein, J. Chem. Phys. 46 (1967) 4905.
- [330] R. E. Roberts, R. B. Bernstein, C. F. Curtiss, J. Chem. Phys. 50 (1969)
 5163.
- ²⁷⁴⁴ [331] M. Menzinger, Chem. Phys. Lett. 10 (1971) 507.
- ²⁷⁴⁵ [332] R. E. Roberts, J. Chem. Phys. 54 (1971) 1422.
- ²⁷⁴⁶ [333] C. Schwartz, R. J. L. Roy, J. Chem. Phys. 81 (1984) 4149.
- [334] S. Morisset, F. Aguillon, M. Sizun, V. Sidis, J. Chem. Phys. 121 (2004)
 6493.
- [335] S. Morisset, F. Aguillon, M. Sizun, V. Sidis, J. Chem. Phys. 122 (2005)
 194702.
- ²⁷⁵¹ [336] B. Kerkeni, D. C. Clary, Chem. Phys. 338 (2007) 1.
- [337] D. Bachellerie, M. Sizun, F. Aguillon, V. Sidis, J. Phys. Chem. A 113 (2009) 108.
- [338] A. J. Farebrother, A. J. Meijer, D. C. Clary, A. J. Fisher, Chem. Phys.
 Lett. 319 (2000) 303.
- ²⁷⁵⁶ [339] X. Sha, B. Jackson, J. Chem. Phys. 116 (2002) 7158.
- [340] S. Morisset, F. Aguillon, M. Sizun, V. Sidis, J. Phys. Chem. A 105
 (2004) 8571.

- [341] D. Bachellerie, M. Sizun, D. Teillet-Billy, N. Rougeau, V. Sidis, Chem.
 Phys. Lett. 448 (2007) 223.
- [342] M. Bonfanti, S. Casolo, G. F. Tantardini, R. Martinazzo, Phys. Chem.
 Chem. Phys. 13 (2011) 16680.
- ²⁷⁶³ [343] N. Rougeau, D. Teillet-Billy, V. Sidis, Phys. Chem. Chem. Phys. 13 ²⁷⁶⁴ (2011) 17579.
- ²⁷⁶⁵ [344] R. Martinazzo, G. F. Tantardini, J. Chem. Phys. 124 (2006) 124702.
- [345] A. J. H. M. Meijer, A. J. Farebrother, D. C. Clary, A. J. Fisher, J.
 Phys. Chem. A 105 (2001) 2173.
- ²⁷⁶⁸ [346] J. Ree, Y. Kim, H. Shin, Chem. Phys. Lett. 353 (2002) 368.
- [347] A. J. H. M. Meijer, A. J. Farebrother, D. C. Clary, J. Phys. Chem. A 106 (2002) 8996.
- ²⁷⁷¹ [348] A. J. H. M. Meijer, A. J. Fisher, D. C. Clary, J. Phys. Chem. A 107 (2003) 10862.
- ²⁷⁷³ [349] R. Martinazzo, G. F. Tantardini, J. Phys. Chem. A 109 (2005) 9379.
- ²⁷⁷⁴ [350] R. Martinazzo, G. F. Tantardini, J. Chem. Phys. 124 (2006) 124703.
- ²⁷⁷⁵ [351] H. M. Cuppen, L. Hornekar, J. Chem. Phys. 128 (2008) 174707.
- 2776 [352] S. Casolo, R. Martinazzo, M. Bonfanti, G. F. Tantardini, J. Phys.
 2777 Chem. A 113 (2009) 14545.
- [353] M. Sizun, D. Bachellerie, F. Aguillon, V. Sidis, Chem. Phys. Lett. 498
 (2010) 32.
- ²⁷⁸⁰ [354] D. R. Flower, G. D. Watt, Mon. Not. R. Astr. Soc. 209 (1984) 25.
- ²⁷⁸¹ [355] D. R. Flower, G. P. des Forêts, Mon. Not. R. Astr. Soc. 316 (2000) 901.
- ²⁷⁸² [356] S. C. O. Glover, T. Abel, Mon. Not. R. Astr. Soc. 388 (2008) 1627.
- [357] L. Pagani, C. Vastel, E. Hugo, V. Kokoouline, C. H. Greene, A. Bacmann, E. Bayet, C. Ceccarelli, R. Peng, S. Schlemmer, Astron. Astrophys. 494 (2009) 623.